

October 13, 2010



Flux Density Monitoring with the SMA

Mark A. Gurwell



A Unique Resource



I. The SMA Calibrator Database

What is it, and what is is used for?

II. Blazar Observations and Multi-Wavelength Campaigns

A few examples

III. Independent Programs



I. What is the Calibrator Database?

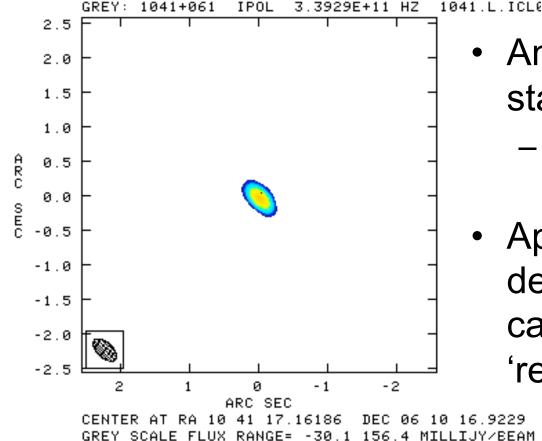


- available, relatively bright sources
 - effectively point sources at resolutions offered by the SMA (sizes <<100 mas)
- nearly exclusively quasars
 - flux densities from 100 mJy to >40 Jy at 1.3 mm
- used to calibrate science observations
 - Phase gains ('seeing')
 - Amplitude gains ('approximate flux scale')



Quasars as Calibrators





- Analogous to guide stars in adaptive optics
 - Timescales generally significantly longer
- Apply corrections determined on calibrator to target to 'remove atmosphere'



Rationale for the Database



- Positions of quasars (very) well known
 - VLA, VLBA positions used, typical $\Delta \sim$ few mas
- Flux densities are variable
 - Often by large factors
 - Timescales of < 1 day through years
- Unlike stock market, past performance is (somewhat) indicative of future value
 - Strong quasars usually strong, but still uncertain

Desirable to (accurately) track calibrator flux densities for planning, executing and calibrating science observations



SMA Calibrator Database



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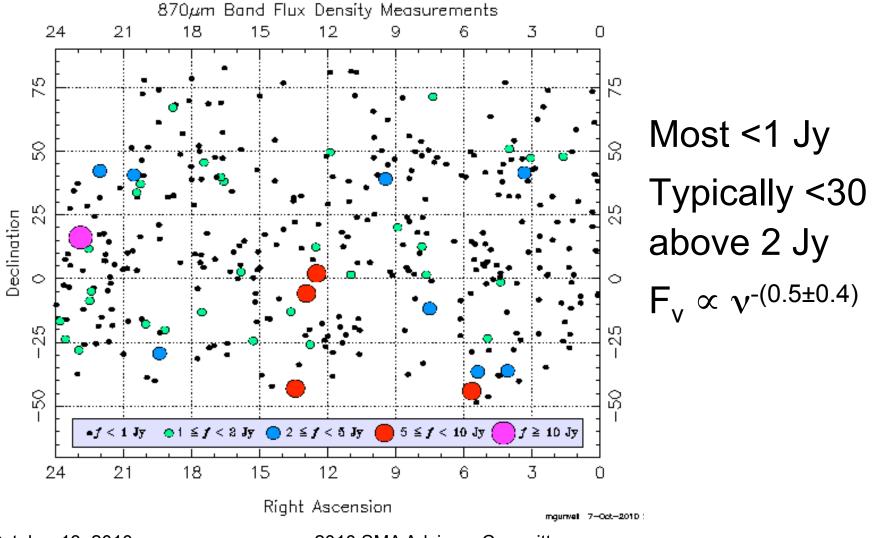
- Data on >400 quasars
- Flux history accessible online
- 1.35 mm, 1.1 mm and 850 μm SMA measurements
- Flux scale tied to solar system 'primary' calibrators (Uranus, Neptune, Mars, Titan, Ganymede, Callisto)

