

The Road from **Exploration** to **Explanation**, and Back

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@AlyssaAGoodman

Traditionally, travel from exploration to explanation is called
“Scholarly Publishing” if its *dry*, and “Public Outreach,” if it’s *beautiful*.

Explore



Explain

Explore



Explain

It's much harder to go the other way.

Explore

And, the *best* roads are two-way.

Explain

VIENNA

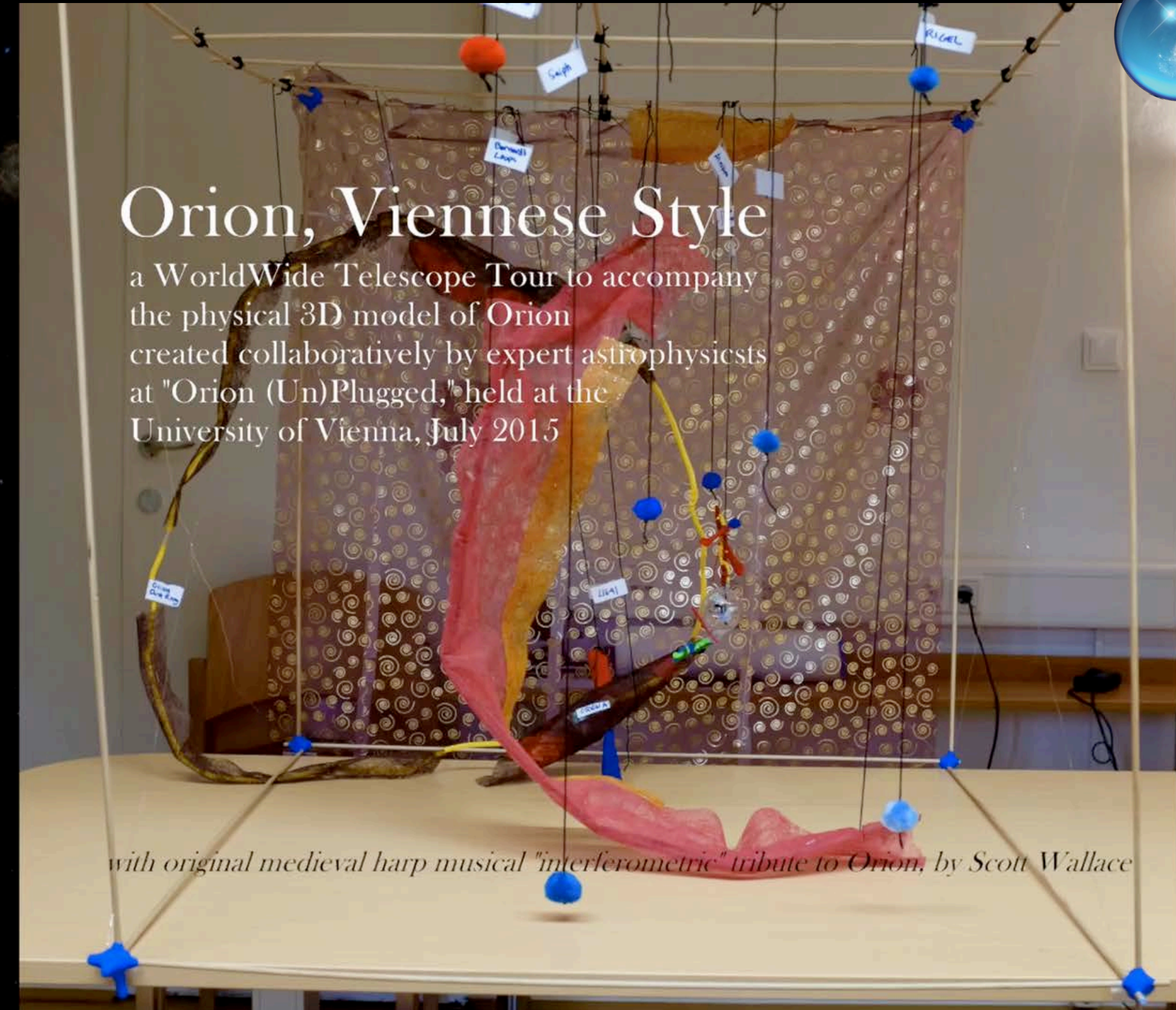
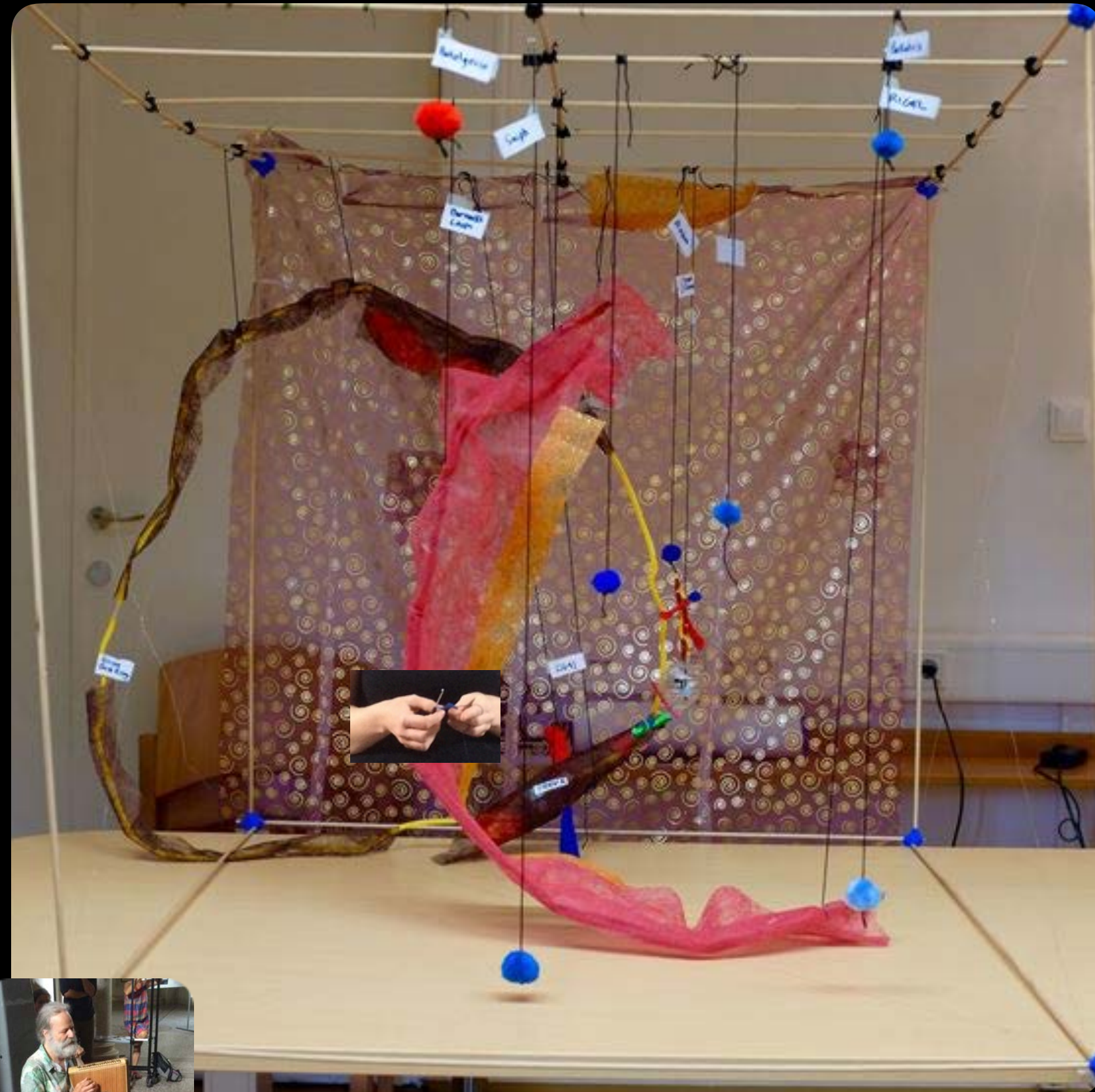


Structure Identification

Human | ✓ | ✓ | ✓ | ✓ | ✓ | ✓ | ✓ | ✓ | ✓ | ✓ | ✓

Available	human	Point Sources	Vel. Coh. Start	Fibers	Filaments	Neuron shape Unbiased	Hier-archical	Clumps + Blobs	Bubbles	Outflows
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✓ CLUMPFIND	✓	(v)	x	x	(v)	x	✓	✓	x	x
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SCIMES			same as dendro but	→ yay!						

