

# Line Parameters and Forward Calculation for Retrieving Carbon Dioxide and Methane (CO<sub>2</sub> & CH<sub>4</sub>) from "GOSAT" Data

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**National Institute for Environmental Studies (NIES), Japan**

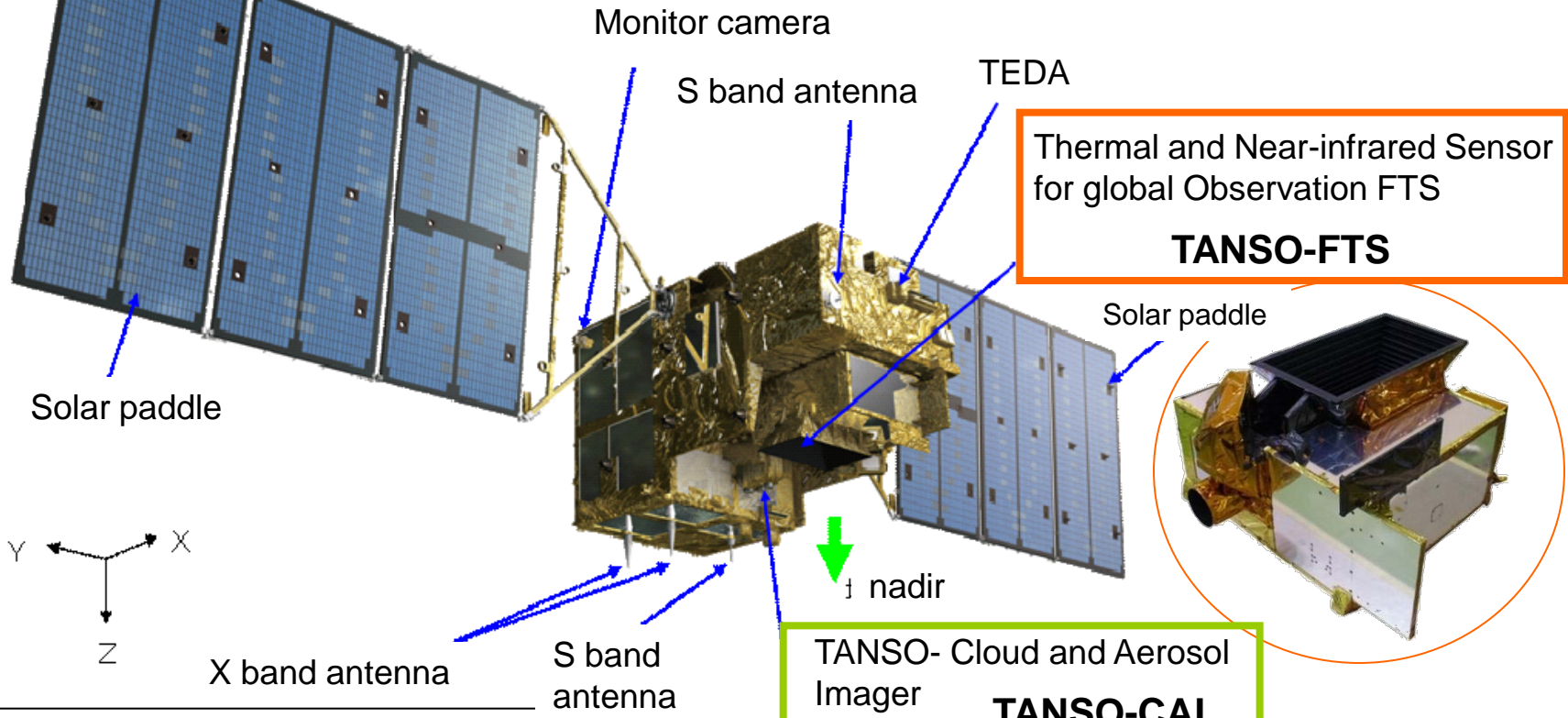
**2) Fujitsu FIP Corporation, @GOSAT DHF, CGER, NIES, Japan**

# Contents

- **GOSAT: Greenhouse gases Observing SATellite**
- **GOSAT Overview**
- **GOSAT First Light Data (spectra) and Present Processed Results**
- **GOSAT Data Product Validation Status**
- **Line Parameters and forward calculations**
  - **Outline of the retrieval method**
- **Future Plan**

GOSAT launch on 23 January 2009 (photo by Mitsubishi)

# "IBUKI"(GOSAT) Sensor System



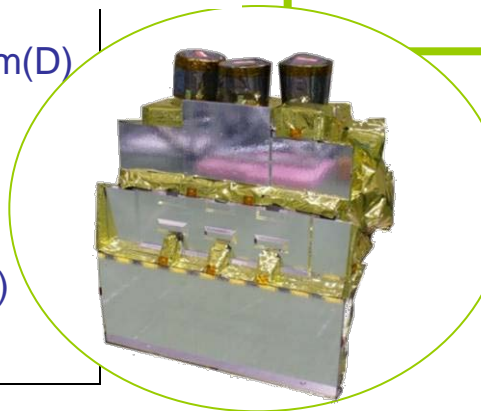
Thermal and Near-infrared Sensor for global Observation FTS

**TANSO-FTS**



TANSO- Cloud and Aerosol Imager

**TANSO-CAI**



Size	
Body	:3.7m(H) × 1.8m(W) × 2.0m(D)
Paddle span	: 13.7 m
Mass	: 1750 kg
Power	: 3.8 kw(EOL)
Life Time	: 5 years
Launch	: 23 Jan 2009 12:54 (JST)
Launch Vehicle	: H-IIA Rocket F-15

O r b i t	<b>Sun Synchronous</b>	
	<b>Local time</b>	<b>13:00+/-0:15</b>
	<b>Altitude</b>	<b>666 km</b>
	<b>Inclination</b>	<b>98 deg.</b>
	<b>Recurrent</b>	<b>3 days</b>



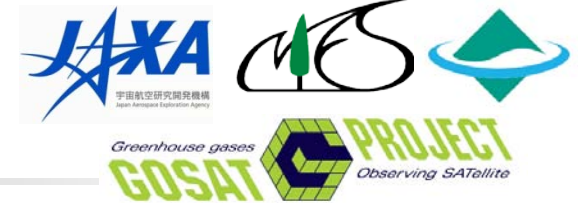
# TANSO-FTS



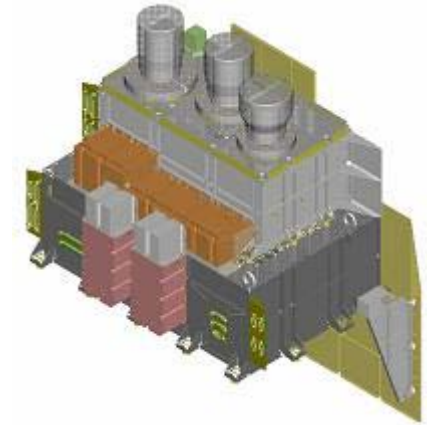
Ground Pointing Mechanism and Fore optics	Configuratio n	2-axes scanner (fully redundant) for ground pointing and calibration			
	Scanning	Cross Track (+/-35 deg) Along Track (+/-20 deg)			
	Field of view	IFOV <10.5 km 790 km (scan width) (latitude of 30 deg)			
Fourier Transform Spectrometer	Speed	0.25, 0.5, 1 (Interferogram)/sec			
	Spectral band	1P, 1S	2P, 2S	3P, 3S	4
	Coverage (micron)	0.75-0.78	1.56-1.72	1.92-2.08	5.5-14.3
	resolution (cm <sup>-1</sup> )	0.5	0.2	0.2	0.2
	Detector	Si	InGaAs	InGaAs	PC-MCT
	Calibration	Solar Irradiance, Deep Space, Moon, Diode Laser (1.55 micron, ILS)			Blackbody, Deep space



# TANSO-CAI



- CAI is operated together with FTS to
- detect aerosol spatial distribution and cloud coverage
  - retrieve scattering spectral characteristics of aerosol

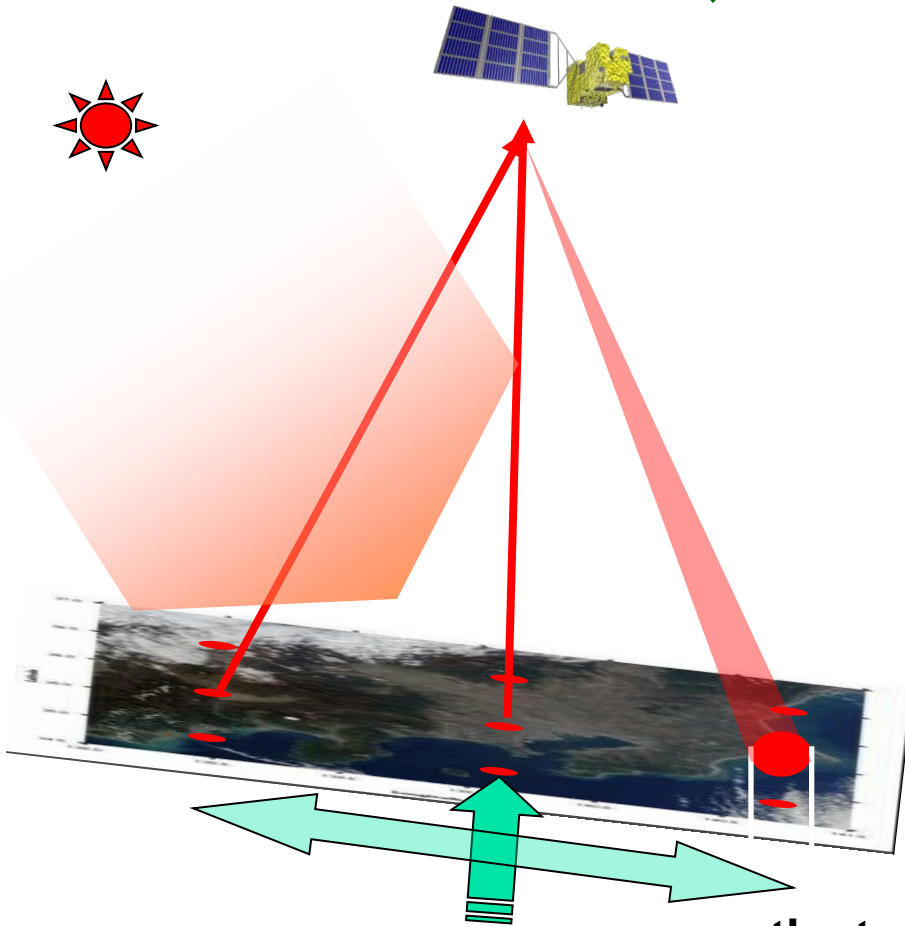


Band No.	Observation Band (nm)	Center Wavelength (nm)	Spatial Resolution (IFOV) (km)	FOV (km)	No. of Pixels (cross track)
1	372-387	380	0.5	1000	2000
2	667-680	678	0.5	1000	2000
3	866-877	870	0.5	1000	2000
4	1560-1640	1620	1.5	750	500