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Sky Distribution





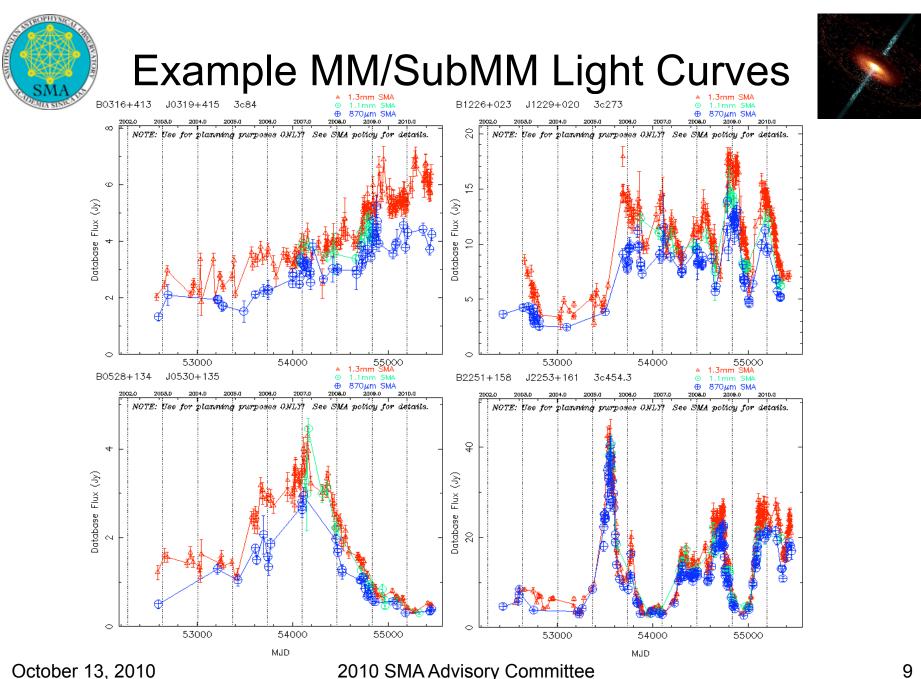


Size and Growth of Database



SMA Calibrator Database Flux Density Measurements Projected 2000 Current 1000 o 2002 2003 2004 2005 2006 2007 2008 2009 2010 Year

- Over 14000 measurements
- Adding ~2400/yr
- From regular science tracks and dedicated 'flux' and baseline calibration observations





II. Blazar/AGN Observations





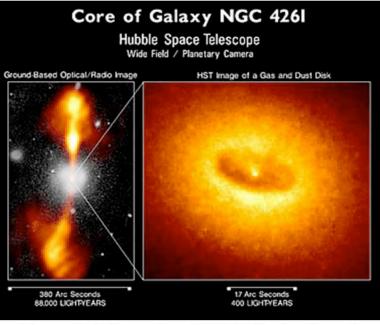
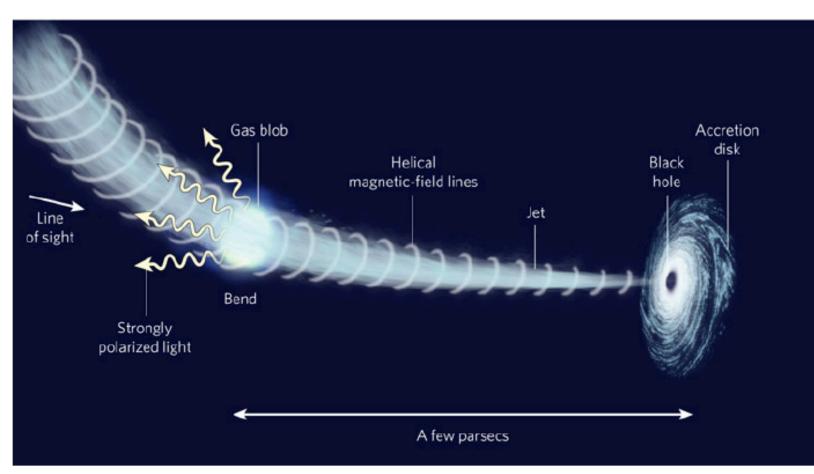


Figure 5: Active galaxy NGC 4261 at radio and optical wavelengths.