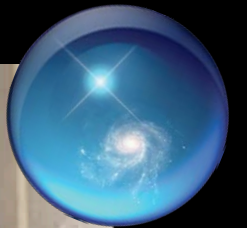
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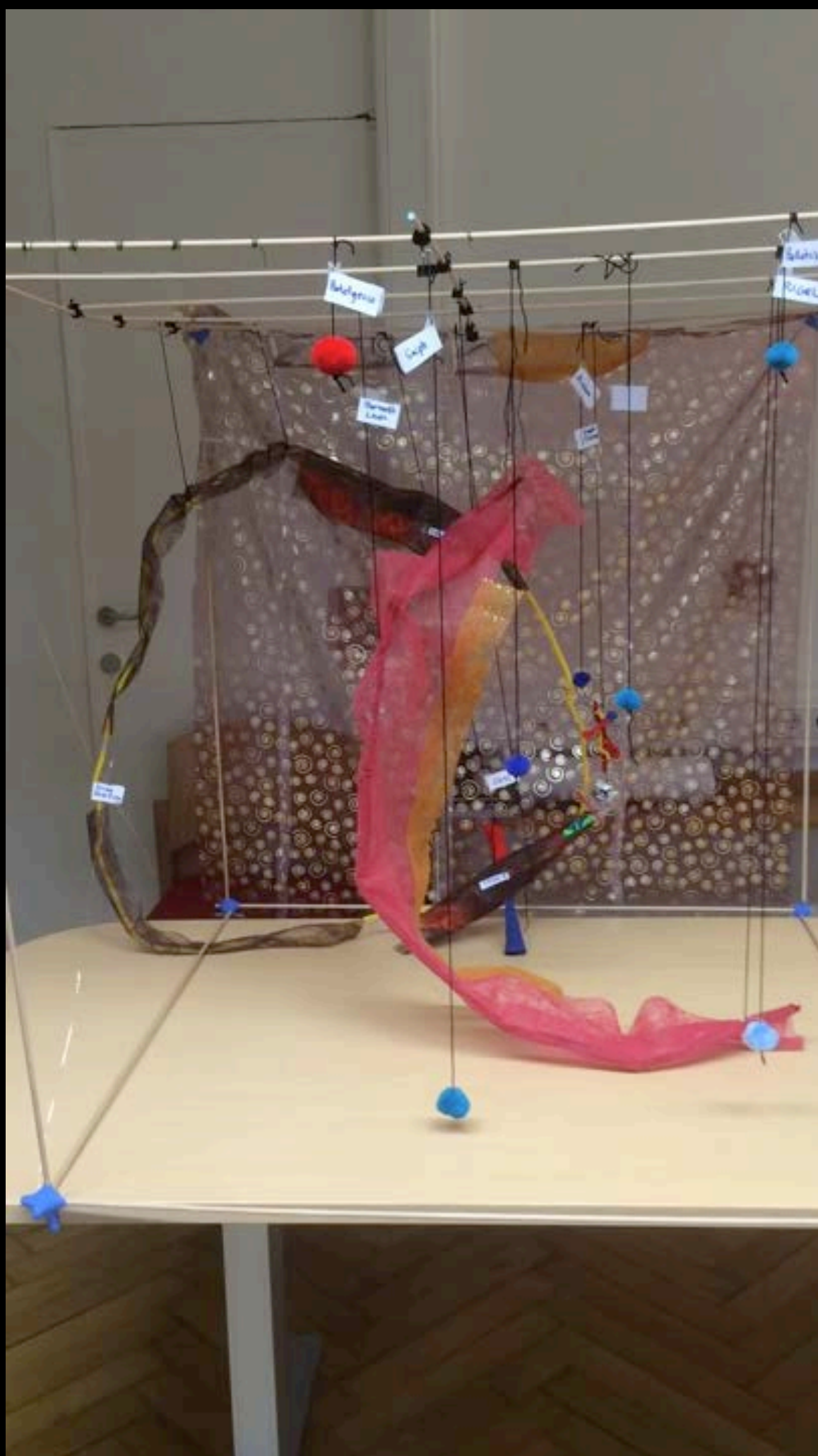
Orion, Viennese Style

a WorldWide Telescope Tour to accompany
the physical 3D model of Orion
created collaboratively by expert astrophysicists
at "Orion (Un)Plugged," held at the
University of Vienna, July 2015

with original medieval harp musical "interferometric" tribute to Orion, by Scott Wallace



VIENNA



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**Astronomy
&
Astrophysics**

VISION – Vienna survey in Orion

I. VISTA Orion A Survey^{★,★★}

Stefan Meingast¹, João Alves¹, Diego Mardones², Paula Stella Teixeira¹, Marco Lombardi³, Josefa Grobschedl¹, Joana Ascenso^{4,5}, Herve Bouy⁶, Jan Forbrich^{1,7}, Alyssa Goodman⁷, Alvaro Hacar¹, Birgit Hasenberger¹, Jouni Kainulainen⁸, Karolina Kubiak¹, Charles Lada⁷, Elizabeth Lada⁹, André Moitinho¹⁰, Monika Petr-Gotzens¹¹, Lara Rodrigues², and Carlos G. Román-Zúñiga¹²

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ABSTRACT

Context. Orion A hosts the nearest massive star factory, thus offering a unique opportunity to resolve the processes connected with the formation of both low- and high-mass stars. Here we present the most detailed and sensitive near-infrared (NIR) observations of the entire molecular cloud to date.

Aims. With the unique combination of high image quality, survey coverage, and sensitivity, our NIR survey of Orion A aims at establishing a solid empirical foundation for further studies of this important cloud. In this first paper we present the observations, data reduction, and source catalog generation. To demonstrate the data quality, we present a first application of our catalog to estimate the number of stars currently forming inside Orion A and to verify the existence of a more evolved young foreground population.

Methods. We used the European Southern Observatory's (ESO) Visible and Infrared Survey Telescope for Astronomy (VISTA) to survey the entire Orion A molecular cloud in the NIR J , H , and K_S bands, covering a total of ~ 18.3 deg². We implemented all data reduction recipes independently of the ESO pipeline. Estimates of the young populations toward Orion A are derived via the K_S -band luminosity function.

Results. Our catalog (799 995 sources) increases the source counts compared to the Two Micron All Sky Survey by about an order of magnitude. The 90% completeness limits are 20.4, 19.9, and 19.0 mag in J , H , and K_S , respectively. The reduced images have 20% better resolution on average compared to pipeline products. We find between 2300 and 3000 embedded objects in Orion A and confirm that there is an extended foreground population above the Galactic field, in agreement with previous work.

THE REAL WORLD

New Ideas
Discoveries

Public Outreach
Scholarly Publishing

Explore

Explain



“It’s much harder to go the other way.”

Astronomy Picture of the Day

[Discover the cosmos!](#) Each day a different image or photograph of our fascinating universe is featured, along with a brief explanation written by a professional astronomer.

2018 March 11



Public Outreach

Duel Particle Beams in Herbig-Haro 24

Image Credit: [NASA](#), [ESA](#), [Hubble Heritage \(STScI/AURA\)](#)/Hubble-Europe Collaboration;
Acknowledgment: D. Padgett ([NASA's GSFC](#)), T. Megeath (U. Toledo), B. Reipurth (U. Hawaii)

Explanation: [This might look like](#) a double-bladed lightsaber, but these two cosmic jets actually beam outward from [a newborn star in a galaxy near you](#). Constructed from Hubble Space Telescope image data, the stunning scene spans about half a light-year across Herbig-Haro 24 (HH 24), some 1,300 light-years away in the [stellar nurseries](#) of the Orion B molecular cloud complex. Hidden from direct view, [HH 24](#)'s central protostar is surrounded by cold dust and gas flattened into a rotating [accretion disk](#). As material from the disk falls toward the young stellar object it heats up. Opposing [jets are blasted out](#) along the system's rotation axis. Cutting through the region's interstellar matter, the narrow, energetic jets produce a series of glowing shock fronts [along their path](#).

March 11, 2018 Astronomy Picture of the Day

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2018 March 11



Duel Particle Beams in Herbig-Haro 24
Image Credit: [NASA](#), [ESA](#), [Hubble Heritage \(STScI/AURA\)](#)/Hubble-Europe Collaboration;
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March 11, 2018 Astronomy Picture of the Day

Herbig-Haro Jet HH 24



Hubble
Heritage

NASA and ESA • Hubble Space Telescope • WFC3/IR • WFPC2 • STScI-PRC15-42a

2015 Hubble Heritage Press Release

Explain

Hubble Heritage

ESA Sky

MAST Portal

MAST Archive

Explore

glueviz

HST Retrieval Submission Success:
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Success! Your HST Request ID: anonymous98615