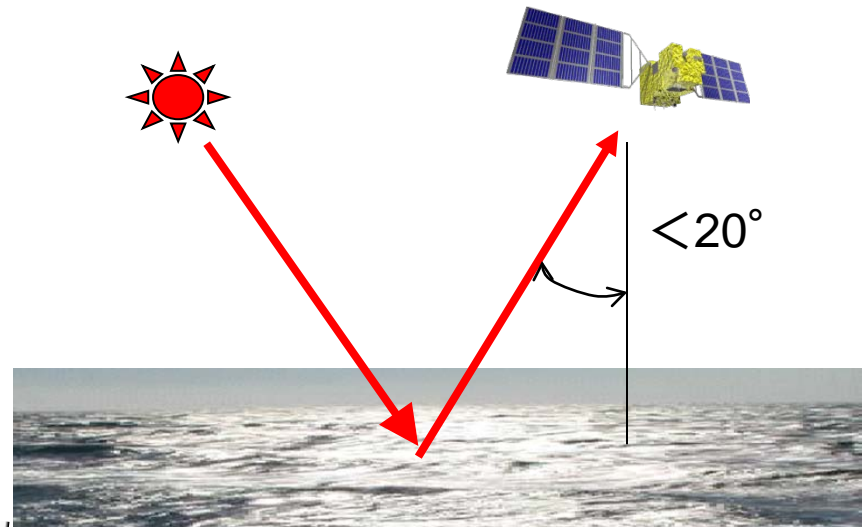


# Observational Mode (TANSO-FTS)



Pointing along the track

cross the track

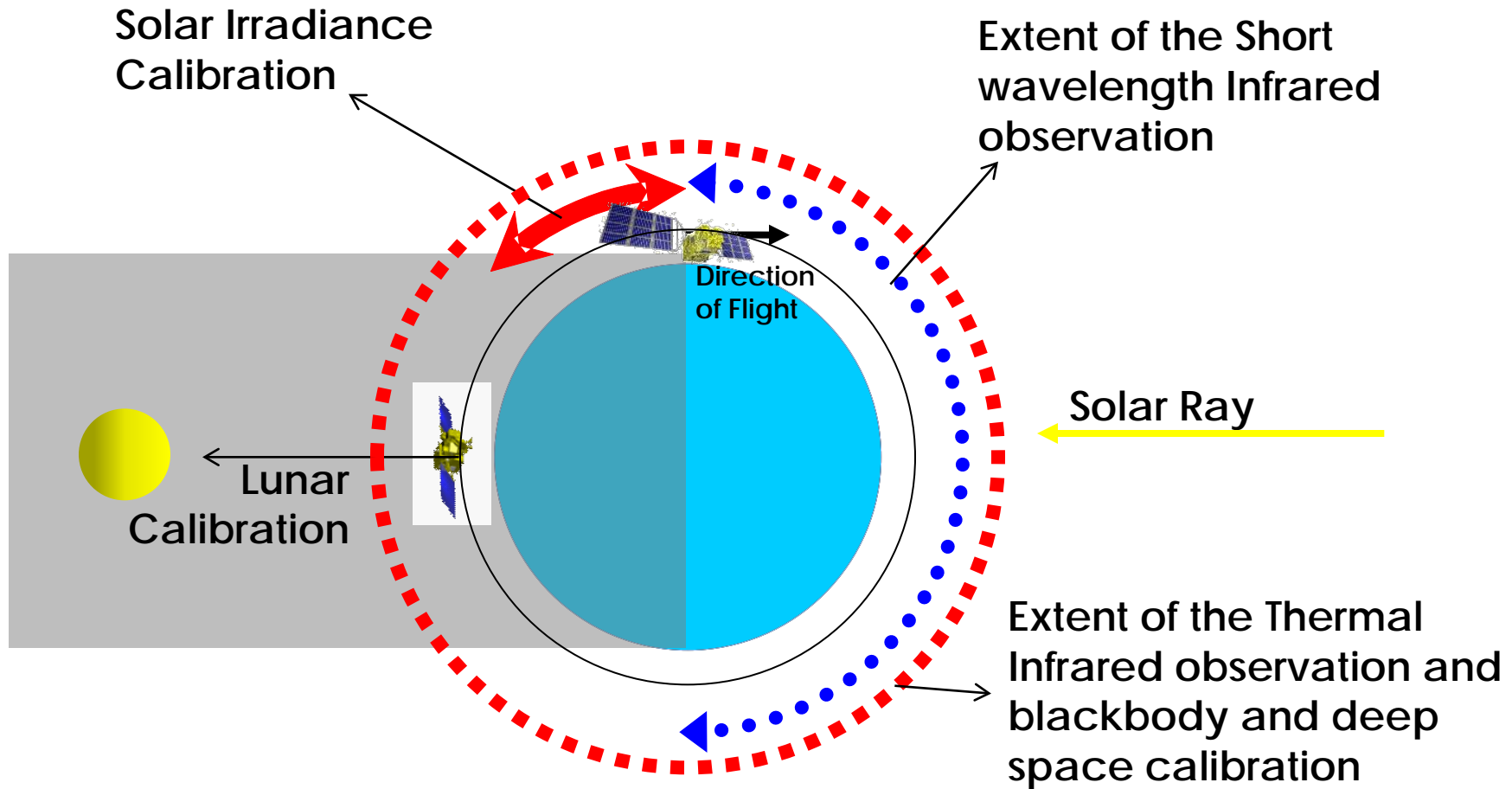


**Sun Glint Pointing Mode**



# TANSO-FTS

## In-Orbit Operation



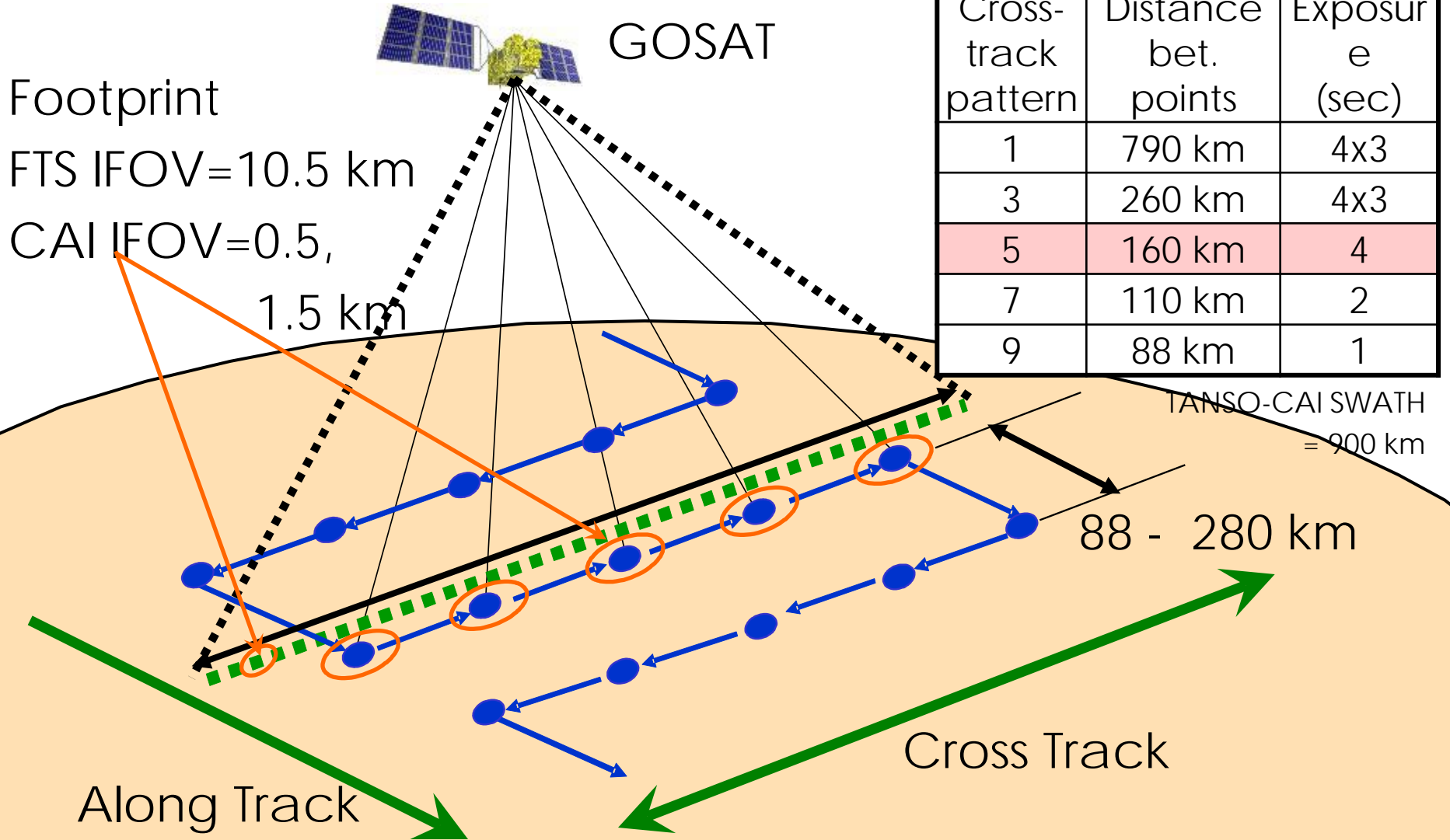
Blackbody and deep space calibration are performed at regular intervals.



# TANSO-FTS



## Observation Pattern

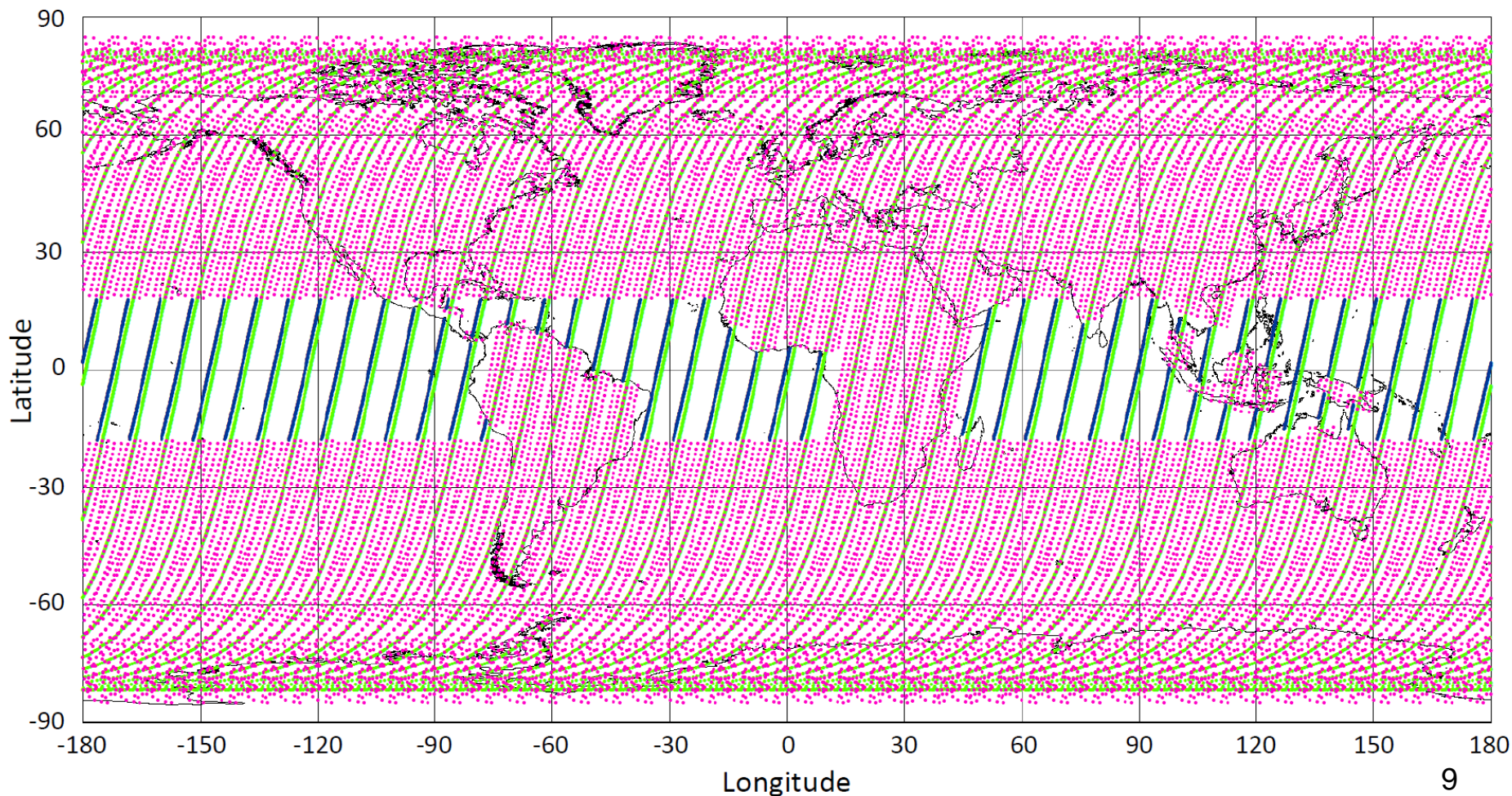


Cross-track pattern	Distance bet. points	Exposure (sec)
1	790 km	4x3
3	260 km	4x3
5	160 km	4
7	110 km	2
9	88 km	1



# 5 point mode + sunglint measurement over the sea (equinox)

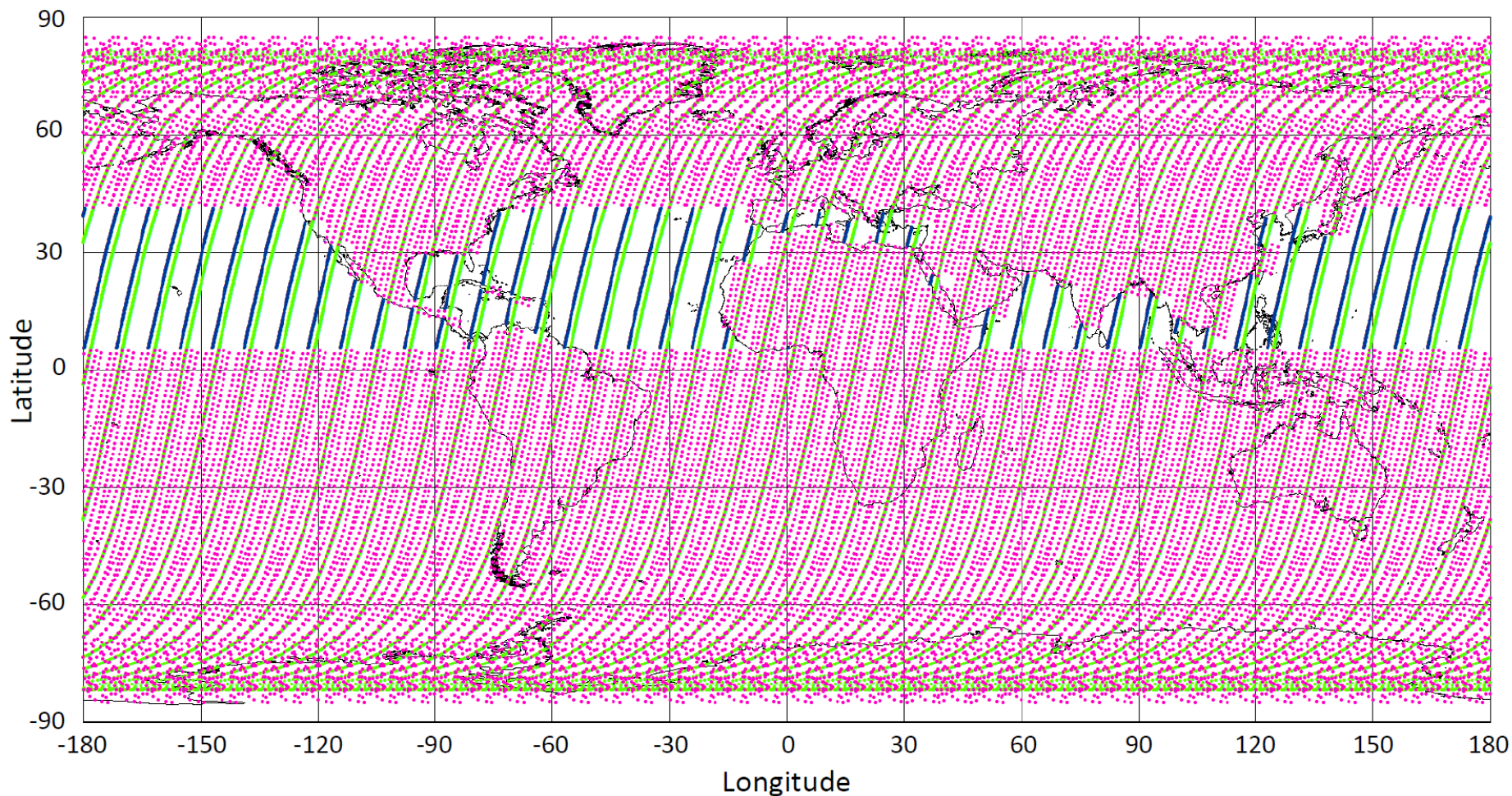
5-point and Sunglint mode : Vernal Equinox