Blazar Model



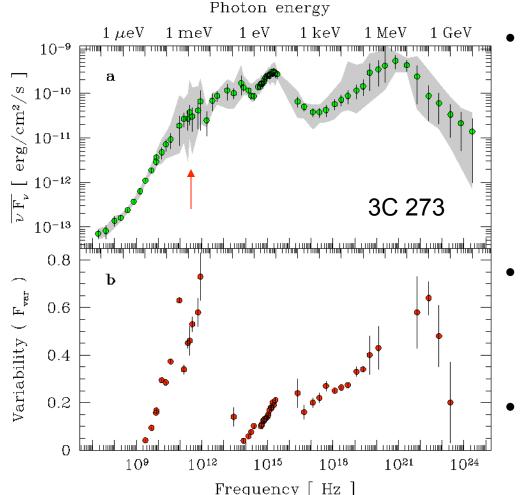


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Quasar/blazar SEDs





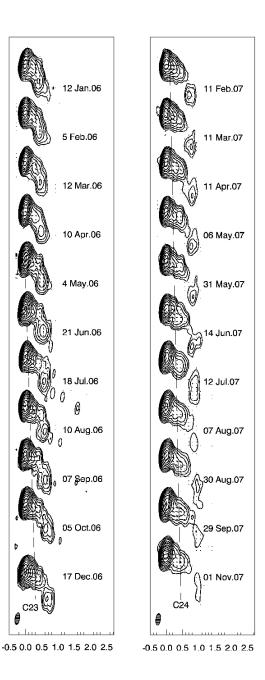
- Quasar/blazar emission
 - bimodal SED
 - Low peak synchrotron,
 high peak inverse
 Compton scattering
- Flare time delay related to relative location of emitting regions
- submm is closest band usable from ground for low end of synchrotron peak



3C279 -Radio Flare

VLBA monitoring of 3C279

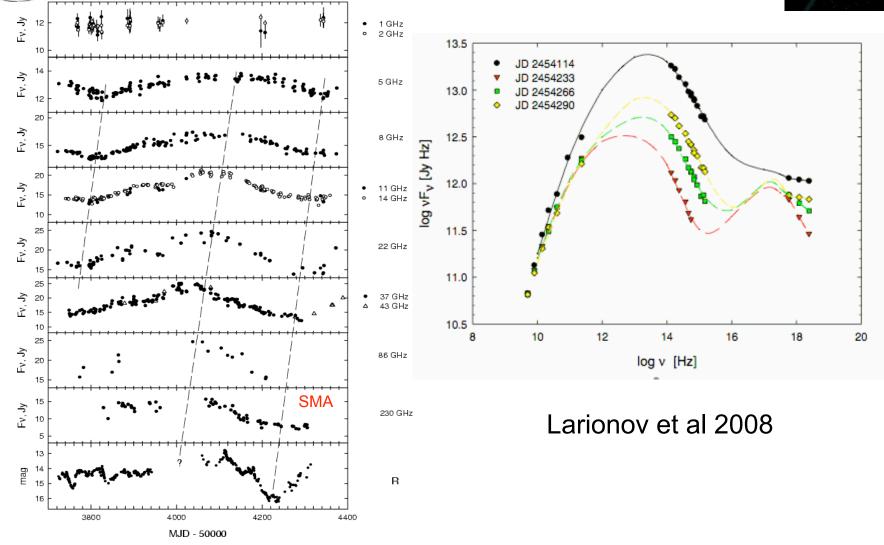
2006-2007 superluminal ejection along jet axis captured