The following datasets will be retrieved:
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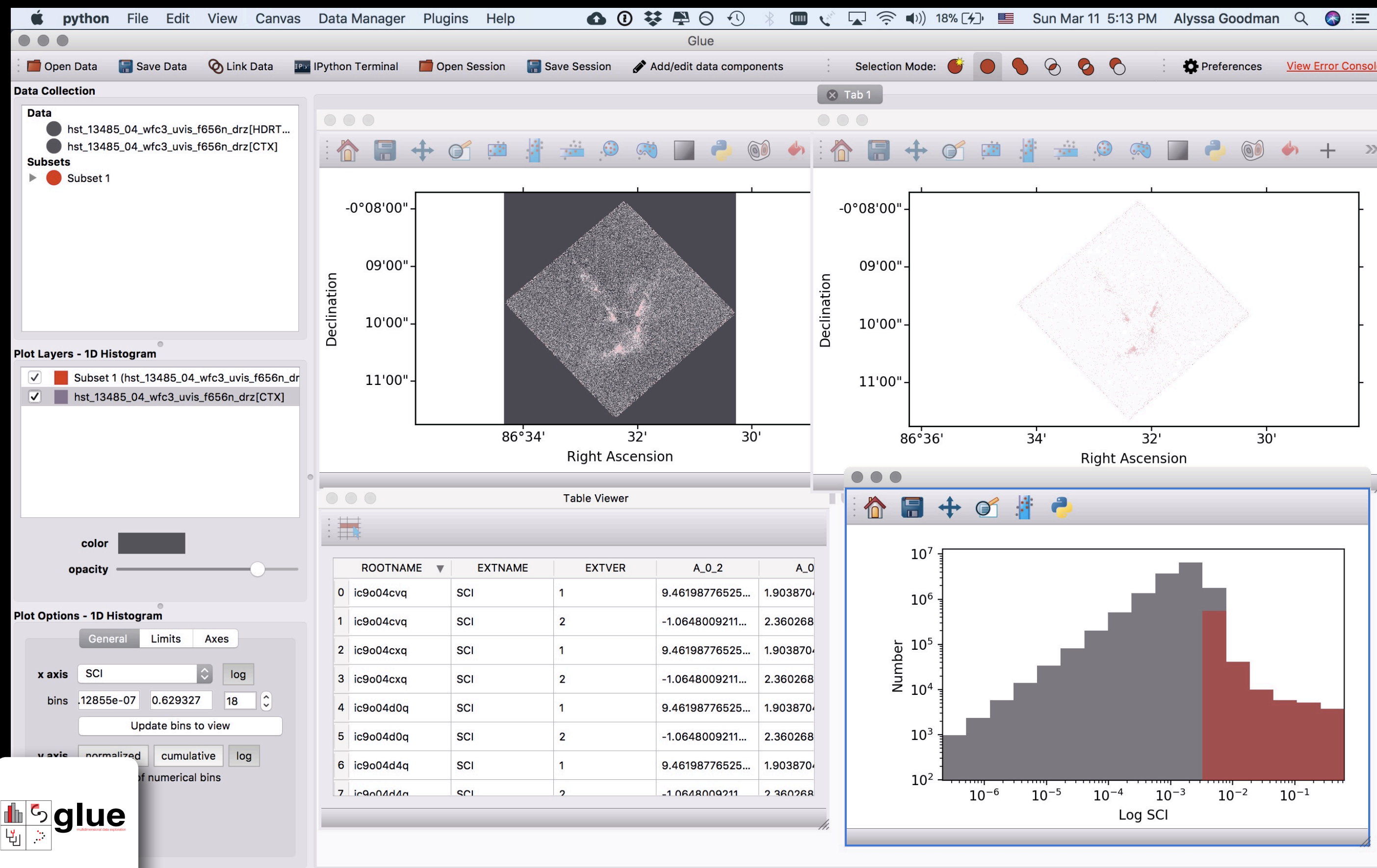
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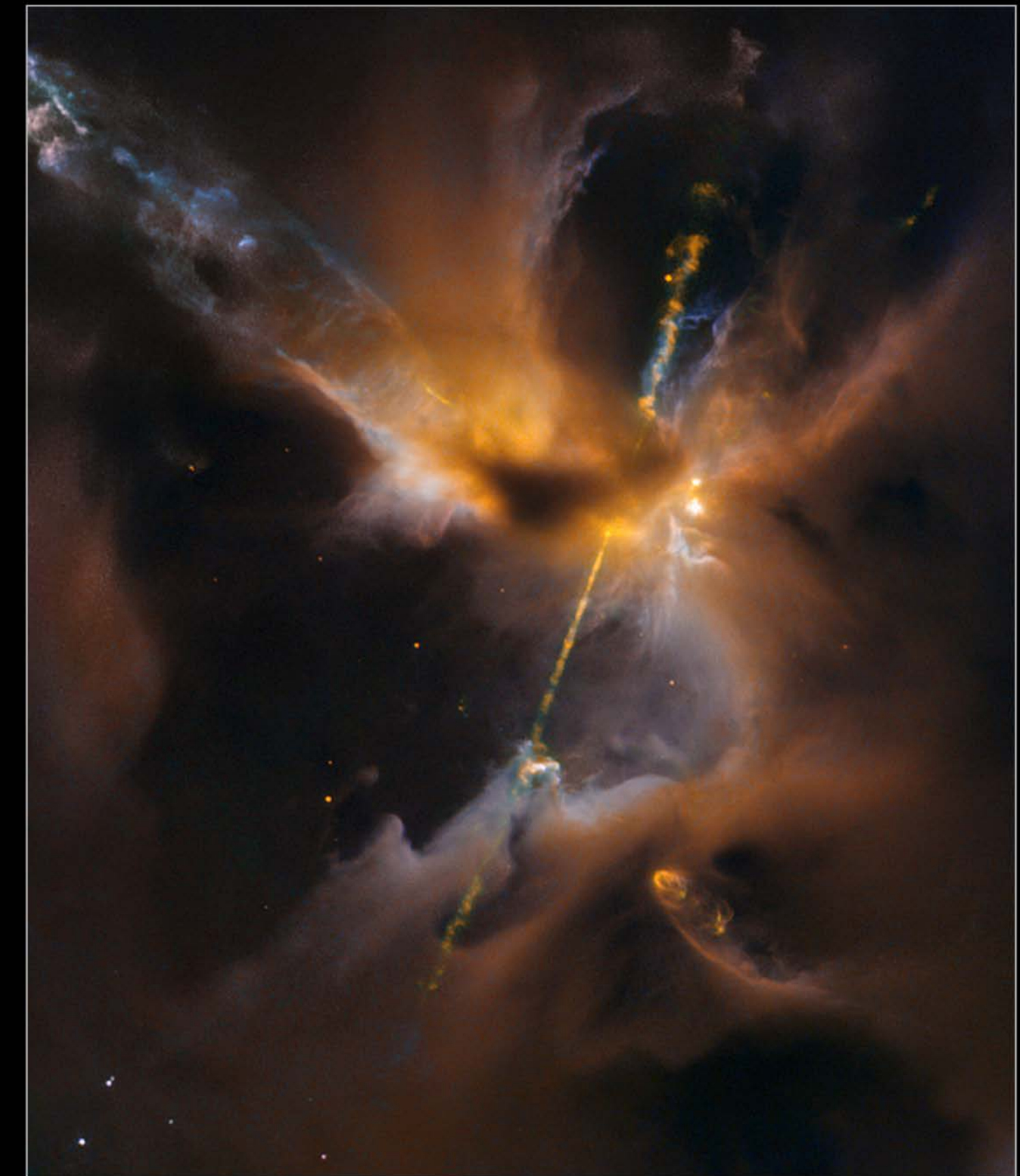
Explore

A 5 minute journey, if you know the way.

Explain



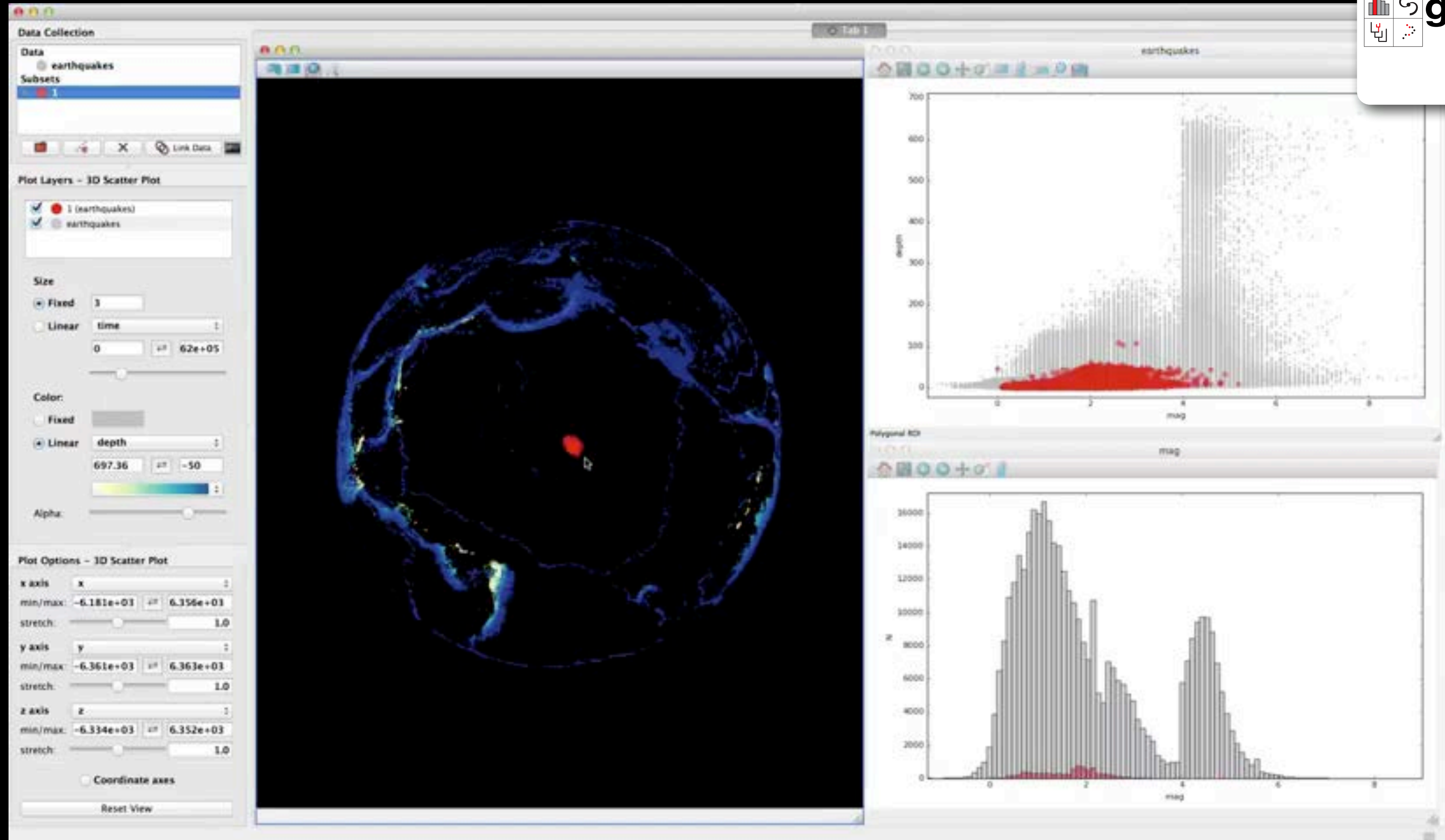
Herbig-Haro Jet HH 24



Hubble
Heritage

NASA and ESA • Hubble Space Telescope • WFC3/IR • WFC2 • STScI-PRC15-42a

“Linked Views of High-dimensional Data” (in Python)