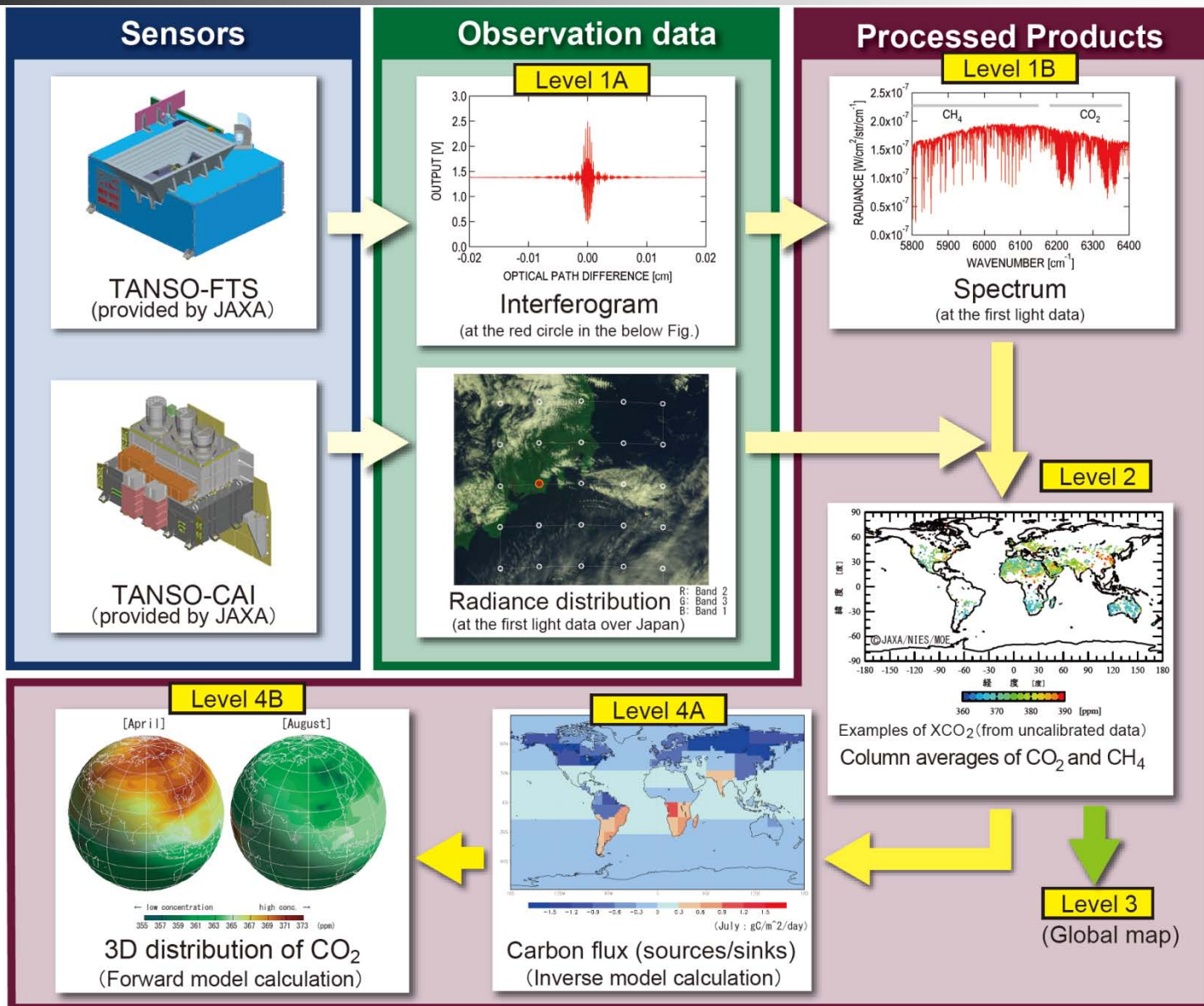




# 5 point mode + sunglint measurement over the sea (summer solstice)

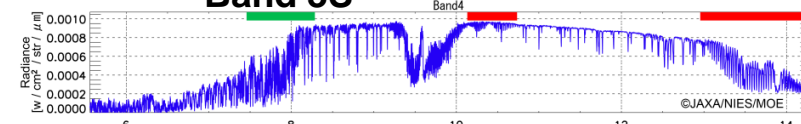
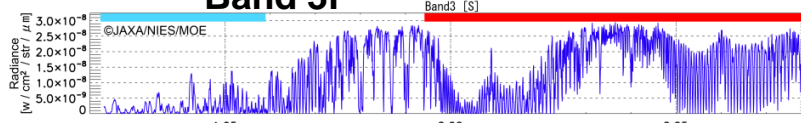
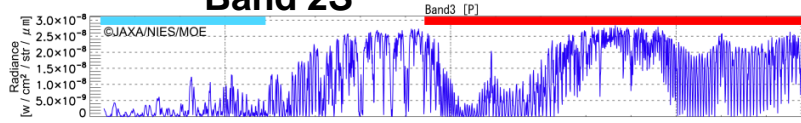
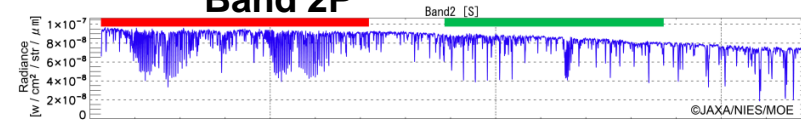
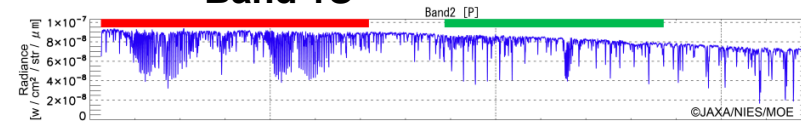
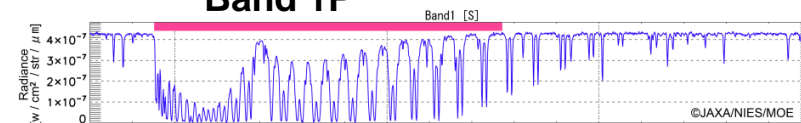
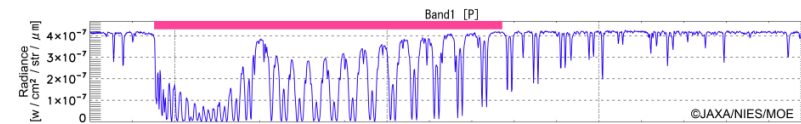
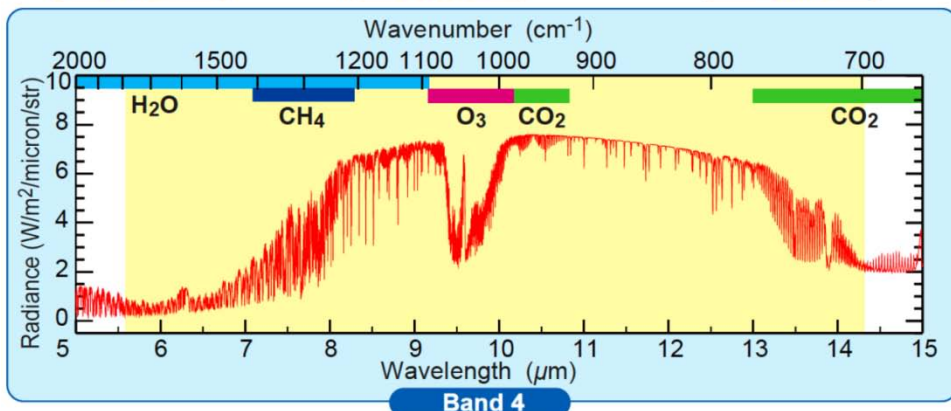
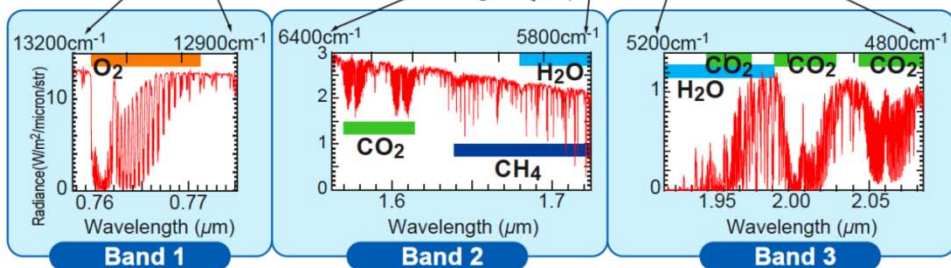
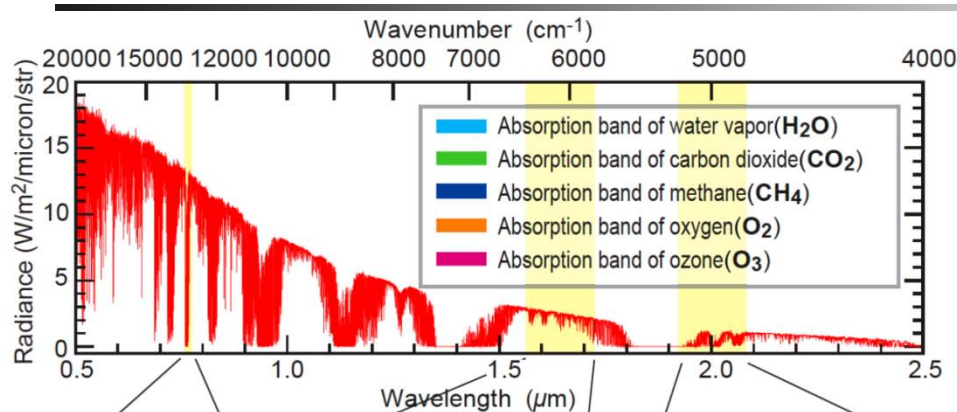
5-point and Sunglint mode : Summer Solstice







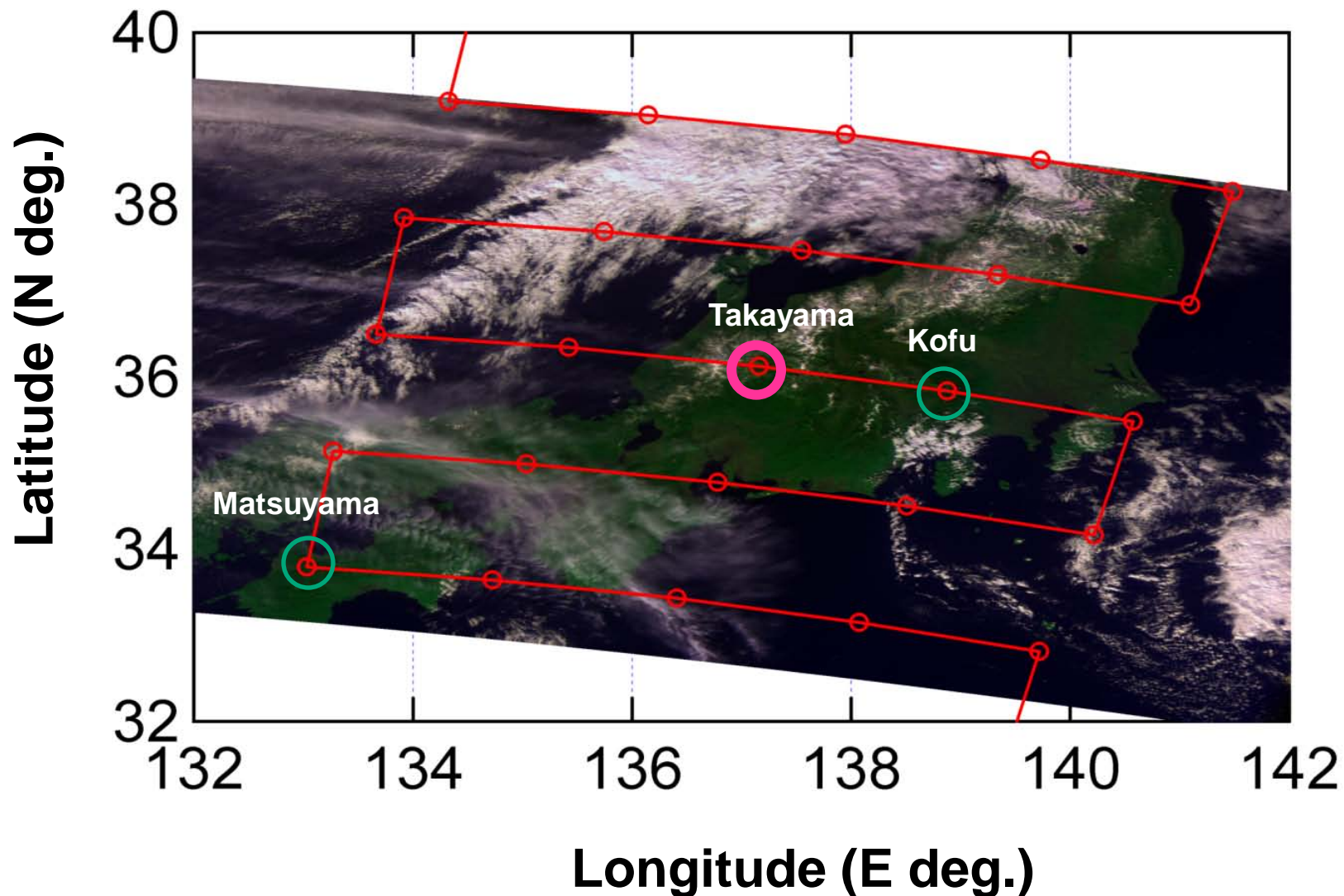
# TANSO-FTS Level 1B



carbon dioxide methane oxygen water vapor

# GOSAT First light spectra over Japan

February 7, 2009 CAI L1B equivalent Image

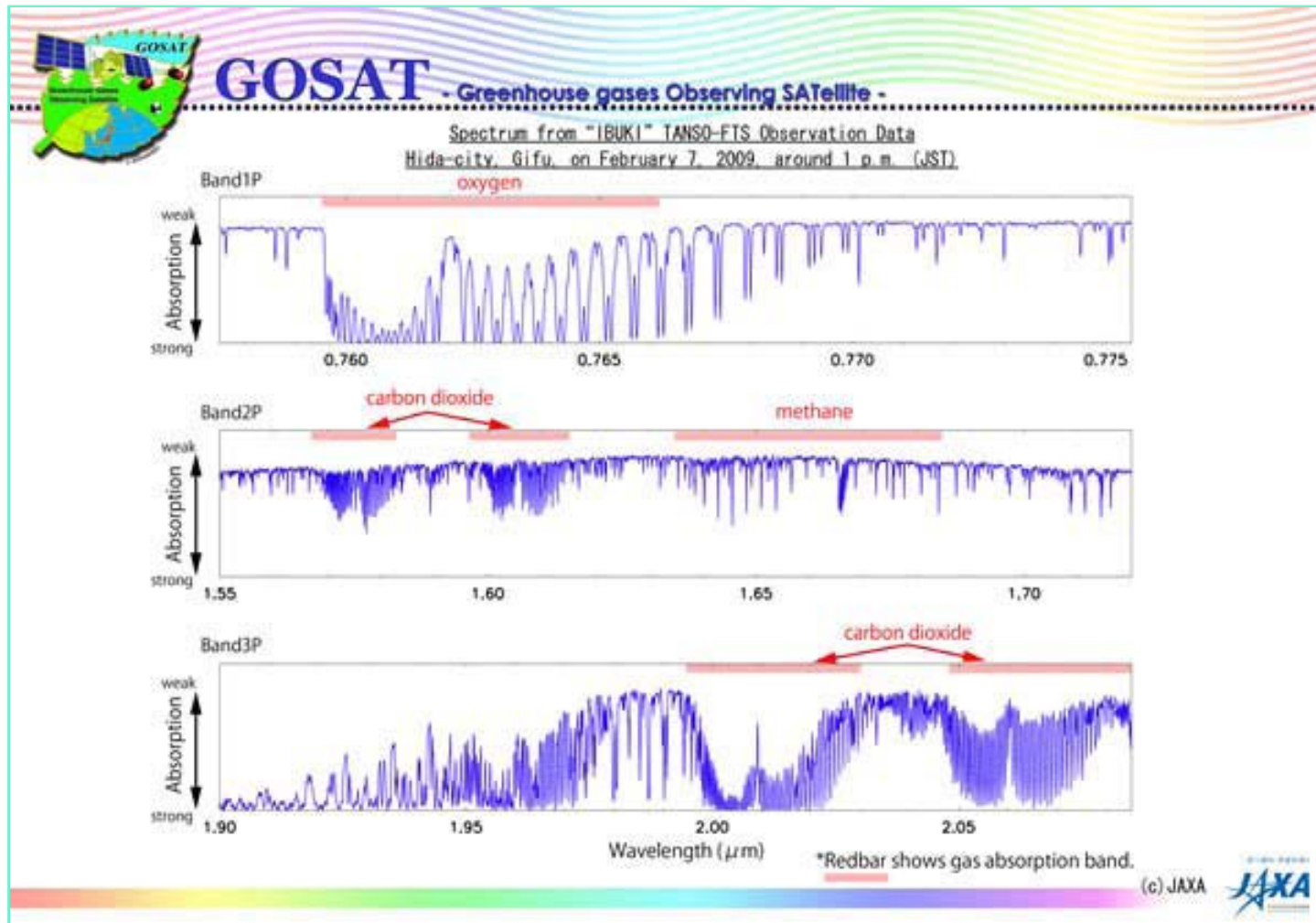


# GOSAT first Light Data

(February 7, 2009)