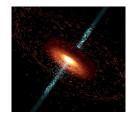
3C279 - Flare Propogation

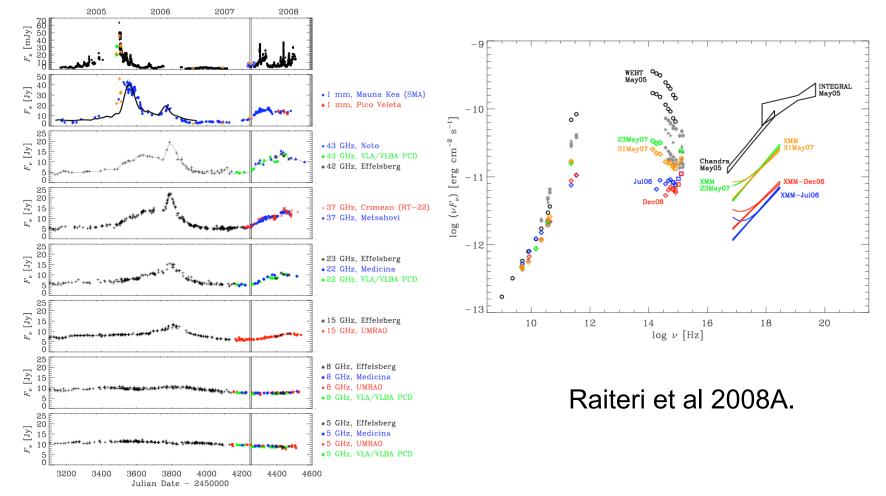


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3C454.3 Flare Propogation

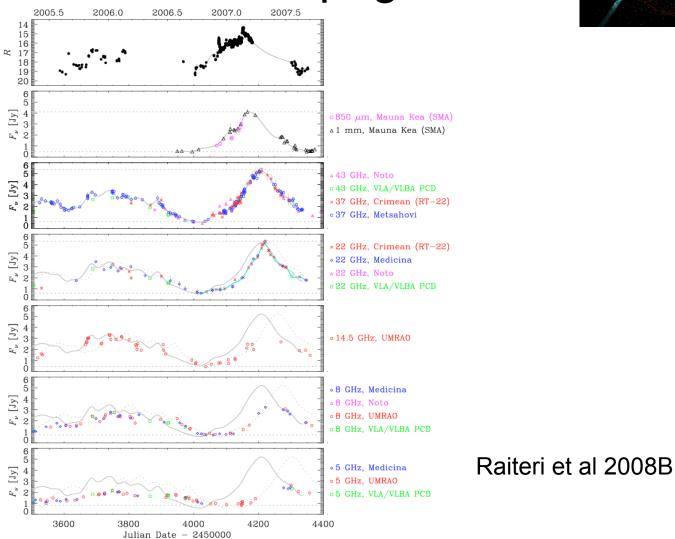




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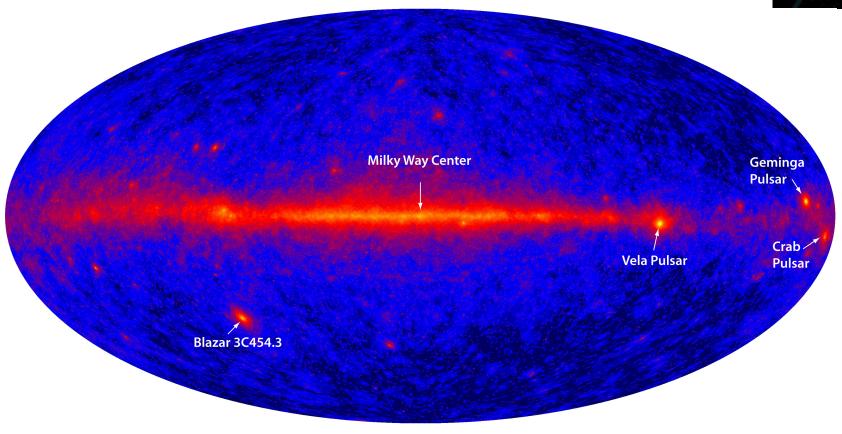


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Blazars Blazing





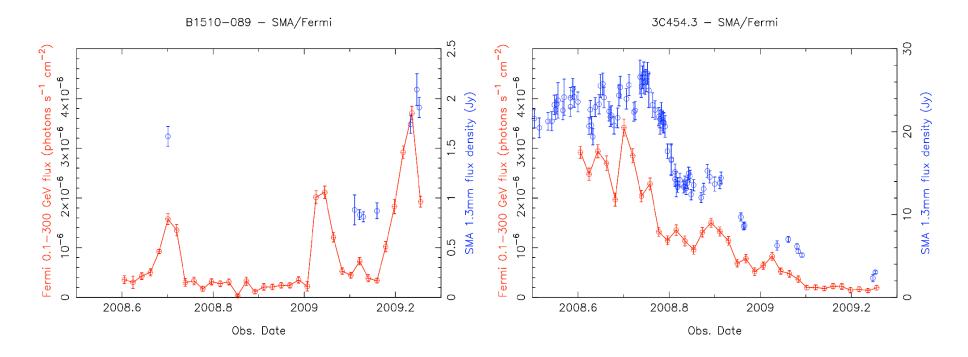
The Gamma Ray Sky - Fermi

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Gamma-ray/mm/submm correlation