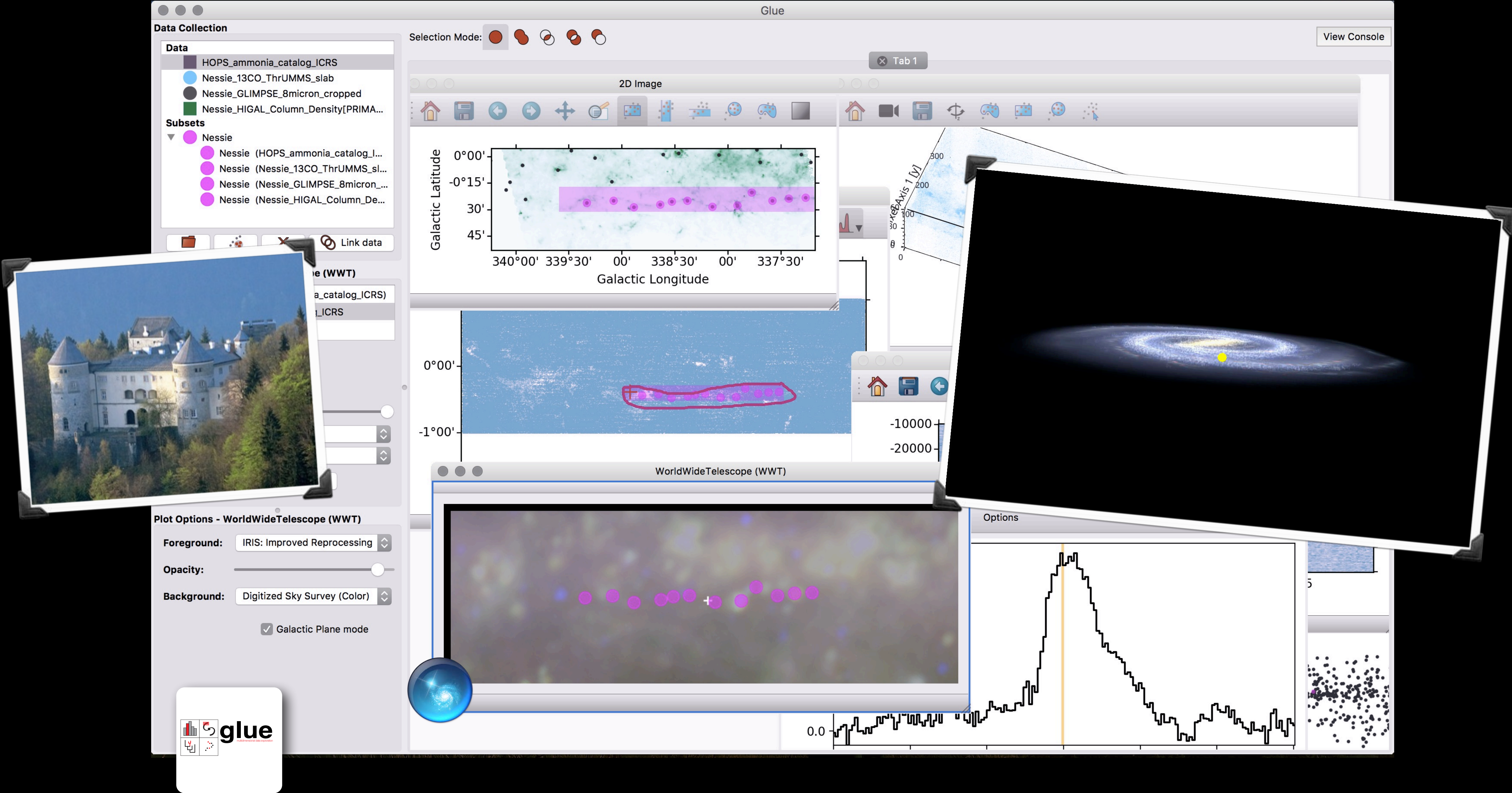
video by Tom Robitaille, lead glue developer glue created by: C. Beaumont, M. Borkin, M. Breddels, P. Qian, T. Robitaille, and A. Goodman, PI

Explore



Explain

glueing together the Milky Way



Citizen Science

Data Collection

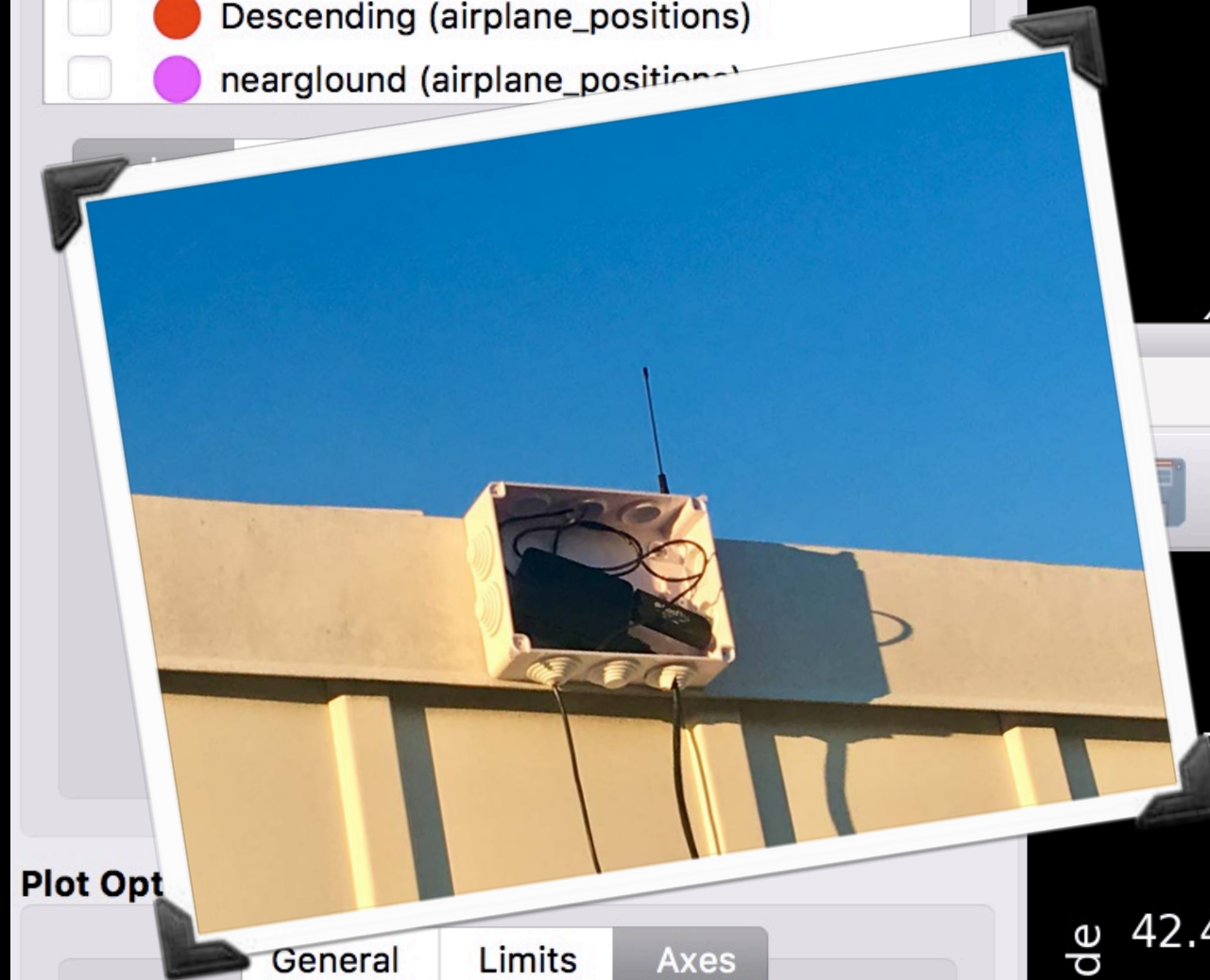
satelliteimages

Subsets

- fastplanes
- nearground
- Descending
- Climbing
- Landing
- A Day in the Life of Logan

Plot Layers - 2D Scatter

- ☐ A Day in the Life of Logan (airplane_positions)
- ☒ Landing (airplane_positions)
- ☐ Climbing (airplane_positions)
- ☐ Descending (airplane_positions)
- ☐ nearground (airplane_positions)



Plot Options

General Limits Axes

x label: Heading (degrees)