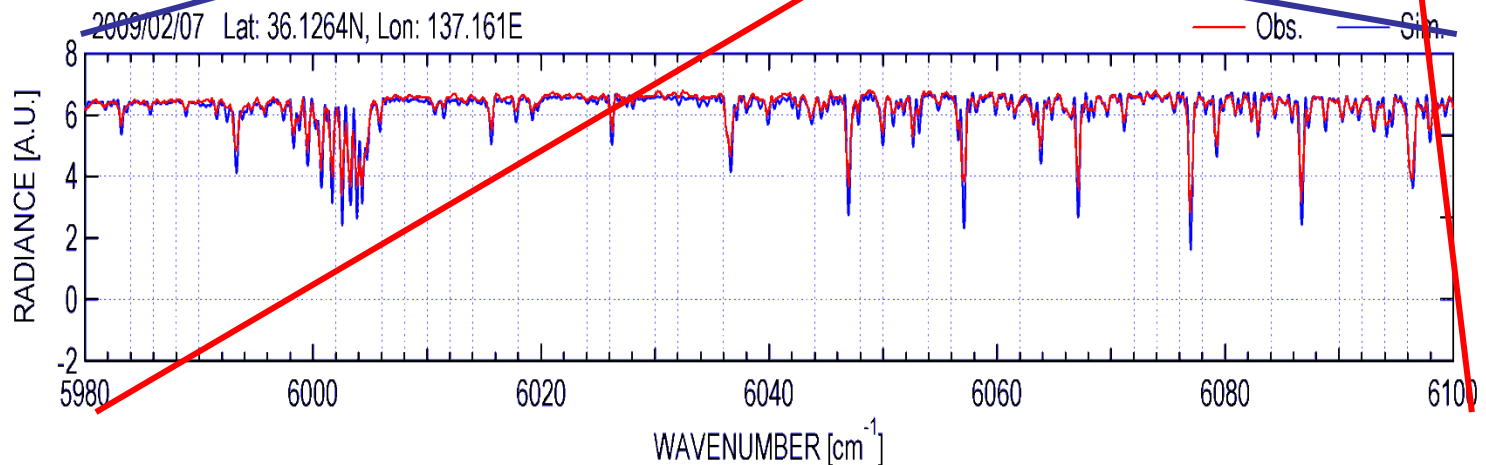
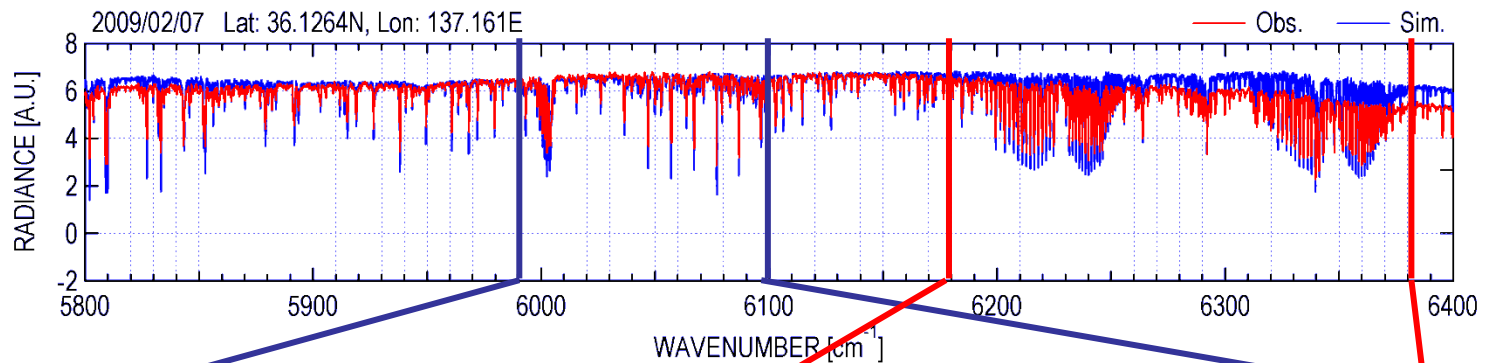


- The GOSAT first light data was press released by JAXA on February 9, 2009.



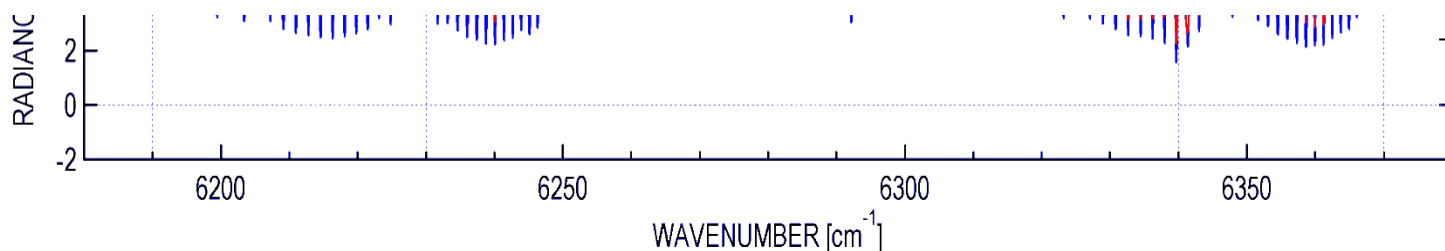


- NIES contributed to show the simulation spectra.

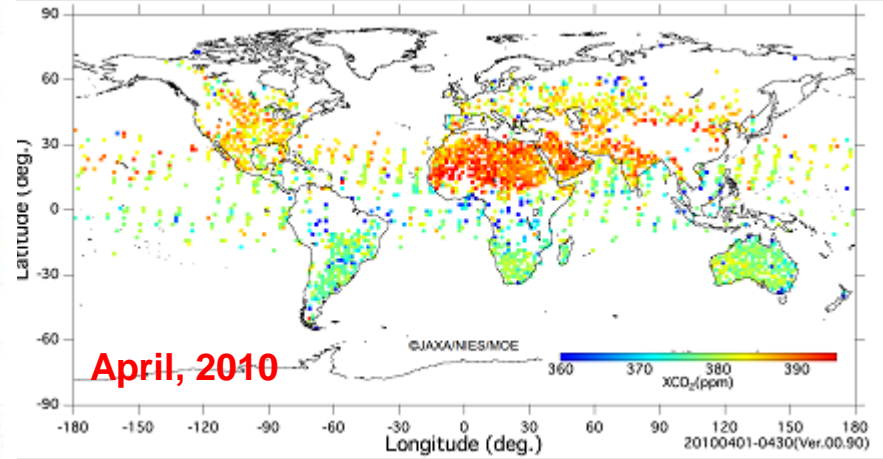
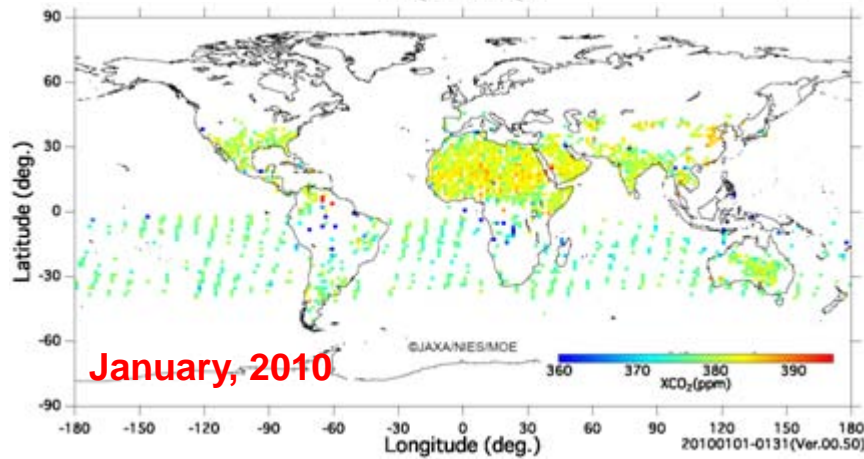
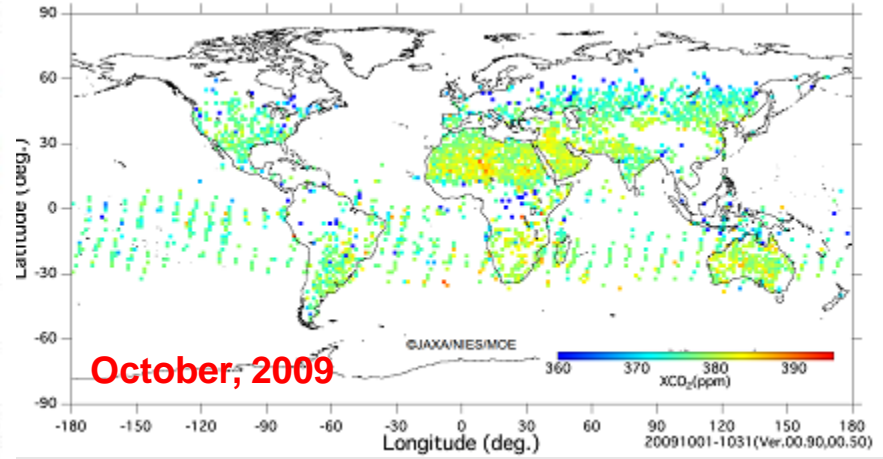
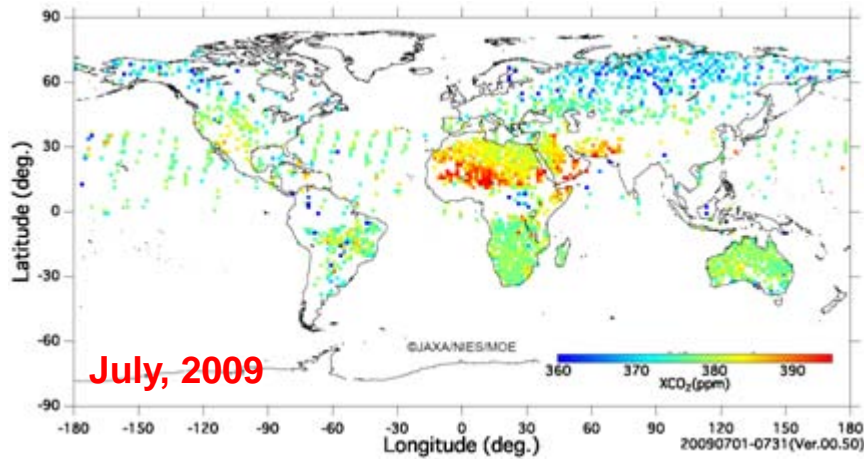


CH<sub>4</sub>

CO<sub>2</sub>

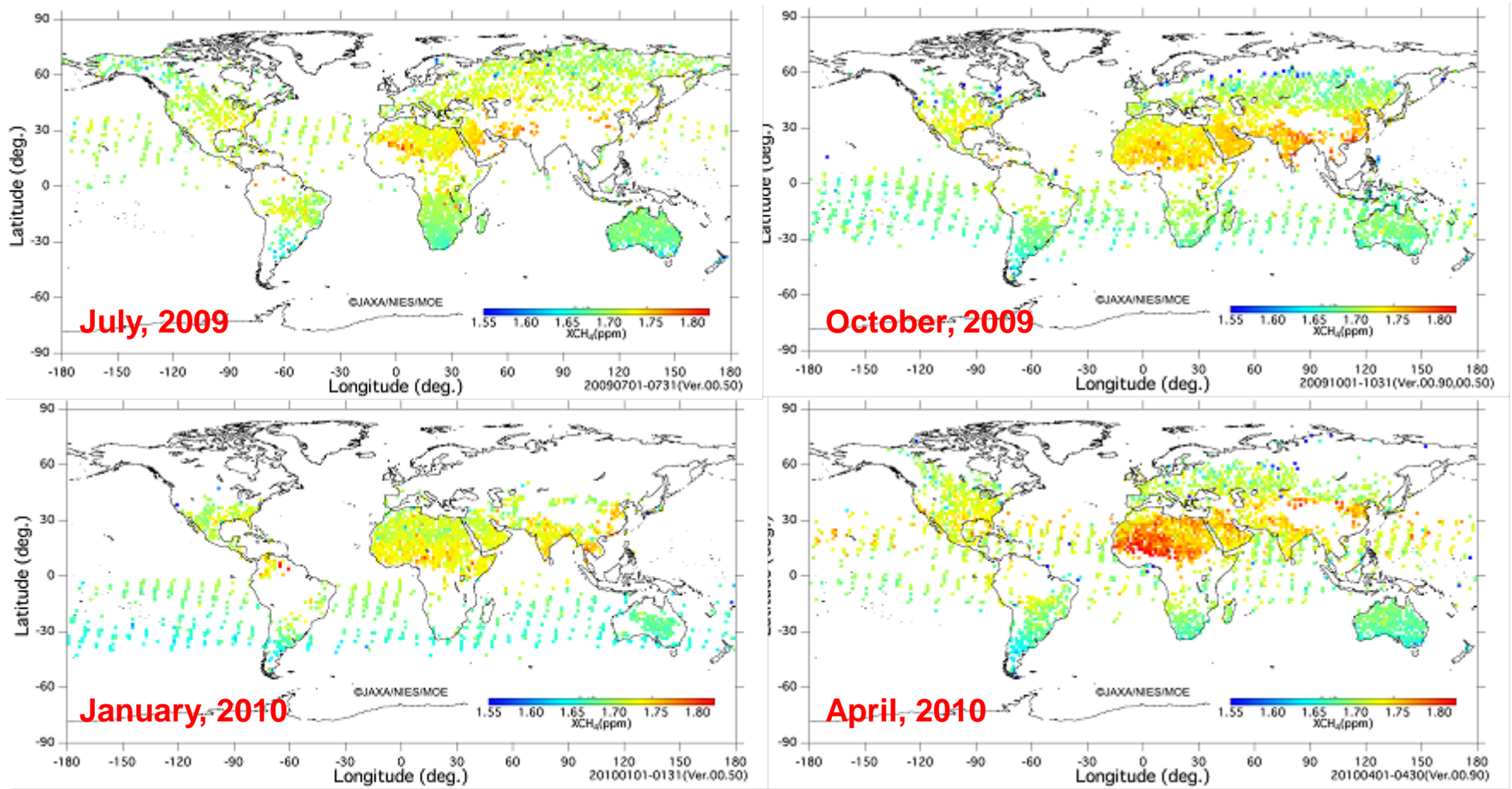


# Seasonal variation of XCO<sub>2</sub>



Areas of elevated concentrations are found in the desert areas of the Arabian Peninsula and northern Africa as well as their surroundings. The extent of the elevated concentration in these areas appears to depend on the season. It is probable that these elevated concentrations were largely influenced by blown dusts of the deserts.

