

# An Improved Version of the CO<sub>2</sub> Line-mixing Database and Software: Update and Extension



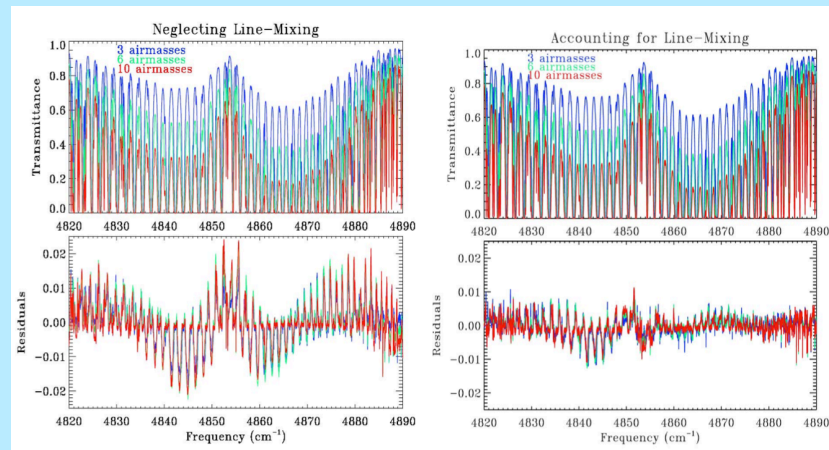
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## The importance line-mixing for CO<sub>2</sub>



From Hartmann *et al.* (2009). Residuals obtained from fits of ground-based atmospheric solar-absorption spectra in the region from 4765 to 4915 cm<sup>-1</sup> for various airmasses.

## Previous work

- In a previous study, Niro *et al.* (2005) presented a database and software package which provided for the calculation of the absorption coefficients of CO<sub>2</sub> that included the effects of line-mixing.
- J.-M. Hartmann *et al.* (2009) concluded that the CO<sub>2</sub> database and software of Niro *et al.* needed to be updated and improved. The present work provides this update.

The present work was driven by the need for a line-mixing database consistent with the CO<sub>2</sub> database on HITRAN. The new database being presented goes hand-in-hand with HITRAN2008.

# Line-mixing

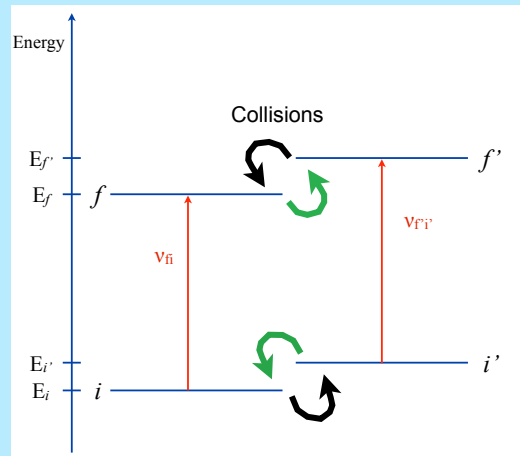


Illustration of line-mixing between two optical transitions,  $f \leftarrow i$  and  $f' \leftarrow i'$  (Hartmann *et al.* 2008).

## The previous database

Spectroscopic files – data files for each band

Bandinfo – statistical information for each band

Software – computes the absorption coefficients

The off-diagonal elements of the relaxation matrices,  $W(T)$ , are computed, and from the  $W$ 's the line-mixing coefficients,  $Y_k$ , are calculated. The  $Y$ 's are then used to calculate absorption coefficients.

## Procedure for the construction of the new database

- Took HITRAN2008, CDS-296 and CDS-1000 and separated into files by isotopologue and by vibrational band.
- Checked for  $J_{\max} \geq 70$  in HITRAN, if not then checked for the missing transitions in CDS-296, and in CDS-1000 (if not in CDS-296).
- For bands where the  $J_{\max} \geq 70$  condition could not be met with the inclusion of the 3 databases, line-mixing effects were not included in the calculation of absorption coefficients. Note that these bands are weak bands and are for minor isotopologues.
- For each band with  $J_{\max} \geq 70$ , the same procedure from the previous model is used to calculate absorption coefficients.

## The new line-mixing database

Previous version (Niro <i>et al.</i> , 2005)	Updated version (Lamouroux <i>et al.</i> , 2010)
HITRAN 2000 (Rothman <i>et al.</i> , 2003)	HITRAN 2008 (Rothman <i>et al.</i> , 2009) CDSD-296 (Tashkun <i>et al.</i> , 2003) CDSD-1000 (Tashkun <i>et al.</i> , 2008)
~ 600 vibrational bands	~ 3600 vibrational bands
~ 78,000 lines	~ 450,000 lines
Half-widths	
CO <sub>2</sub> -CO <sub>2</sub> CO <sub>2</sub> -air	CO <sub>2</sub> -CO <sub>2</sub> (Rothman <i>et al.</i> , 1992) CO <sub>2</sub> -air (Rothman <i>et al.</i> , 1992) CO <sub>2</sub> -H <sub>2</sub> O (Sung <i>et al.</i> , 2009) mass dependence (Lamouroux <i>et al.</i> , 2010)
Temperature Dependence	
CO <sub>2</sub> -air	CO <sub>2</sub> -air (Rothman <i>et al.</i> , 1992) CO <sub>2</sub> -H <sub>2</sub> O (Sung <i>et al.</i> , 2009)
Line Shifts	
Some CO <sub>2</sub> -air	CO <sub>2</sub> -air (J.-M. Hartmann, 2009)