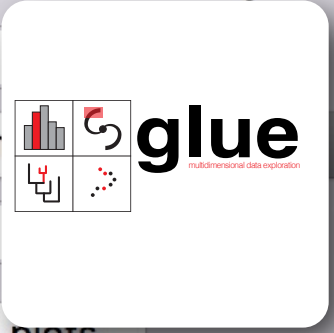
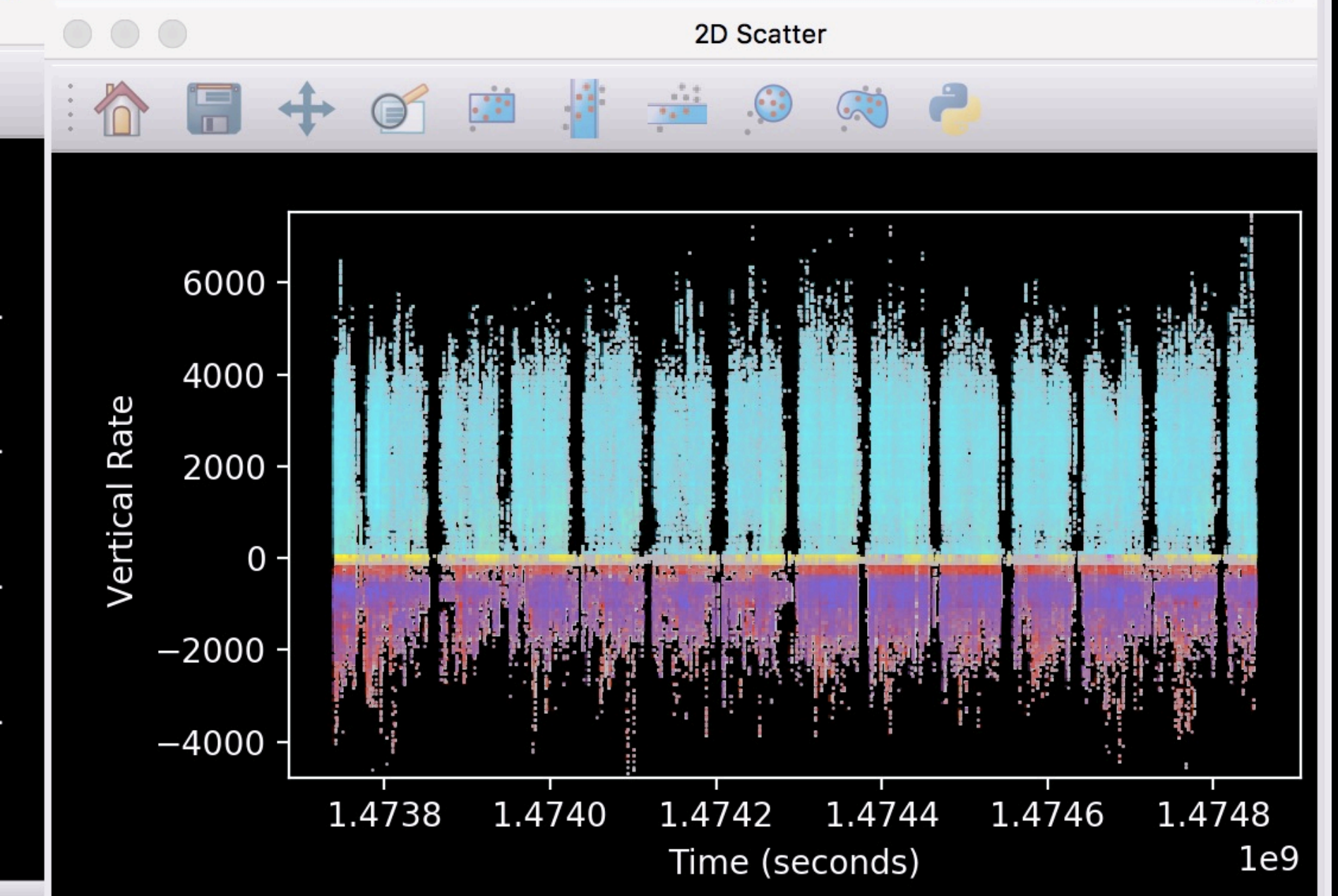
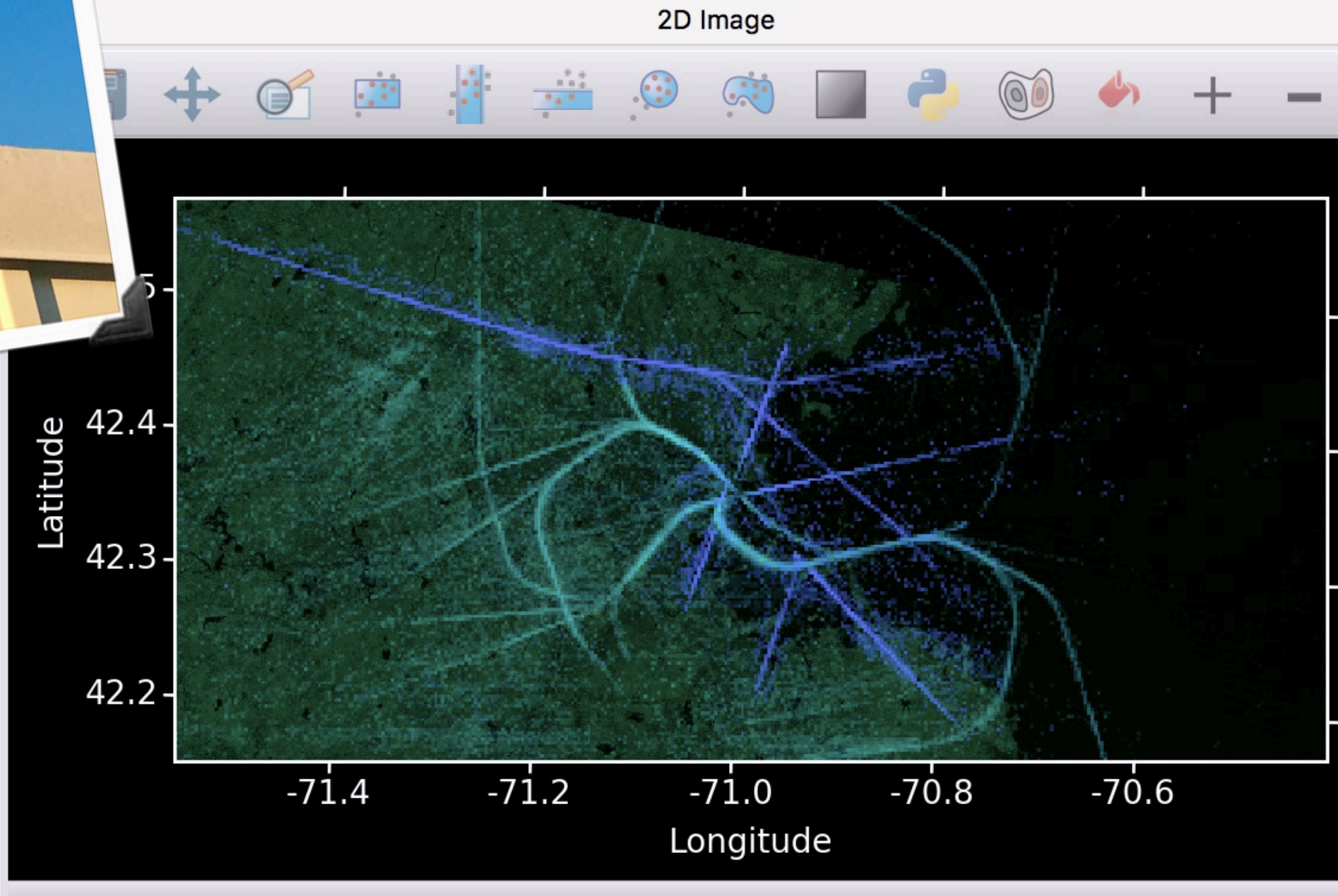
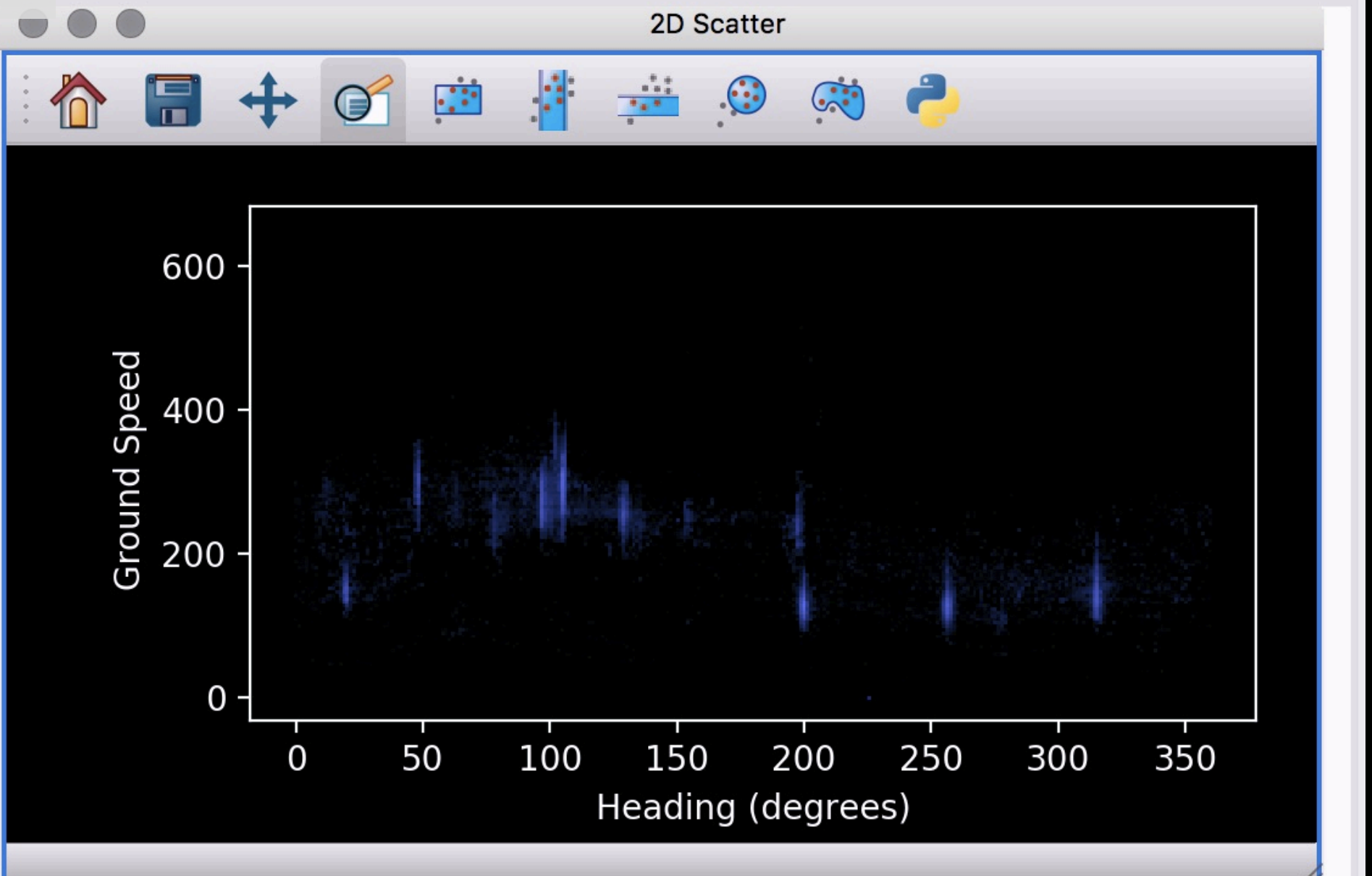
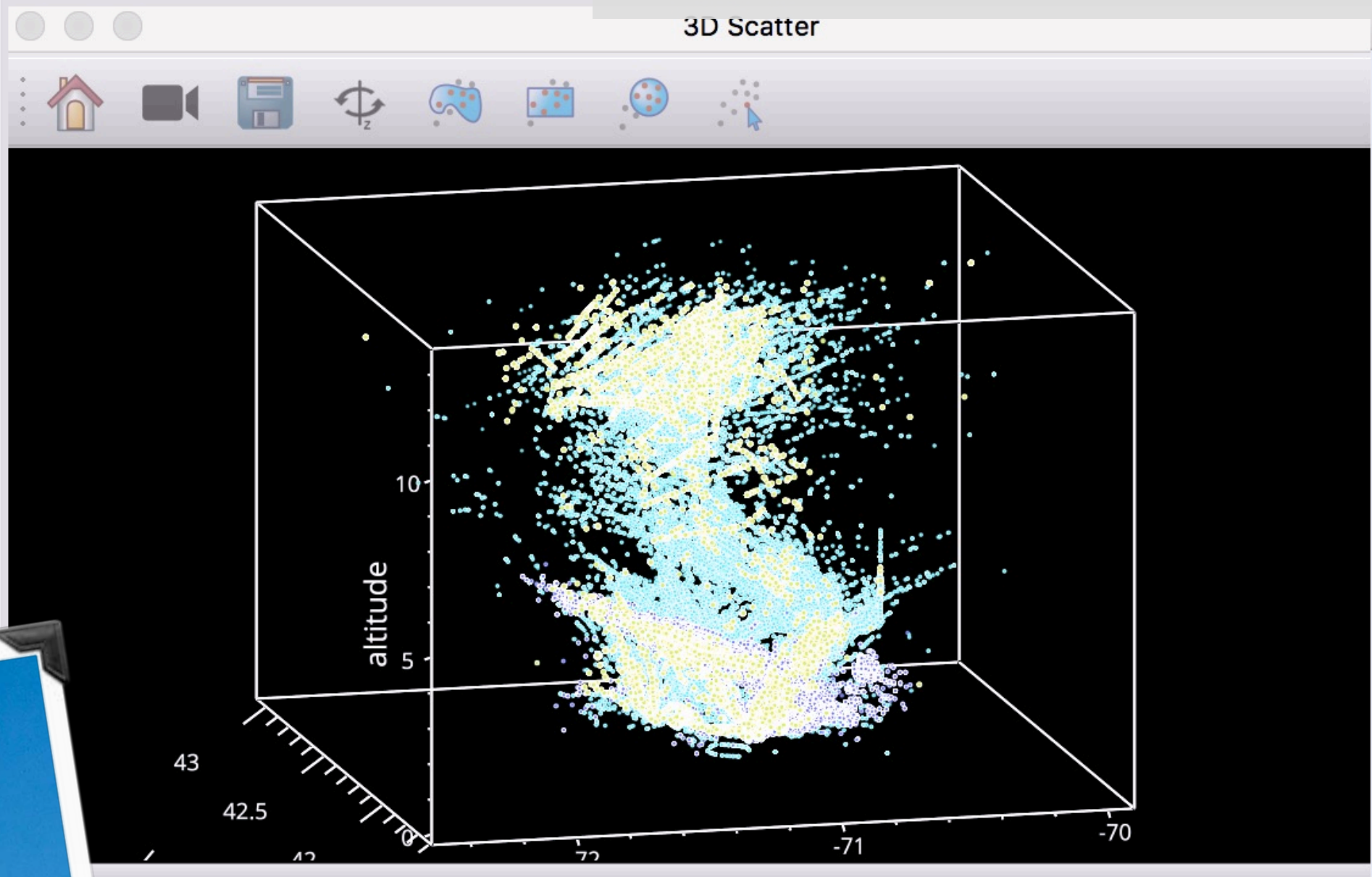
y label: Ground Speed