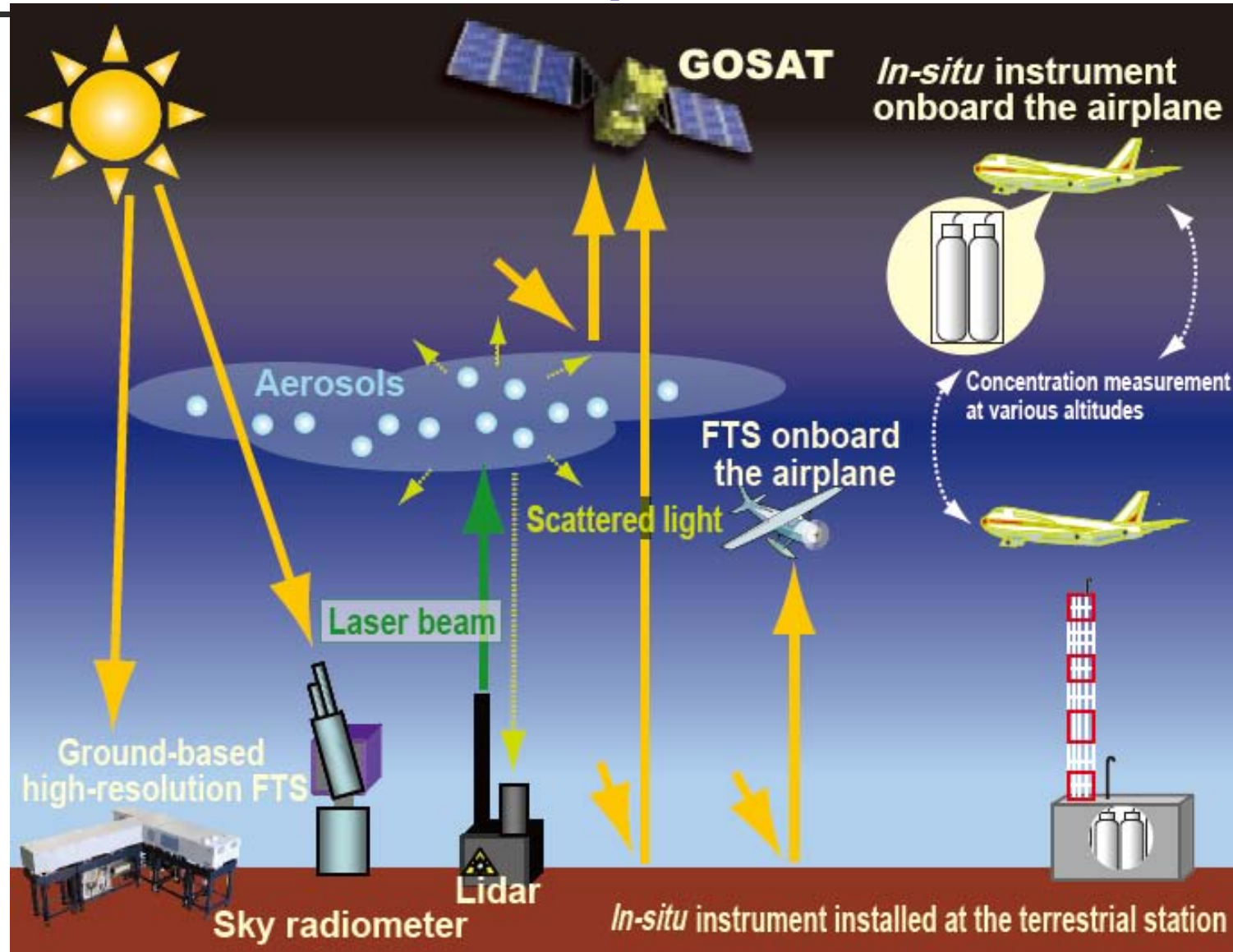
# Seasonal variation of XCH<sub>4</sub>



Areas of elevated concentrations are found in the desert areas of the Arabian Peninsula and northern Africa as well as their surroundings. The extent of the elevated concentration in these areas appears to depend on the season. It is probable that these elevated concentrations were largely influenced by blown dusts of the deserts.



# Schematic illustration of validation experiments

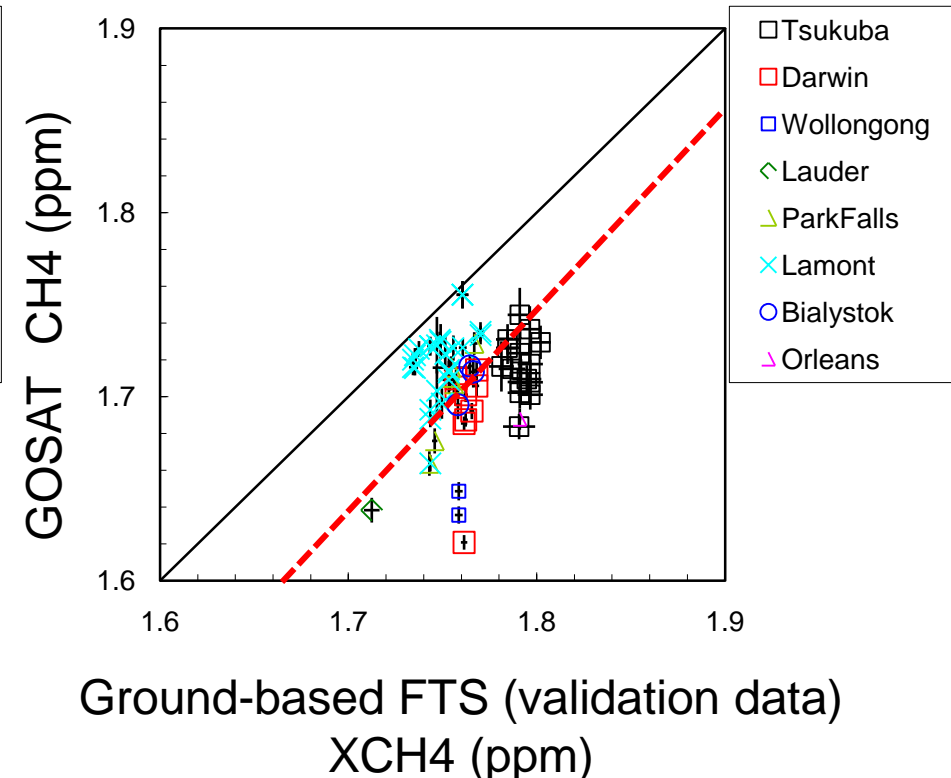
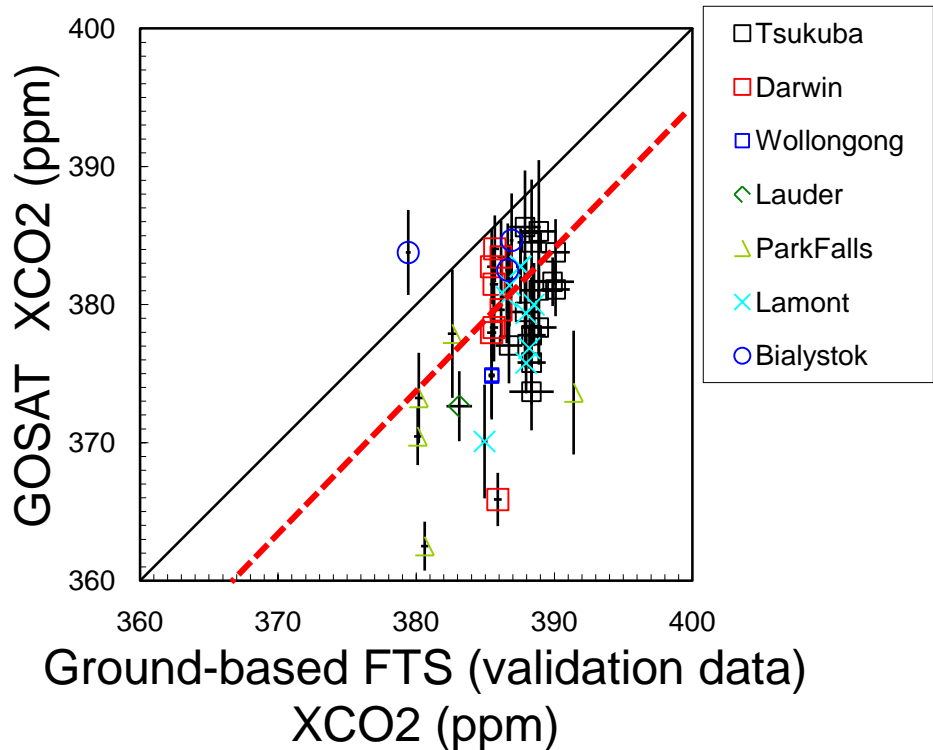




# Comparison of GOSAT data with ground-based FTS data

XCO<sub>2</sub>

XCH<sub>4</sub>



**GOSAT product: V00.50 & V00.80**

Validation data: TCCON and other FTSs data.

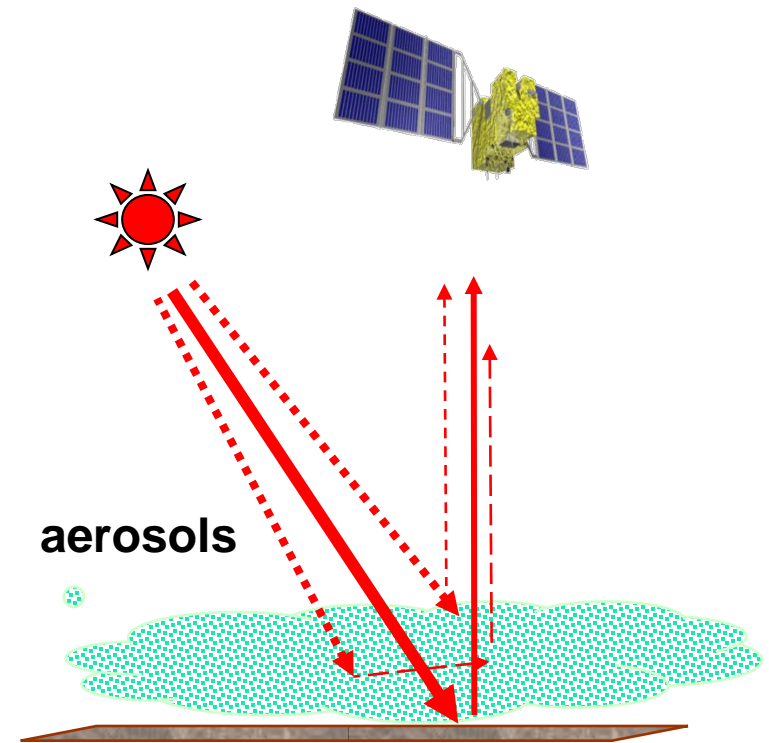
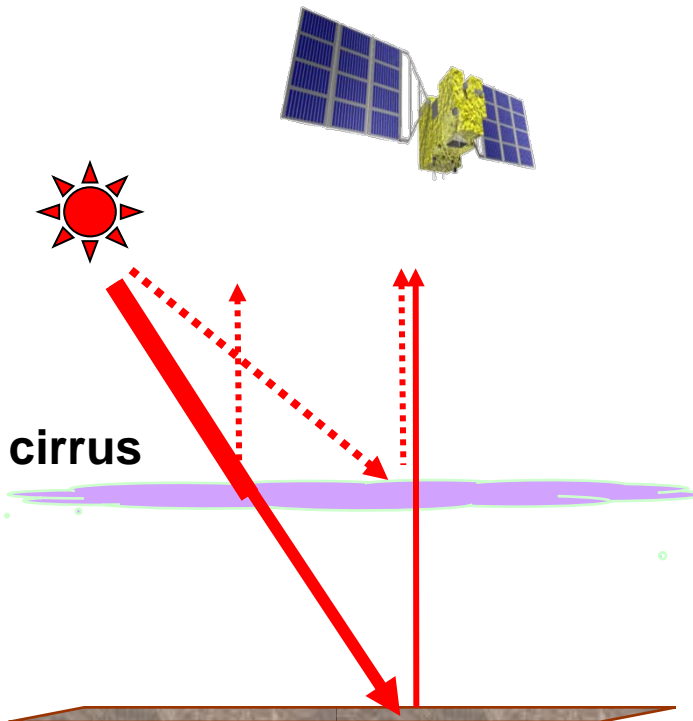


# GOSAT Spectral Calculation (Forward Model for V00.xx)



- Assumption: Plane parallel atmosphere, Lambert surface over the land
- Discrete ordinate method, and a modified Duran's (2005) high-speed computation approach.
- HITRAN 2008 is used as a molecular spectroscopic database, except for H<sub>2</sub>O and CH<sub>4</sub>. Lyulin et al. (2009) is used for CH<sub>4</sub>. HITRAN 2006 is used for water vapor.
- Gas absorption coefficient: cross section table is computed by LBLRTM (v11.3)
- Solar irradiance spectra: provided by R. L. Kurucz
- Rayleigh scattering, cirrus cloud, and aerosol scattering are computed.

# Reflection from cirrus/aerosol

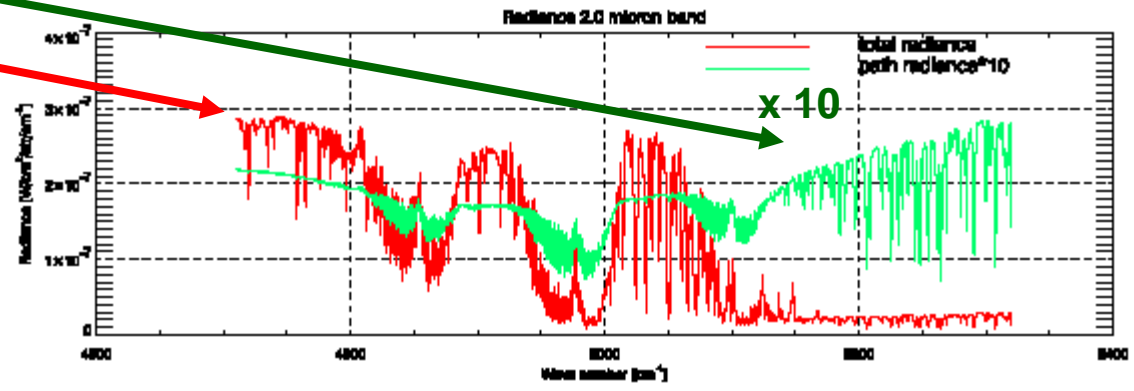
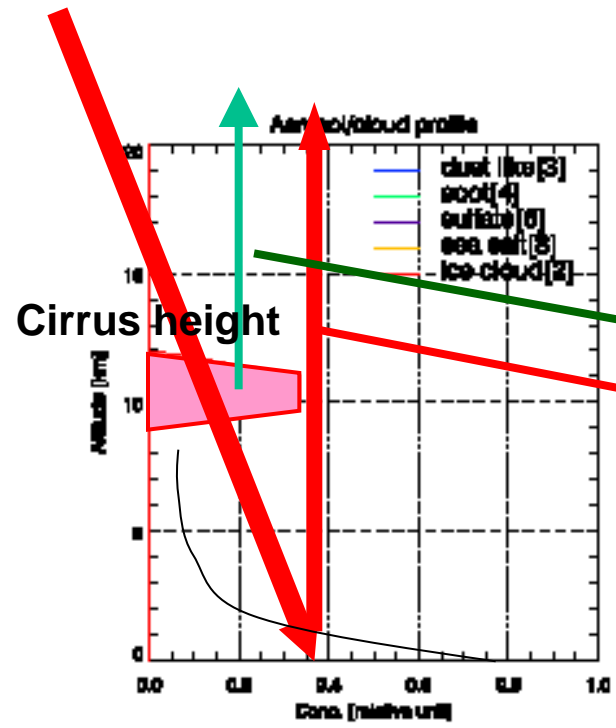
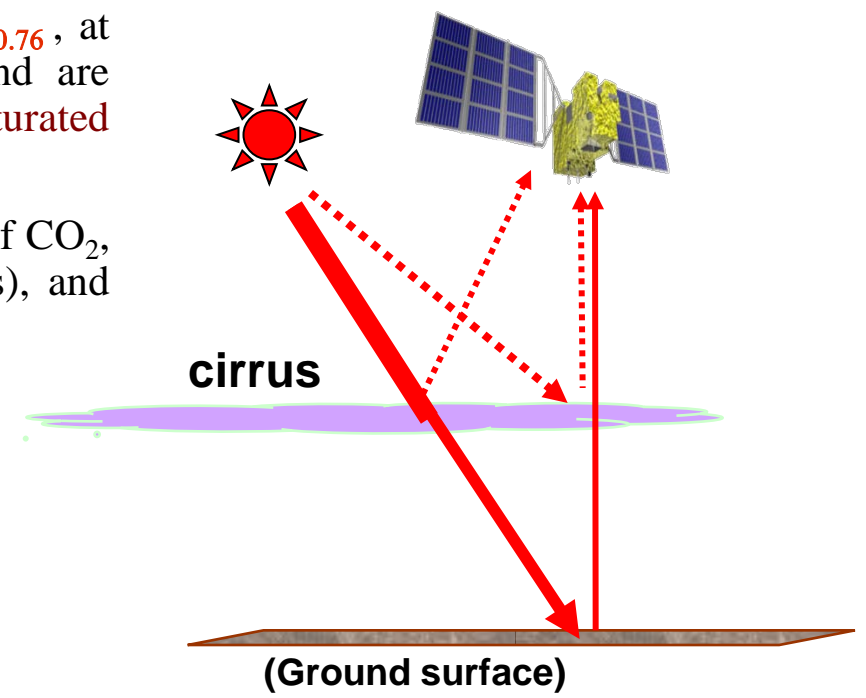