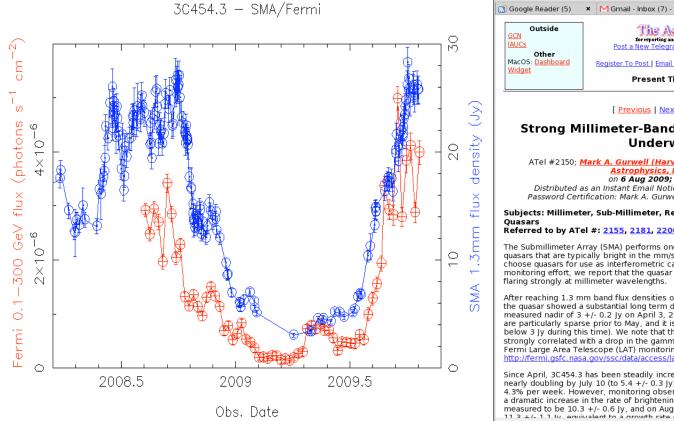


Wehrle et al (ongoing)



Gamma Ray/MM Correlation II





ATel #2150: Strong Millimeter-Band Flaring of 3C454.3 Underway - Mozilla Firefox - 🗆 🗙 Edit View History Bookmarks Tools Help File × M Gmail - Inbox (7) - ... × K FSSC: Data » Data... × 🔬 ATel #2150: Stron... 4 RSS The Astronomer's Telegram Тор for reporting and comm Supernovae Post a New Telegram | Search | Information | Mirror Software Transients Telegram Index SGRs Register To Post | Email and RSS Subscriptions | Forget your password Gamma Ray Present Time: 29 Oct 2009; 18:00 UT Bursts Related [Previous | Next | ADS] 2223 Strong Millimeter-Band Strong Millimeter-Band Flaring of 3C454.3 Flaring of 3C454.3 Continues Underway 2201 Optical behavior of 3C 454.3 in the ATel #2150; Mark A. Gurwell (Harvard-Smithsonian Center for current flaring Astrophysics, MA, US state on 6 Aug 2009; 14:17 UT 2200 Fermi LAT Distributed as an Instant Email Notice (Request for Observations) detection of GeV Password Certification: Mark A. Gurwell (mgurwell@cfa.harvard.edu) flares from 3C454.3 and 3C273 Subjects: Millimeter, Sub-Millimeter, Request for Observations, AGN, 2182 Discovery of a ~205 Hz X-ray Referred to by ATel #: 2155, 2181, 2200, 2223 pulsar in the globular cluster The Submillimeter Array (SMA) performs ongoing flux density monitoring of NGC 6440 guasars that are typically bright in the mm/submm bands, in order to optimally 2181 Optical and IR flare choose guasars for use as interferometric calibration sources. Through this of 3C 454.3 monitoring effort, we report that the quasar 3C454.3 (J2253+161) is currently 2180 NGC 6440 active again 2168 <u>Fermi LAT</u> detection of a GeV After reaching 1.3 mm band flux densities of 27 +/- 1.8 Jy on October 3 2008, the quasar showed a substantial long term drop in intensity, reaching a flare from 3C 273 measured nadir of 3 +/- 0.2 Jy on April 3, 2009 (though observations in 2009 2155 Optical and IR flare are particularly sparse prior to May, and it is possible that the source dipped of blazars 3C 454.3 below 3 Jy during this time). We note that this drop in millimeter flux density is and 3C 279 strongly correlated with a drop in the gamma ray intensity as measured by the 2154 Fermi LAT and Fermi Large Area Telescope (LAT) monitoring program (see INTEGRAL http://fermi.gsfc.nasa.gov/ssc/data/access/lat/msl_lc/_) detection of increasing Since April, 3C454.3 has been steadily increasing in 1.3 mm band flux density, high-energy nearly doubling by July 10 (to 5.4 +/- 0.3 Jy), equivalent to a growth rate of about activity of blazar 4.3% per week. However, monitoring observations over the past 25 days show 3C279 a dramatic increase in the rate of brightening. On August 3, the flux density was 2150 Strong measured to be 10.3 +/- 0.6 Jy, and on August 4 the flux density had grown to illimeter-Band equivalent to a growth rate of 21-23% per week Done