axis label size: 10

axis label weight: medium

tick label size: 10

Apply to all plots



New Ideas
Discoveries

Public Outreach
Scholarly Publishing

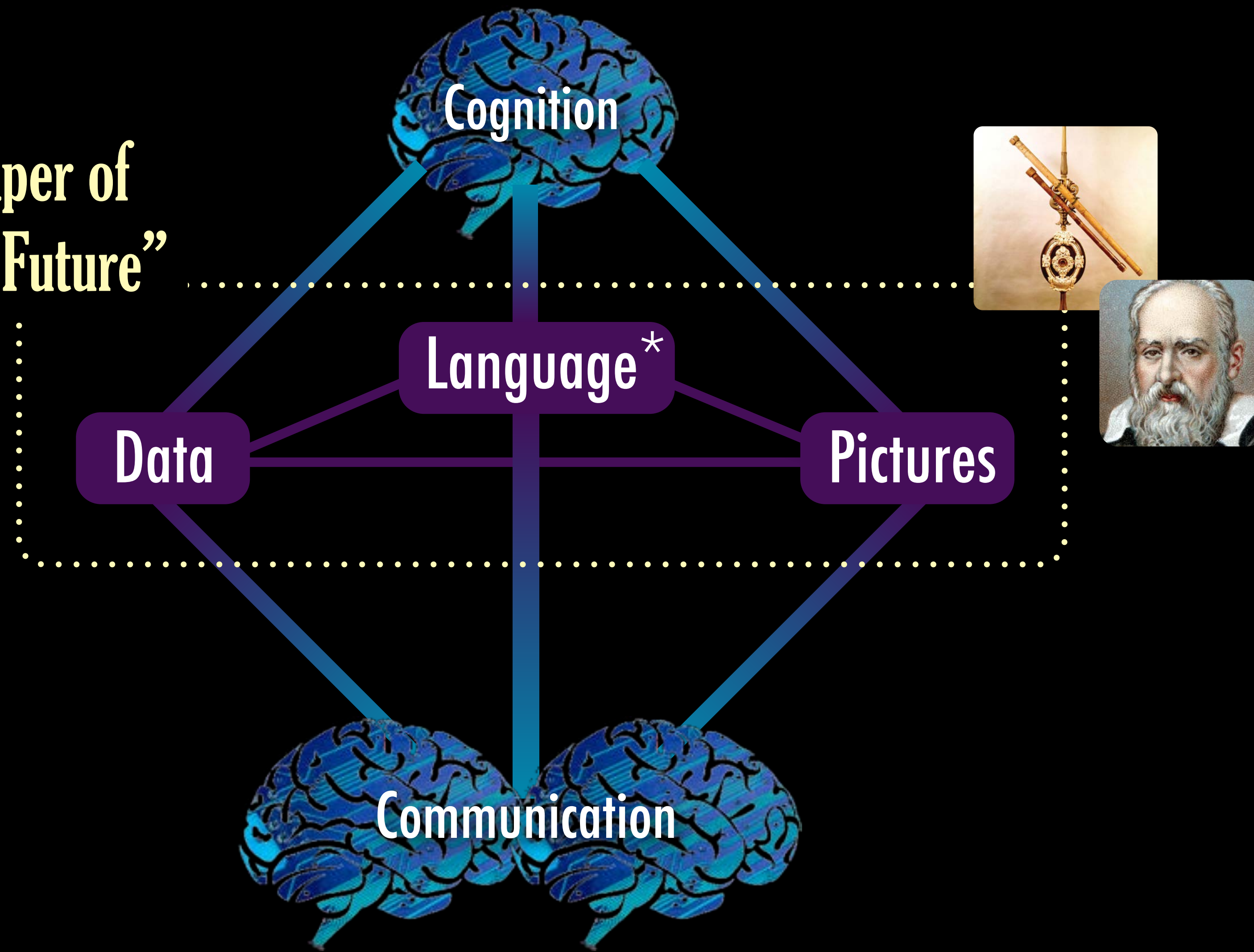
Explore

Explain



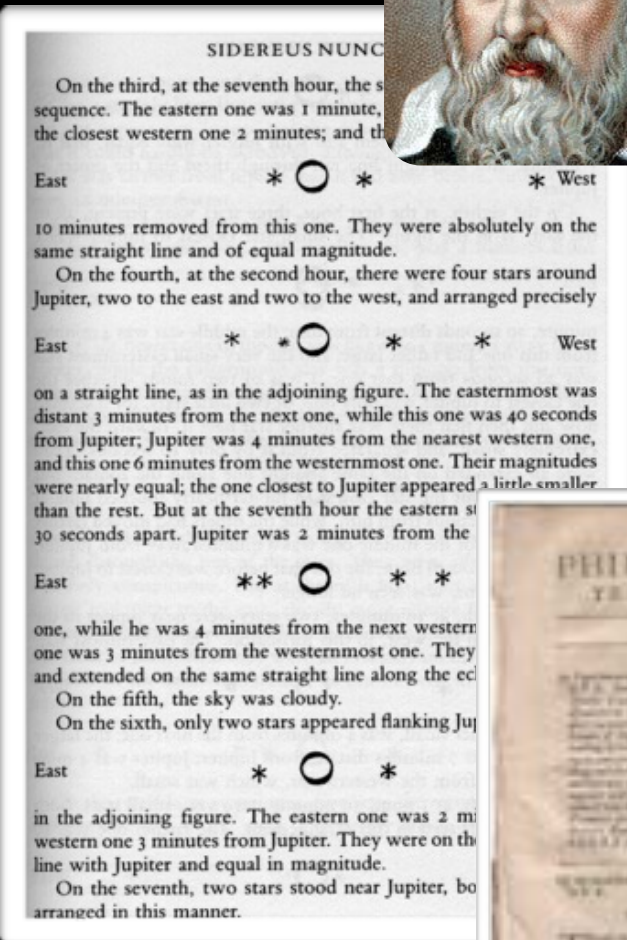
“It’s much harder to go the other way.”

“Paper of
the Future”