**Cirrus effects will be cancelled with a two-step estimation method by using the H<sub>2</sub>O saturated spectral region of the 2.0 μm band**

**Some types of the aerosols (e.g. Including black carbon) will affect serious error on the retrieval. We are now investigating a way to overcome the aerosol effects.**

# Retrieval concept for cirrus screening (Two step method)

- **Step #1:** The cirrus optical depth ( $\tau$ ), the cirrus cloud center height ( $h$ ), and ground surface albedo ( $\alpha_{0.76}$ , at 50  $\text{cm}^{-1}$  intervals) at the 0.76  $\mu\text{m}$  ( $\text{O}_2\text{-A}$ ) band are estimated from the 0.76  $\mu\text{m}$  band and the  $\text{H}_2\text{O}$  saturated spectral region of the 2.0  $\mu\text{m}$  band
- **Step #2:** Simultaneous retrieval of column density of  $\text{CO}_2$ , ground surface albedo ( $\alpha_{1.6}$  at 25  $\text{cm}^{-1}$  intervals), and re-estimation of  $\tau$  from the 1.6  $\mu\text{m}$  band



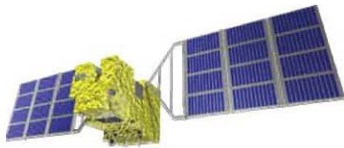
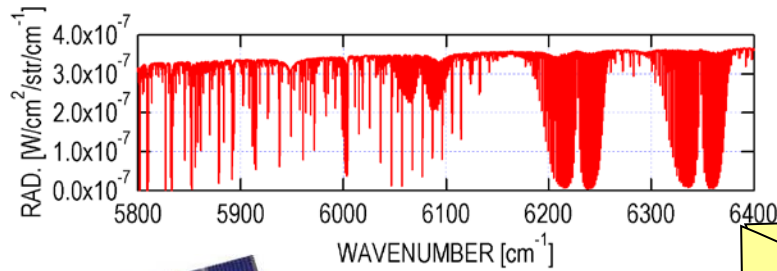
# TANSO-FTS SWIR Level 2 Data

## Processing (up to Versions 00.90)

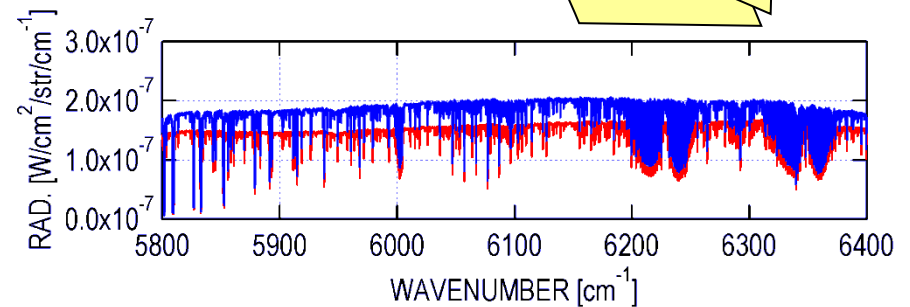
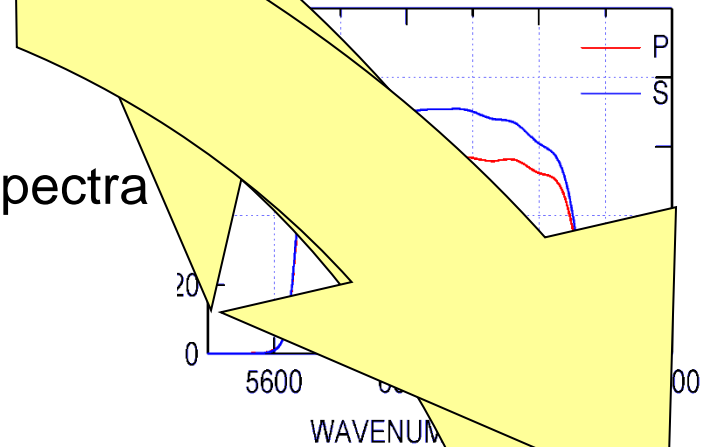
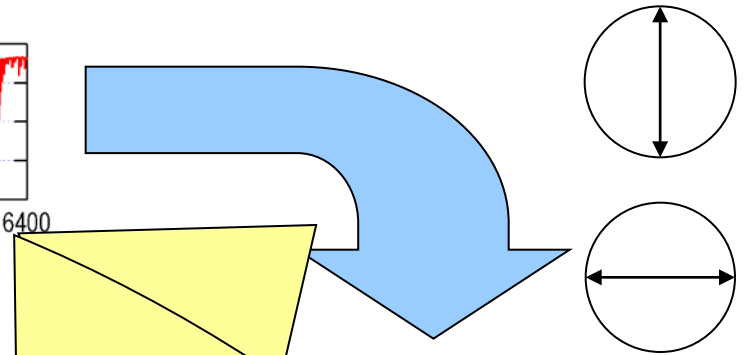
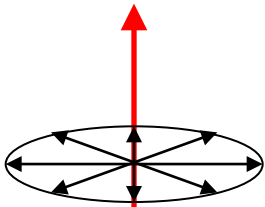
Three steps for data processing

- 1) Data Screening (filtering)
- 2) Optimal Estimation (retrieval)
- 3) Quality Check (screening)

Incident light is divided by two polarized component by PBS.



Version 2  
(Vector RT code)  
Simulate measured spectra  
directly.



**MEASURED SPECTRA 25**

# Radiative transfer model: Hstar and Pstar2.

## ■ High spectral resolution version of the System for Transfer of Atmospheric Radiation (Hstar); [Nakajima et al.]

- Scalar radiative transfer (radiance calculation only).
- **Hybrid model: Discrete-ordinate / matrix-operator (adding) method** (Nakajima and Tanaka, 1986).
- Multilayer multiple scattering media.
- The delta-M truncation method.
- Exact single scattering correction; TMS- and IMS-method (Nakajima and Tanaka, 1988).
- Rough ocean surface model (Nakajima and Tanaka, 1983) / Lambert surface.
- Direct multi-solar beam (for LUT calculation).
- Arbitrary viewing geometry (Source function integration).

## ■ Vector radiative transfer model (Pstar2).

- Same as the above radiative transfer scheme.
- Coupled atmosphere-ocean system including polarization effect.



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*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 99, Issue 1-3, May 2006, Pages 283-294*  
 Hoarty, D.J.; Bentley, C.D.; Crowley, B.J.B.; Davidson, S.J.; Gales, S.G.; Graham, P.; Harris, J.W.O.; Iglesias, C.A.; James, S.F.; Smith, C.C.  
[Cited by Scopus \(12\)](#)
4. **Radiation characteristics of Botryococcus braunii, Chlorococcum littorale, and Chlorella sp. used for CO<sup>2</sup> fixation and biofuel production**  
*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 110, Issue 17, November 2009, Pages 1879-1893*  
 Berberoglu, H.; Gomez, P.S.; Pilon, L.
5. **Light scattering theories and computer codes**  
*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 110, Issue 11, July 2009, Pages 833-843*  
 Wriedt, T.  
[Cited by Scopus \(3\)](#)
6. **Spectra of dense pure hydrogen plasma in Balmer area**  
*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 83, Issue 3-4, February 2004, Pages 387-405*  
 Vitel, Y.; Gavrilova, T.V.; D'yachkov, L.G.; Kurilenkov, Y.K.  
[Cited by Scopus \(2\)](#)
7. **The HITRAN 2004 molecular spectroscopic database**  
*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 96, Issue 2, December 2005, Pages 139-204*  
 Rothman, L.S.; Jacquemart, D.; Barbe, A.; Chris Benner, D.; Birk, M.; Brown, L.R.; Carleer, M.R.; Chackerian, C.; Chance, K.; Coudert, L.H.; Dana, V.; Devi, V.M.; Flaud, J.M.; Gamache, R.R.; Goldman, A.; Hartmann, J.M.; Jucks, K.W.; Maki, A.G.; Mandin, J.  
[Cited by Scopus \(1141\)](#)
25. **Matrix formulations of radiative transfer including the polarization effect in a coupled atmosphereocean system**  
*Journal of Quantitative Spectroscopy and Radiative Transfer, Volume 111, Issue 6, April 2010, Pages 878-894*  
 Ota, Y.; Higurashi, A.; Nakajima, T.; Yokota, T.  
[Cited by Scopus \(1\)](#)

Measurement

Prior constraint

Cost Function:  $J(\mathbf{x}) = \underbrace{[\mathbf{y} - \mathbf{F}(\mathbf{x})]^T \mathbf{S}_\varepsilon^{-1} [\mathbf{y} - \mathbf{F}(\mathbf{x})]}_{\text{Measurement}} + \underbrace{(\mathbf{x} - \mathbf{x}_a)^T \mathbf{S}_a^{-1} (\mathbf{x} - \mathbf{x}_a)}_{\text{Prior constraint}}$

$$\mathbf{x}_{i+1} = \mathbf{x}_i + \left( \mathbf{K}_i^T \mathbf{S}_\varepsilon^{-1} \mathbf{K}_i + \mathbf{S}_a^{-1} + \lambda_i \mathbf{D}^2 \right)^{-1} \left\{ \mathbf{K}_i^T \mathbf{S}_\varepsilon^{-1} [\mathbf{y} - \mathbf{F}(\mathbf{x}_i)] + \mathbf{S}_a^{-1} (\mathbf{x}_i - \mathbf{x}_a) \right\}$$

- $\mathbf{x}_i$  State vector
- $\mathbf{x}_a$  A priori values for state vector
- $\mathbf{y}$  Observed spectrum
- $\mathbf{S}_a$  Covariance matrix for a priori values for state vector
- $\mathbf{S}_\varepsilon$  Covariance matrix for observed spectrum
- $\mathbf{F}(\mathbf{x}_i)$  Forward model
- $\mathbf{K} = \frac{\partial \mathbf{F}(\mathbf{x})}{\partial \mathbf{x}}$  Jacobian

# Optimal Estimation



CO<sub>2</sub> and CH<sub>4</sub> are retrieved under the assumption of cloud-free condition.

*CO<sub>2</sub> retrieval: 6180 ~ 6380 cm<sup>-1</sup> (TANSO-FTS Band 2)*

~~CO<sub>2</sub> profile~~

**HITRAN 2008**

aerosol optical thickness

surface albedo (land) / surface wind speed (ocean)

adjustment factor for relative radiometric calibration

solar Fraunhofer lines, H<sub>2</sub>O absorption lines are masked

*CH<sub>4</sub> retrieval: 5900 ~ 6150 cm<sup>-1</sup> (TANSO-FTS Band 2)*

~~CH<sub>4</sub> profile~~

**Lyulin et al. (2009)**

H<sub>2</sub>O profile (HITRAN 2006)

**Dry air column is  
calculated from the  
meteorological data.**

aerosol optical thickness


surface albedo (land) / surface wind speed (ocean)

adjustment factor for relative radiometric calibration

**→ converted to  
XCO<sub>2</sub> and XCH<sub>4</sub>**

solar Fraunhofer lines are masked.

## Next Improvement of Forward Calculation

- Changing Line Parameter Database
    - CH<sub>4</sub>: Lyulin et al. (2009) → HITRAN 2008
    - H<sub>2</sub>O: HITRAN 2006 → HITRAN 2008
  - Line mixing considerations for all bands
- 
- Retrieval Improvements
    - Simultaneous retrieval from Band 1 (O<sub>2</sub>-A band) and Band 2 (CO<sub>2</sub> 1.6 micron)
    - Parameters: CO<sub>2</sub> profile, CH<sub>4</sub> profile, H<sub>2</sub>O profile, aerosol optical thickness (AOT), surface albedo/wind speed, temperature shift, surface pressure, wavenumber correction coefficient

## Incorporation of LM + CIA for the O<sub>2</sub> A-band

- Three types of the forward calculation were tested

1) LM+ CIA :  $P_s = 989$  hPa (a priori: 981 hPa),  
 $\chi^2 = 1.5038$

Residuals (obs. – calc.) of spectra are sufficiently small.