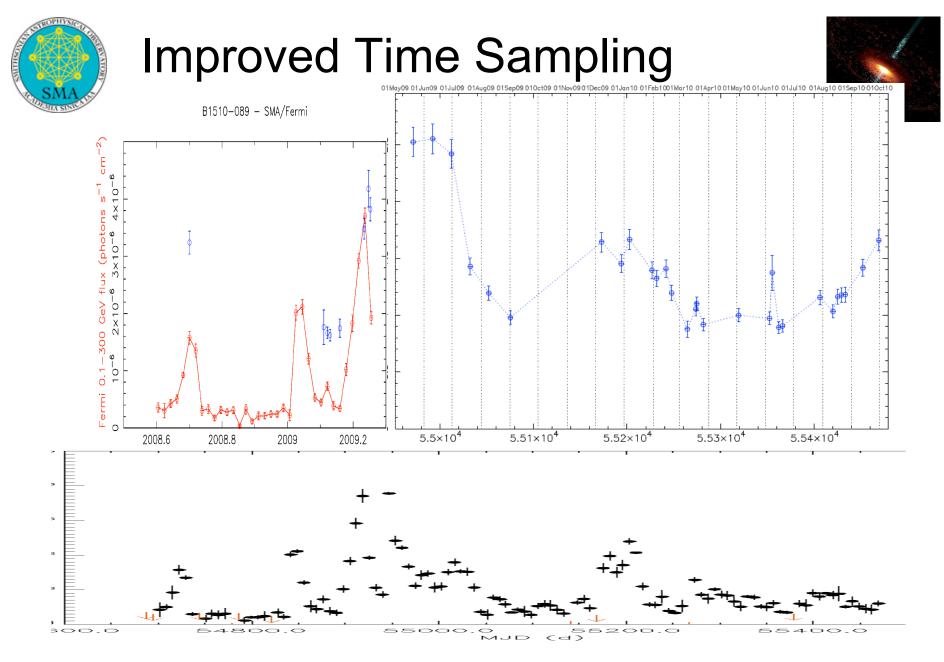
- Currently 16 publications using general SMA flux density monitoring results
- Many more in progress as Fermi observation baseline increases and more flaring sources are discovered



III. Independent Programs



- General flux monitoring not sampled enough in some cases, or not at all in others
- Several proposals accepted for monitoring of specific sources with a regular cadence, with impacts on scheduling
 - Ann Wehrle, Fermi LAT Bright Source catalog
 - Ian McHardy, M81*
 - Aneta Siemiginowska, FSRQs

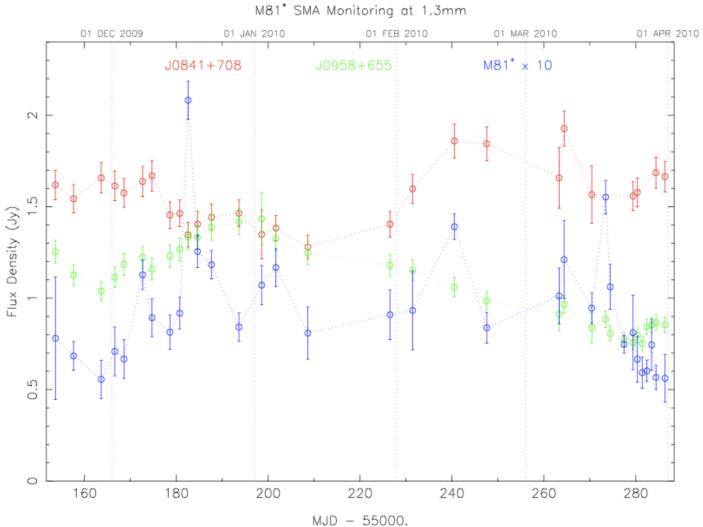


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M81* Monitoring





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2010 SMA Advisory Committee





- SMA Calibrator Database is a nearly unique, accessible resource for the mm/submm flux density history of several hundred quasars
- The SMA data occupy the highest frequency band easily available from the ground for the synchrotron peak in blazars, providing critical information on the shape of the SED of these sources and the location of the emitting regions