*"Language" includes words & math

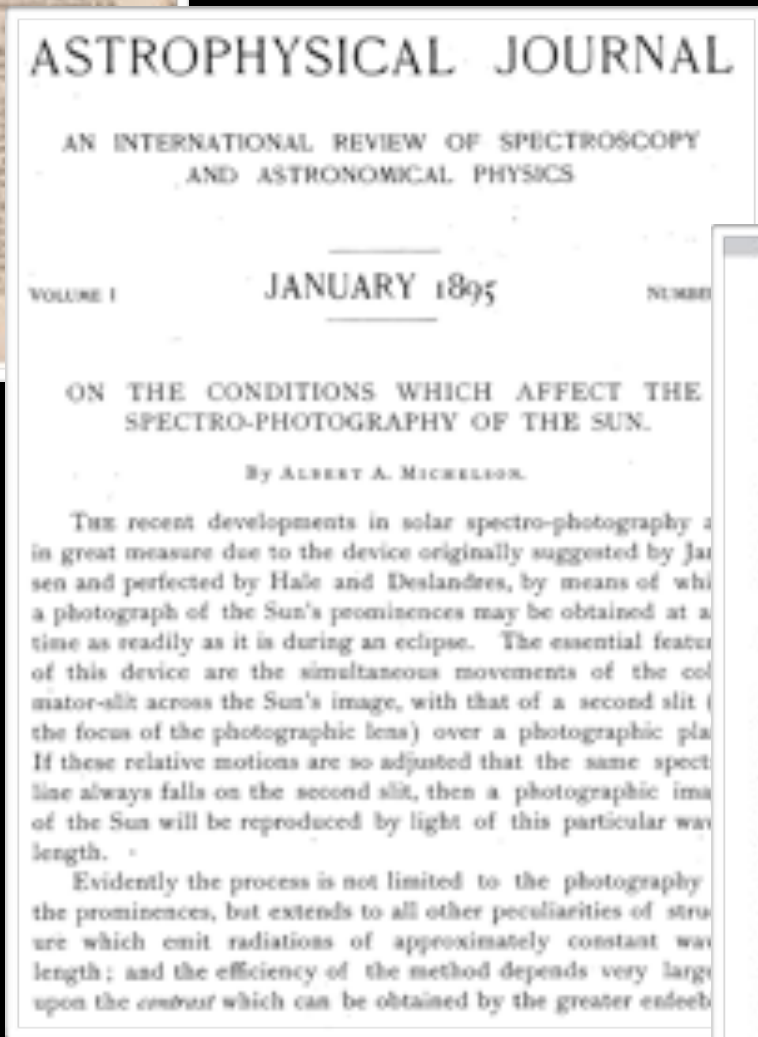
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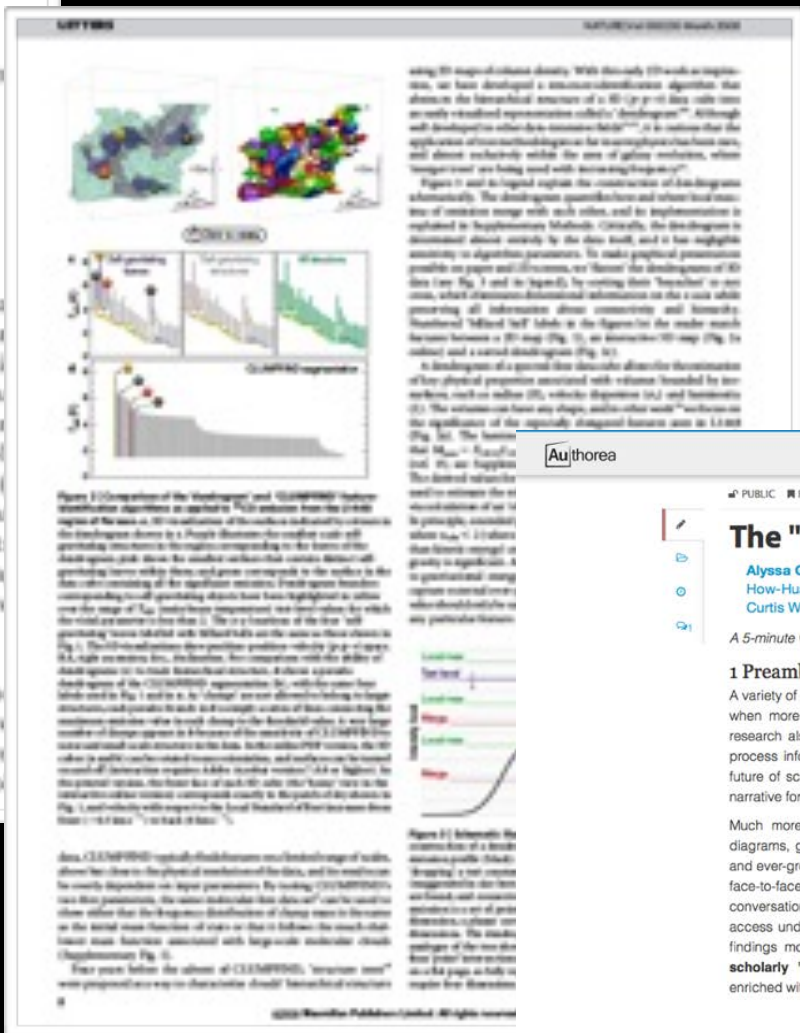
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