## Example of the new database - Spectroscopic files

File name:  
S10001100001.dat

21	2206.530510	8.450E-30	1.836E+02	05500.051	6016.39980	41-.004369	0 0 0 11	0 0 0 01	P124e	4405521221	1 1 1 5	247.0	249.10534	538	0.70568546D+00	0.17292991D-12
21	2209.545610	2.100E-29	1.840E+02	05500.051	5825.63690	41-.004348	0 0 0 11	0 0 0 01	P122e	4405521221	1 1 1 5	243.0	245.10694	547	0.70566223D+00	0.43006583D-12
21	2212.537920	5.150E-29	1.847E+02	05520.051	5637.90010	42-.004326	0 0 0 11	0 0 0 01	P120e	4445521221	1 1 1 5	239.0	241.10864	555	0.70563821D+00	0.10536490D-11
21	2215.507400	1.240E-28	1.848E+02	05520.051	5453.19240	43-.004304	0 0 0 11	0 0 0 01	P118e	4445521221	1 1 1 5	235.0	237.11020	562	0.70561342D+00	0.25429797D-11
21	2218.454020	2.960E-28	1.861E+02	05520.052	5271.51690	45-.004282	0 0 0 11	0 0 0 01	P116e	4445521221	1 1 1 5	231.0	233.11180	570	0.70558776D+00	0.60459587D-11
21	2221.377740	6.940E-28	1.868E+02	05520.052	5092.87650	46-.004259	0 0 0 11	0 0 0 01	P114e	4445521221	1 1 1 5	227.0	229.11340	578	0.70556118D+00	0.14159646D-10

2004 HITRAN format

$\gamma_{CO_2-H_2O}$

rigid rotor dipole  
matrix element

$d_k^0$

$n_{CO_2-H_2O}$

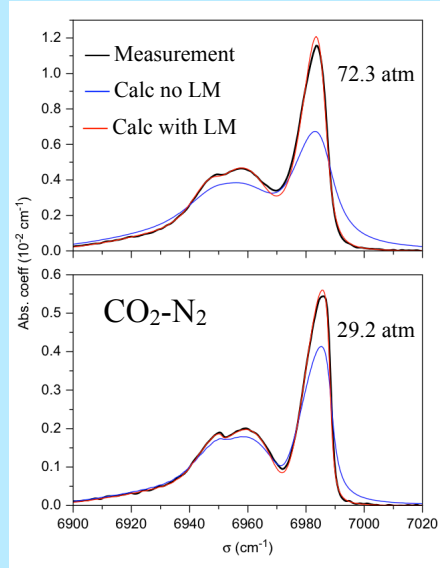
$\rho$   
population

The user will receive:

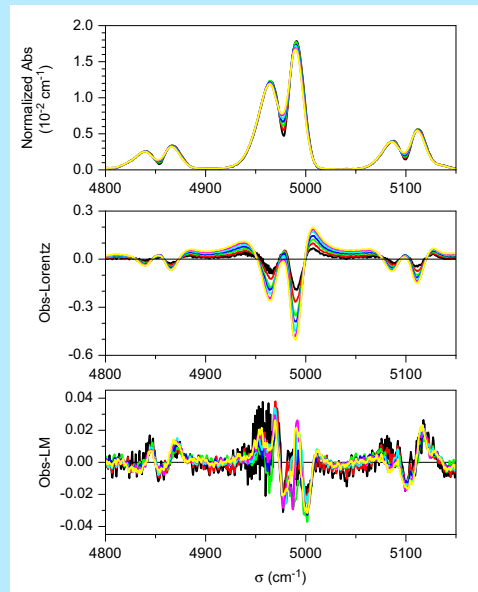
- Database files (like above) and a file with information on each band
- Software to process these files (i.e. calculate absorption coefficients via  $W$  and  $Y_k$ )



## Comparison with measured laboratory spectra



Comparison between measured and calculated  $\text{CO}_2$  absorption coefficients in the  $3\nu_3$  band. The black line is measured values, red and blue are calculated values from the new version of the database and software taking line mixing into account and neglecting line-mixing, respectively. (Lamouroux *et al.*, 2010)



Absorption coefficients and measured-calculated residuals for 1% CO<sub>2</sub> in seven pressures (from 19 to 76.0 atm) of N<sub>2</sub> at 295.15 K in the region of the ( $2\nu_1 + \nu_3$ ) triad. The top panel displays the measured absorptions normalized to unity area through integration over  $\sigma$ . The middle and lower panels display the residuals (meas-calc) obtained from calculations with our new database, respectively, neglecting line-mixing and taking this process into account (Lamouroux *et al.*, 2010).

## Conclusions

- The new line-mixing database includes the best available data for CO<sub>2</sub> line-mixing calculations.
- It includes many more lines than the previous Niro *et al.* 2005 database.
- More parameters are included such H<sub>2</sub>O-broadening parameters of CO<sub>2</sub>.
- The new database is consistent with HITRAN2008.
- The calculated absorption coefficients agree very well with measurement.
- Important for GOSAT, OCO-2, ACE-FTS, and any other high-precision remote sensing measurements.

## Future research

- Software is being tested on atmospheric spectra simulations.
- Procedure for creating future CO<sub>2</sub> line-mixing database has been somewhat automated so that updates corresponding to future HITRAN editions can be easily done.
- Go from the empirically determined values of half-width and line shift used in this work to experimental/calculated values.

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