2) Speed dependence are also included → show very small effect :

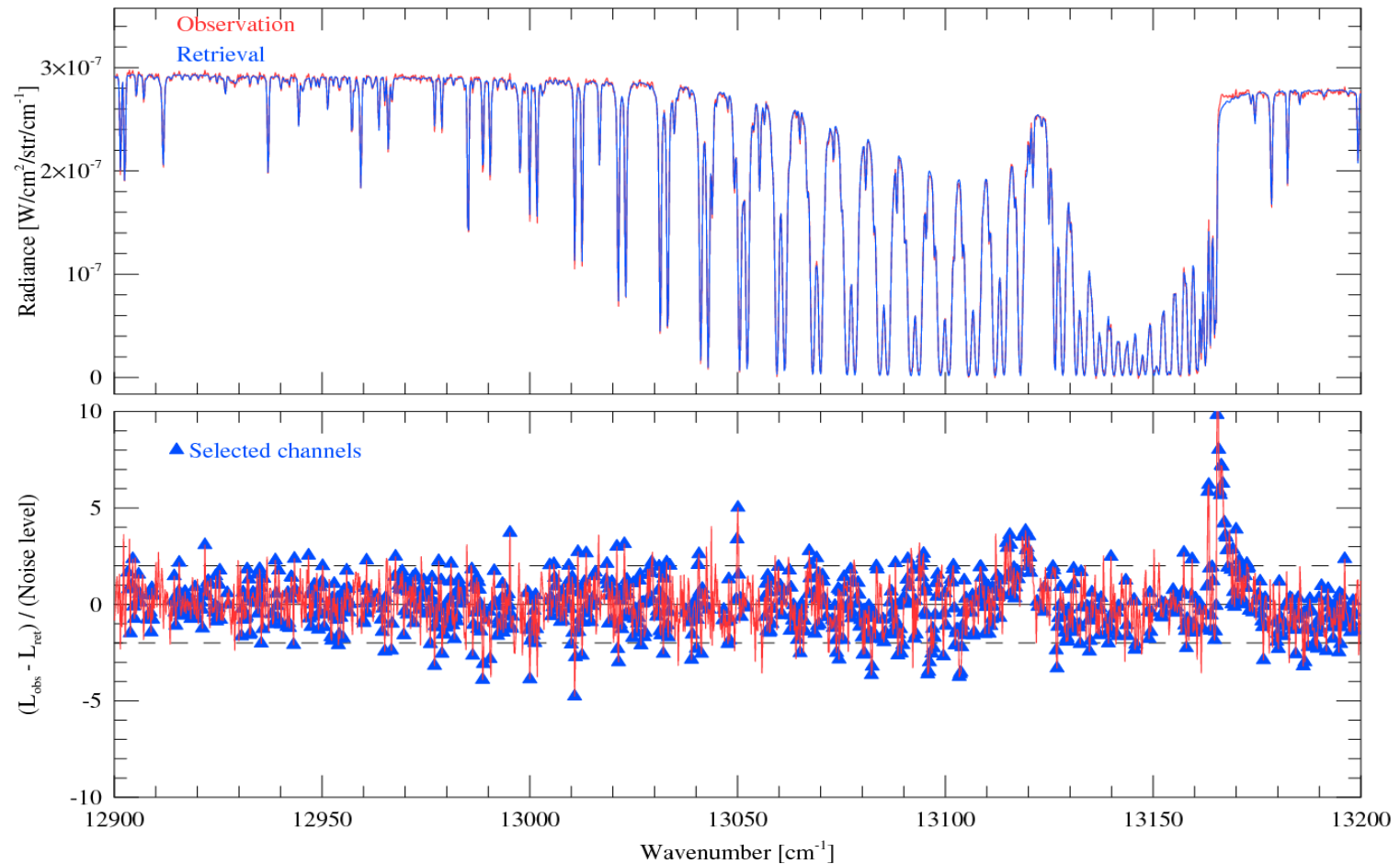
$P_s = 989.05$  hPa and  $\chi^2 = 1.5073$ .

3) Voigt (without CIA and LM) :

$P_s = 1018$  hPa and  $\chi^2 = 2.5647$ .

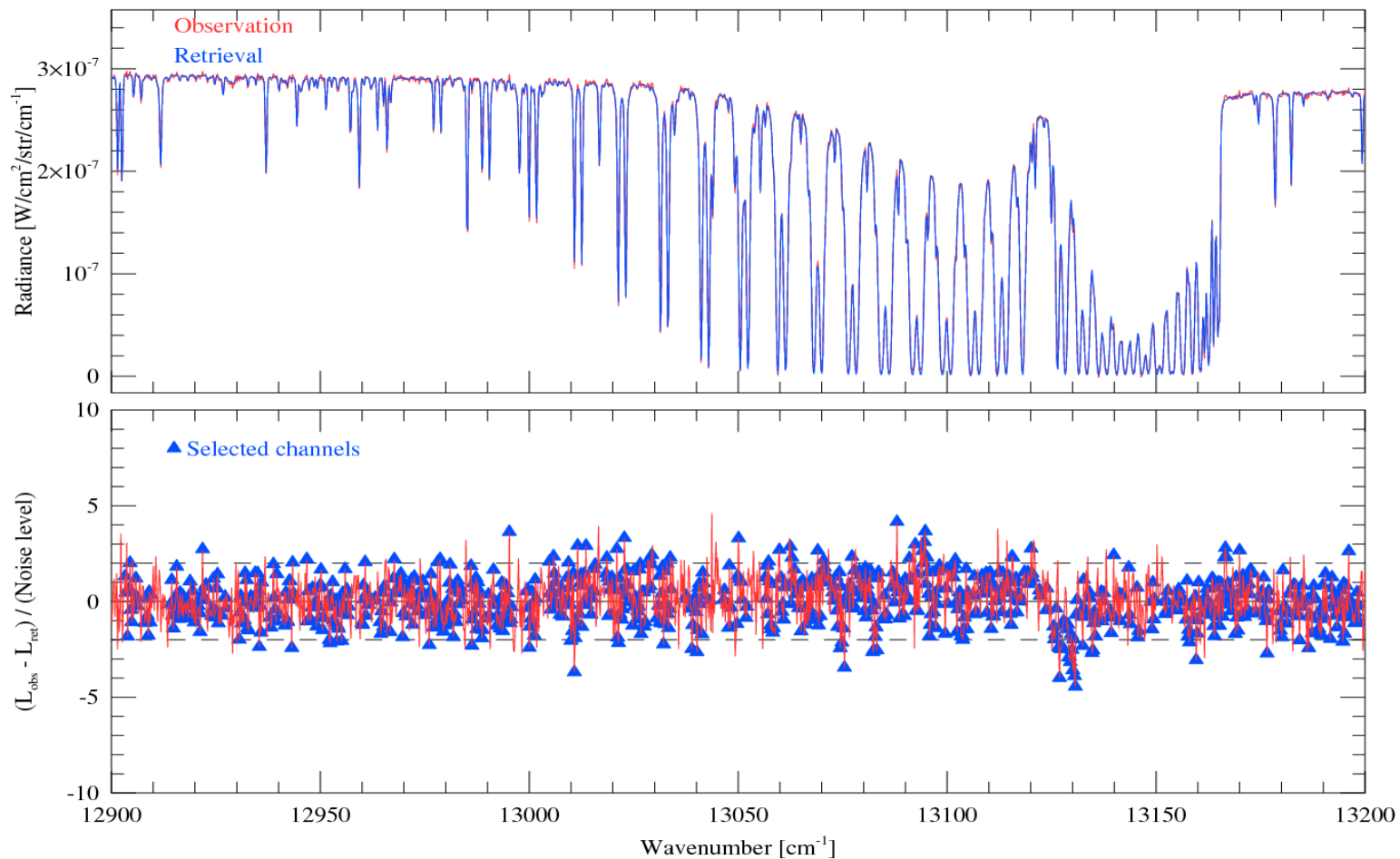
# O<sub>2</sub> A-band: Result with Voigt profiles

Ps = 1018 hPa,  $\chi^2 = 2.5647$



# O<sub>2</sub> A-band: Result with LM+CIA codes

Ps = 989 hPa,  $\chi^2 = 1.5038$





# Line Mixing combination Tests for Retrievals

- H<sub>2</sub>O: HITRAN2008 (2009 update)
- CH<sub>4</sub>: Lyulin et al. (2009)
- Comparison among the following three cases

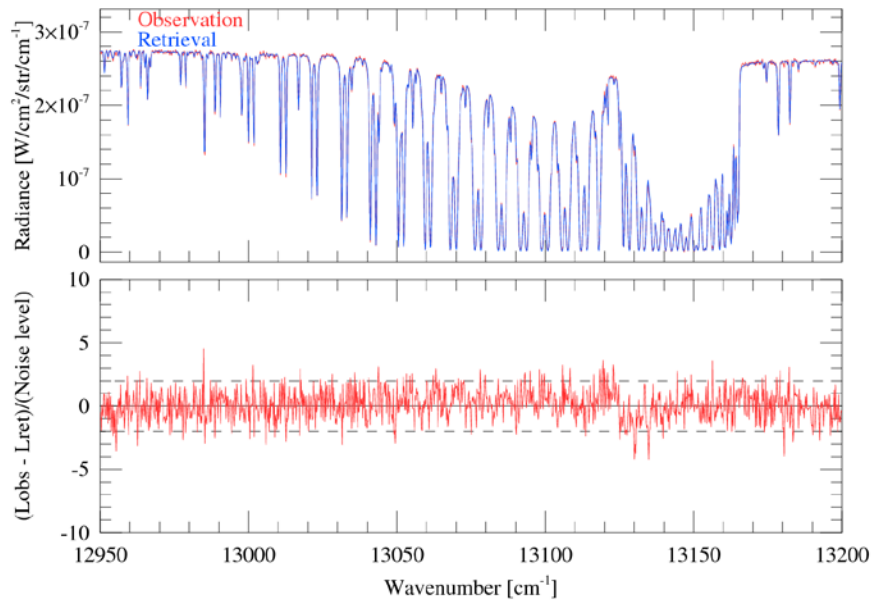
	O2 – A Band	Weak CO2	Note
Case 0	Tran LM+CIA	Hartmann	
Case 1	HITRAN2008	Hartmann	Without CIA
Case 2	Tran LM+CIA	HITRAN2008	

# Retrieval conditions

- TANSO-FTS Band 1 & 2
- Simultaneous retrieval of H<sub>2</sub>O, CO<sub>2</sub>, CH<sub>4</sub>, aerosol optical thickness (AOT), ground surface albedo, and irradiance correction coefficient
- Assumed cloud free and no cirrus contaminated
- 2009/04/26 – 28, over the land, (346 scans)

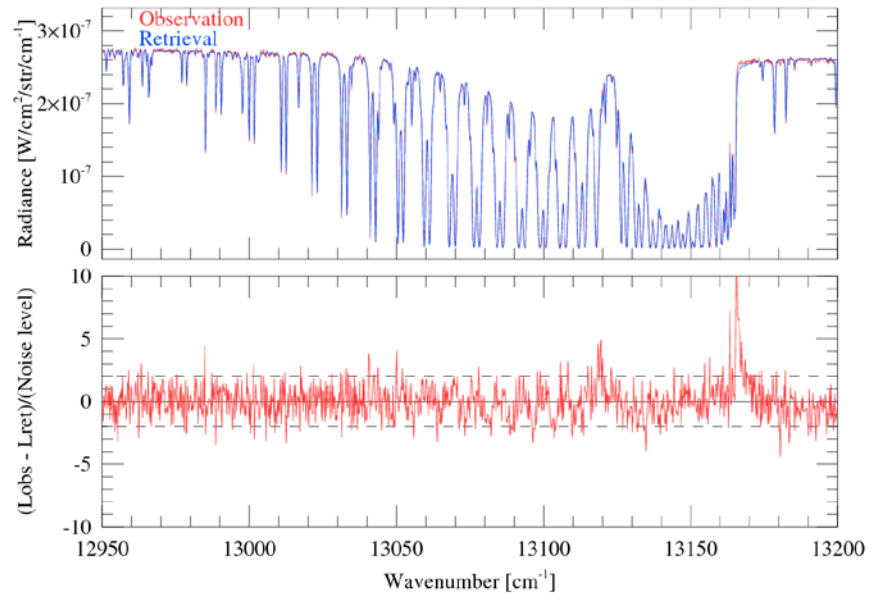
# Spectral residuals after retrievals

## Case 0



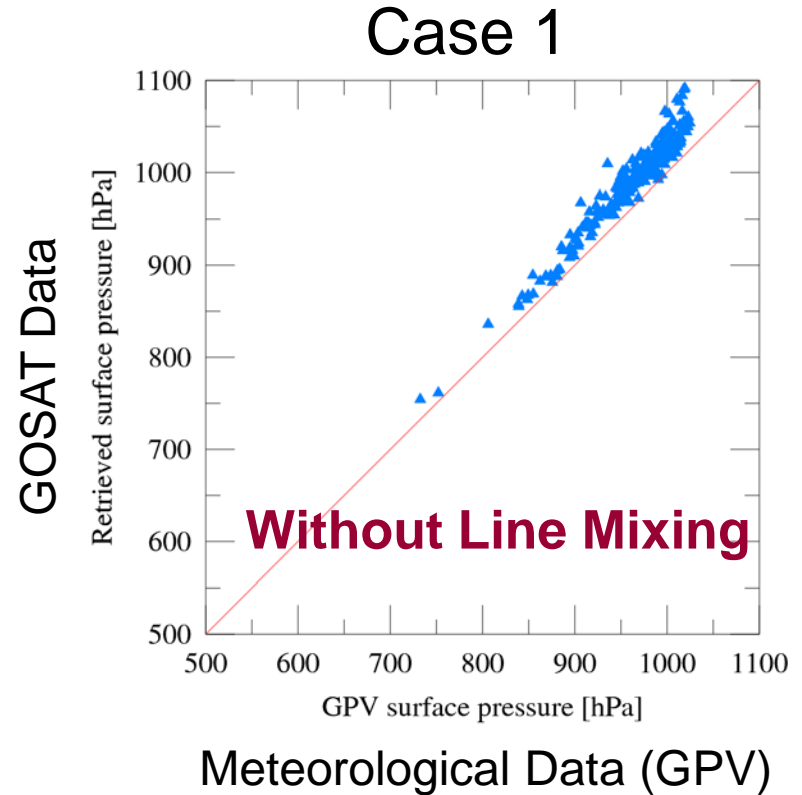
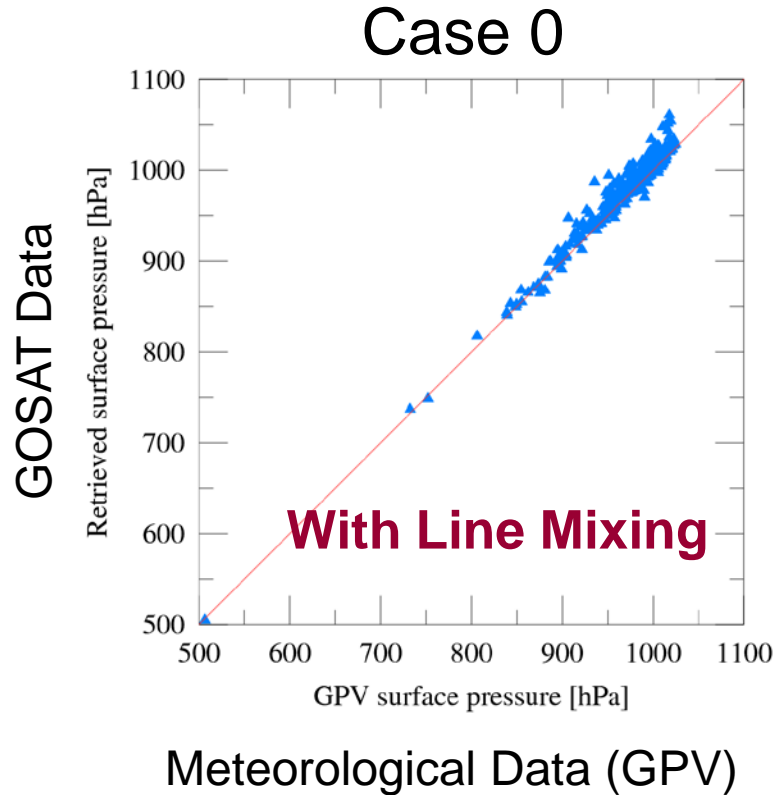
**With Line Mixing**

## Case 1



**Without Line Mixing**

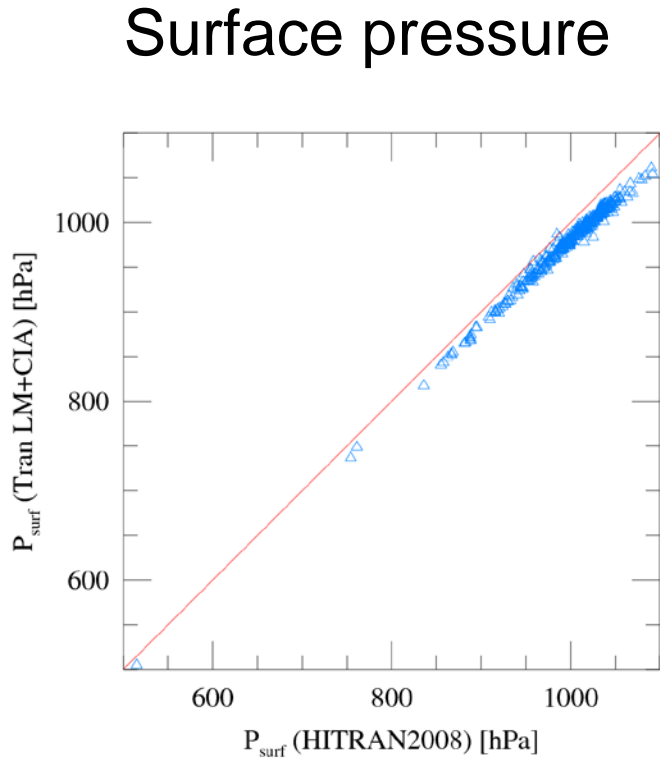
# Surface pressure (Meteor. vs Retrieved)



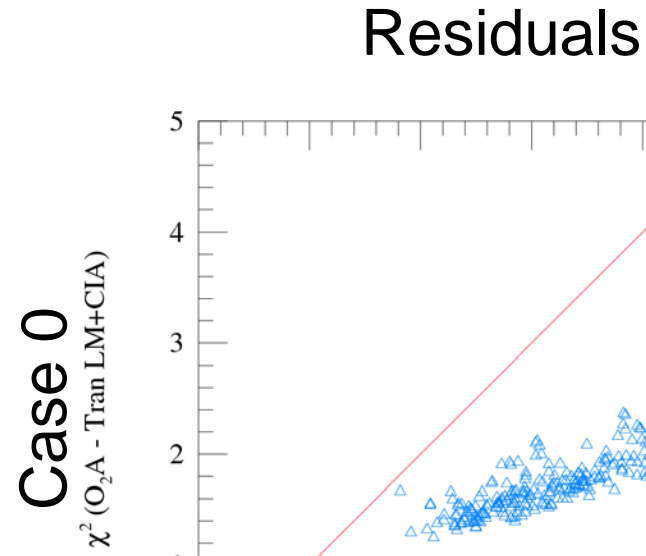
# Case 0 vs. Case 1

With Line Mixing in O<sub>2</sub> A-band

Case 0



Case 1



Case 1

Without Line Mixing in O<sub>2</sub> A-band

# Line Mixing combination Tests for Retrievals

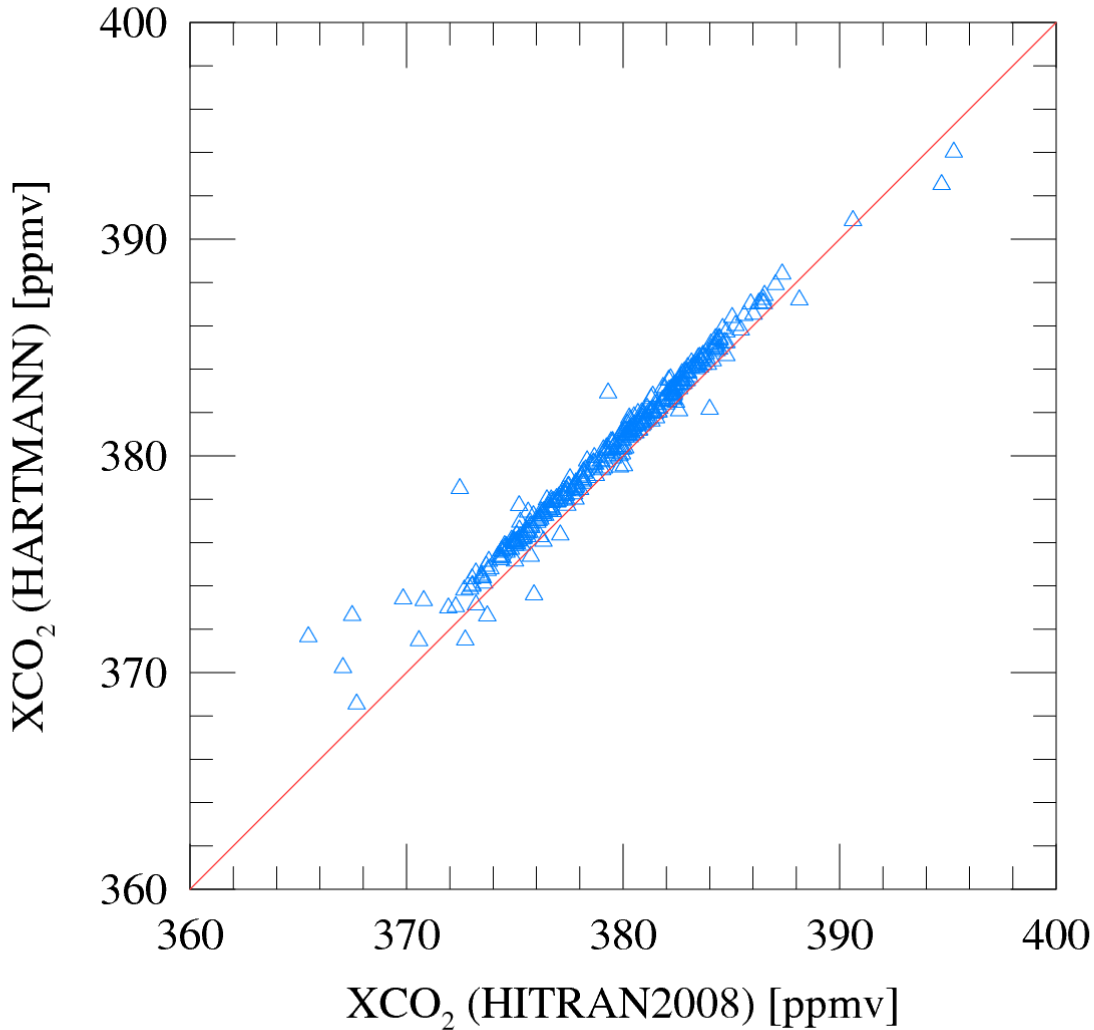
- H<sub>2</sub>O: HITRAN2008 (2009 update)
- CH<sub>4</sub>: Lyulin et al. (2009)
- Comparison among the following three cases

	O2 – A Band	Weak CO2	Note
Case 0	Tran LM+CIA	Hartmann	
Case 1	HITRAN2008	Hartmann	Without CIA
Case 2	Tran LM+CIA	HITRAN2008	



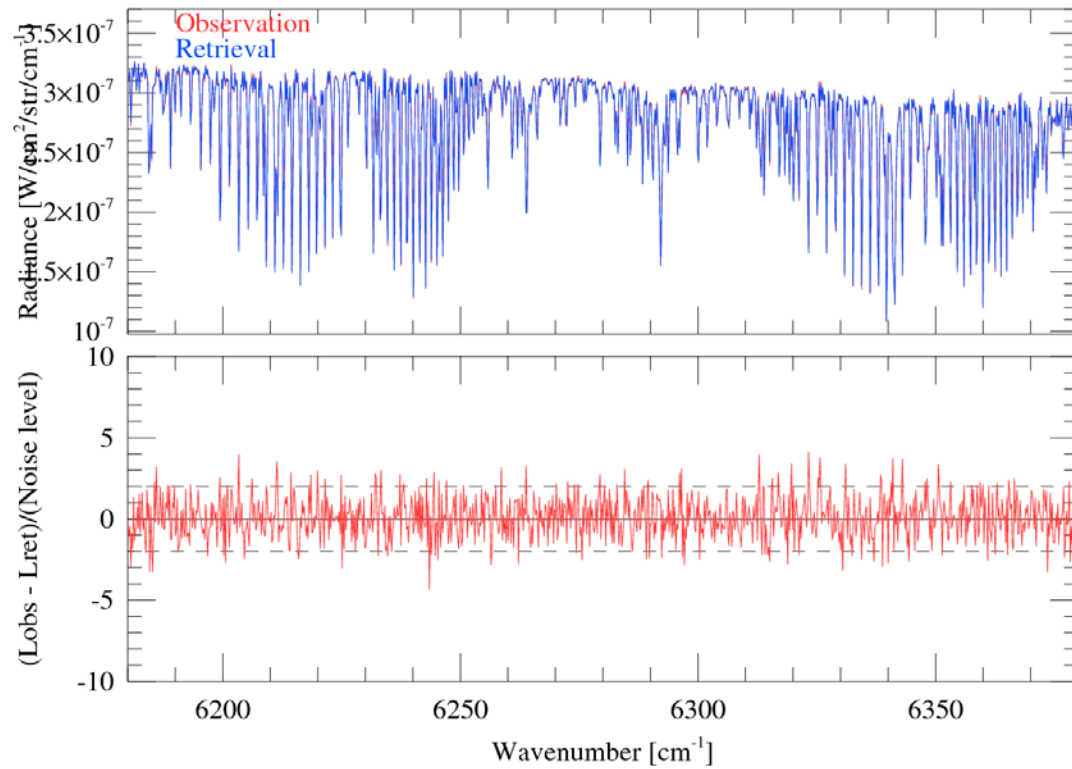
# XCO<sub>2</sub> : Case 0 vs. Case 2

**With Line Mixing in weak CO<sub>2</sub> band**

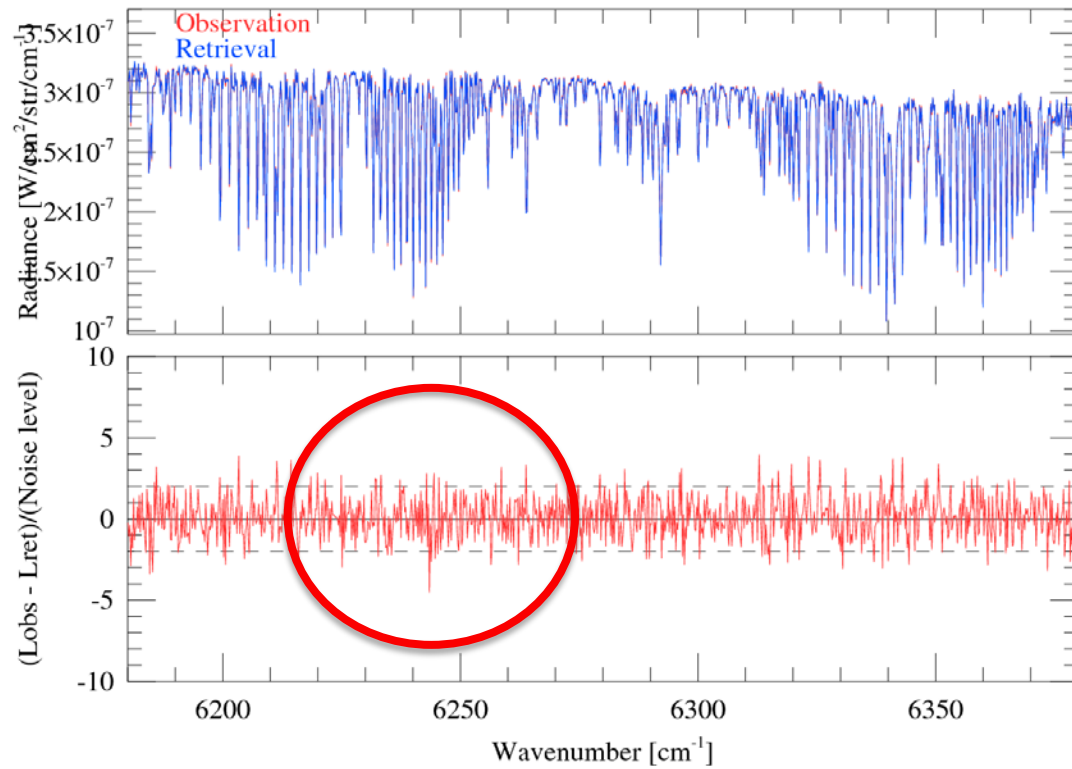


**Without Line Mixing in weak CO<sub>2</sub> band**

# Case 0: spectral residuals



# Case 2: spectral residuals



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# Announcement



# GOSAT Research Activities

## ■ Research Announcement

### ➤ Research Topics

- 1) Calibration
- 2) Data Processing Algorithm
- 3) Validation
- 4) Carbon Balance Estimation and Atmospheric Transport Models
- 5) Data application

### ➤ The 1st GOSAT RA in 2008

52 research themes were selected.

### ➤ The 2nd GOSAT RA in 2009

36 research themes were selected.

### ➤ The 3rd GOSAT RA in August 2010 (Planned)

## ■ Data Release

- CAI & FTS-L1B data product: October 30, 2009
- CAI & FTS-L2 data product: February 18, 2010
- **Improved FTS-L2 data product: August, 2010 (Planned)**



National Institute for  
Environmental Studies  
(NIES)



Ministry of the Environment  
(MOE)

## GOSAT Project

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service are here](#)



Global Greenhouse Gas Observation by Satellite

## GOSAT Project

[http://www.gosat.nies.go.jp/index\\_e.html](http://www.gosat.nies.go.jp/index_e.html)

NIES organized the research team dedicated to the GOSAT project within its organization in April 2004, and since then has been working for the research and development with respect to GOSAT "IBUKI".

### News

- ▶ (April 14,2009)  
["GOSAT Project" pamphlet](#) was revised.
- ▶ (April 7,2009)  
[We have just launched the 2nd Research Announcement on GOSAT.](#)
- ▶ (April 1,2009)  
We made the "[News](#)" page.
- ▶ (February 13,2009)  
[The first data of GOSAT were acquired. The results are shown on following JAXA's home page.](#)
- ▶ (February 9,2009)  
[GOSAT path calendar has been disclosed.](#)
- ▶ (February 6,2009)  
[GOSAT Data Policy has been disclosed.](#)



(If you click the above image, and it opens in another window.)

- GOSAT Level 2 data products (Vers. 00.xx) of CO<sub>2</sub> and CH<sub>4</sub> have negative biases of 2 – 3%, however the overall tendency of global seasonal variations seems good.
- Some part of the retrieval results are affected by dust aerosols.
- Precise line parameters and line shape calculation are quite important to retrieve XCO<sub>2</sub> and XCH<sub>4</sub> correctly.
- Some line parameters will be changed, and line mixing for O<sub>2</sub>-A band in forward calculation will be included in the next version of the algorithm.