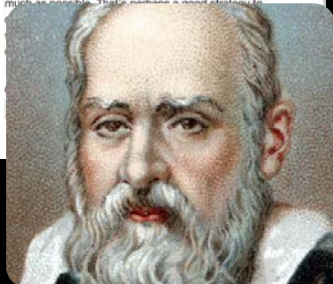
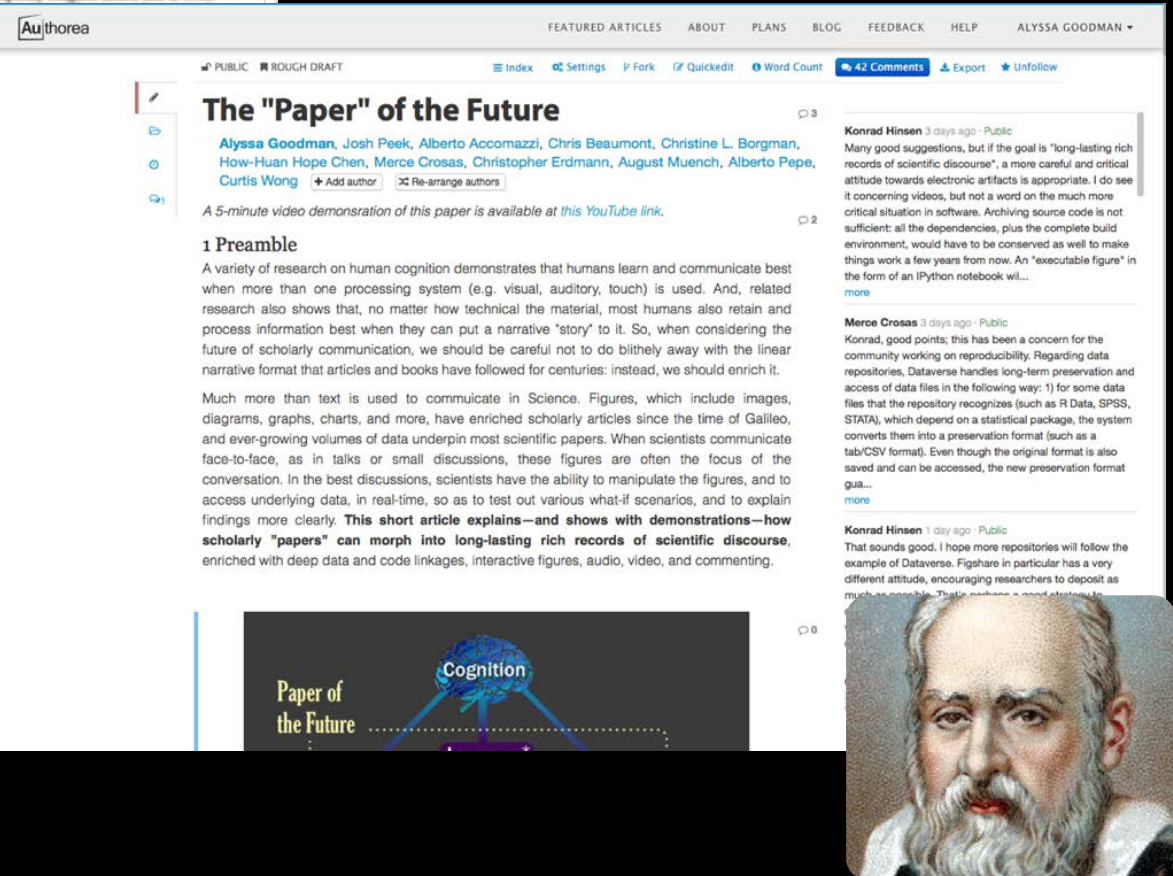
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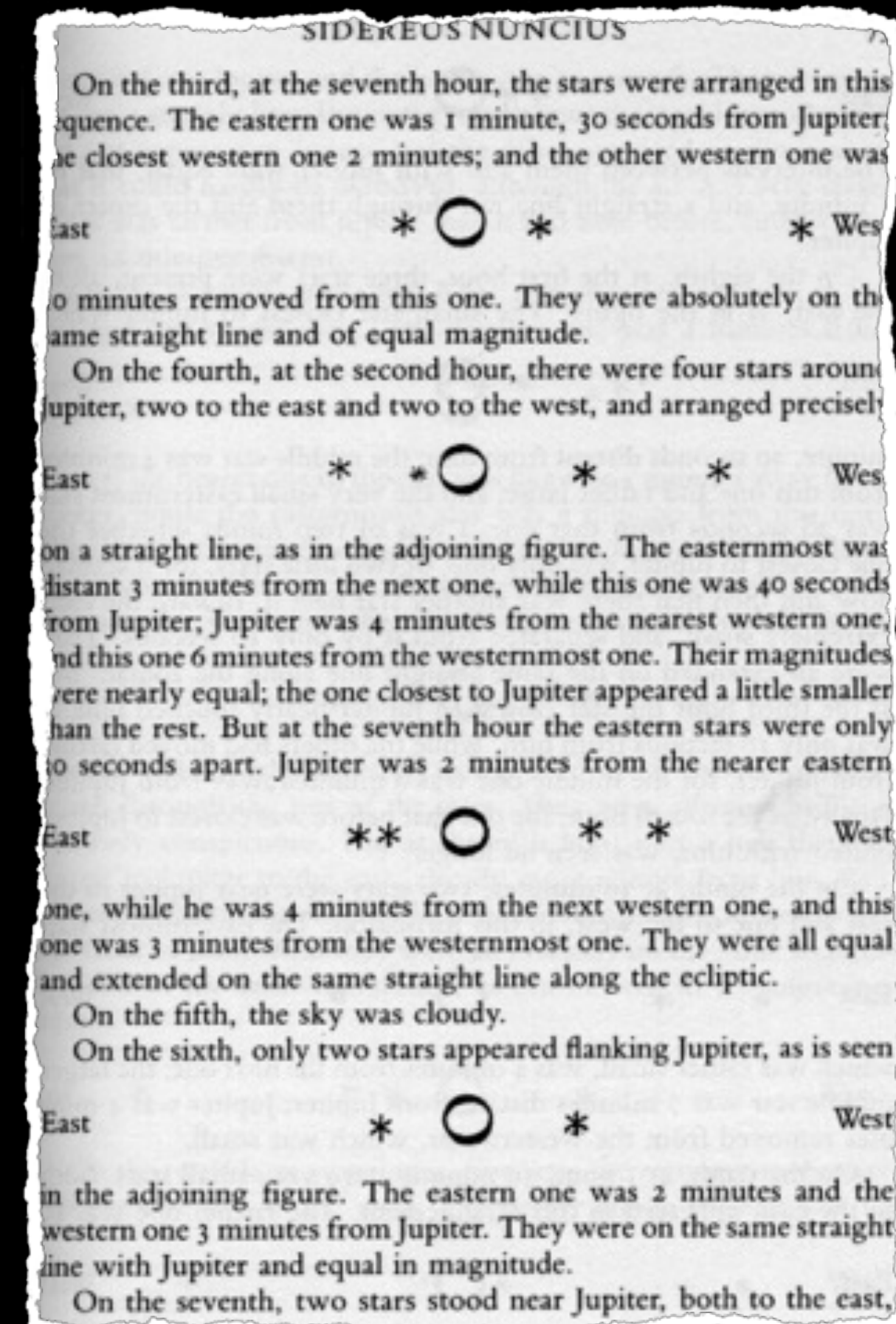
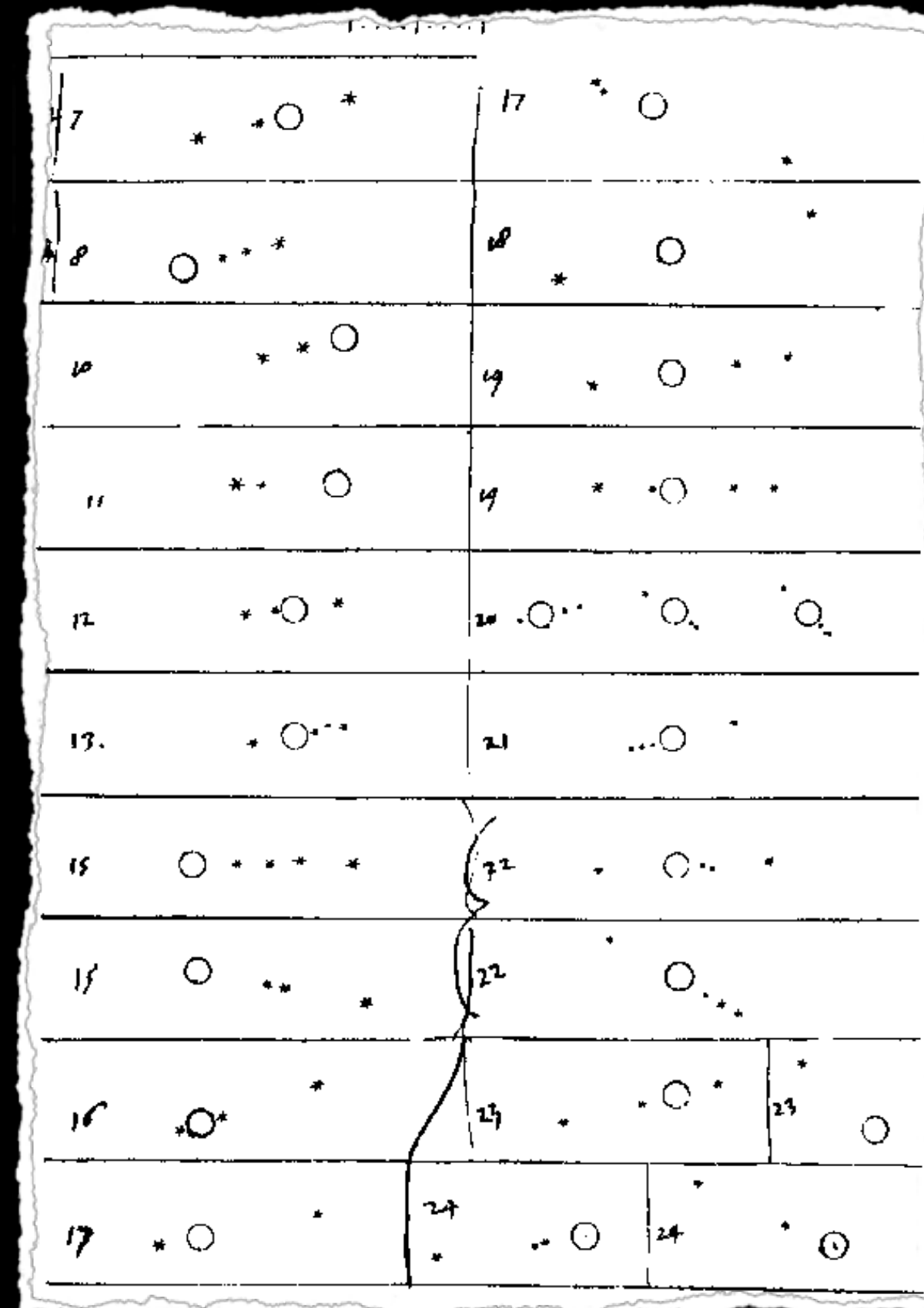
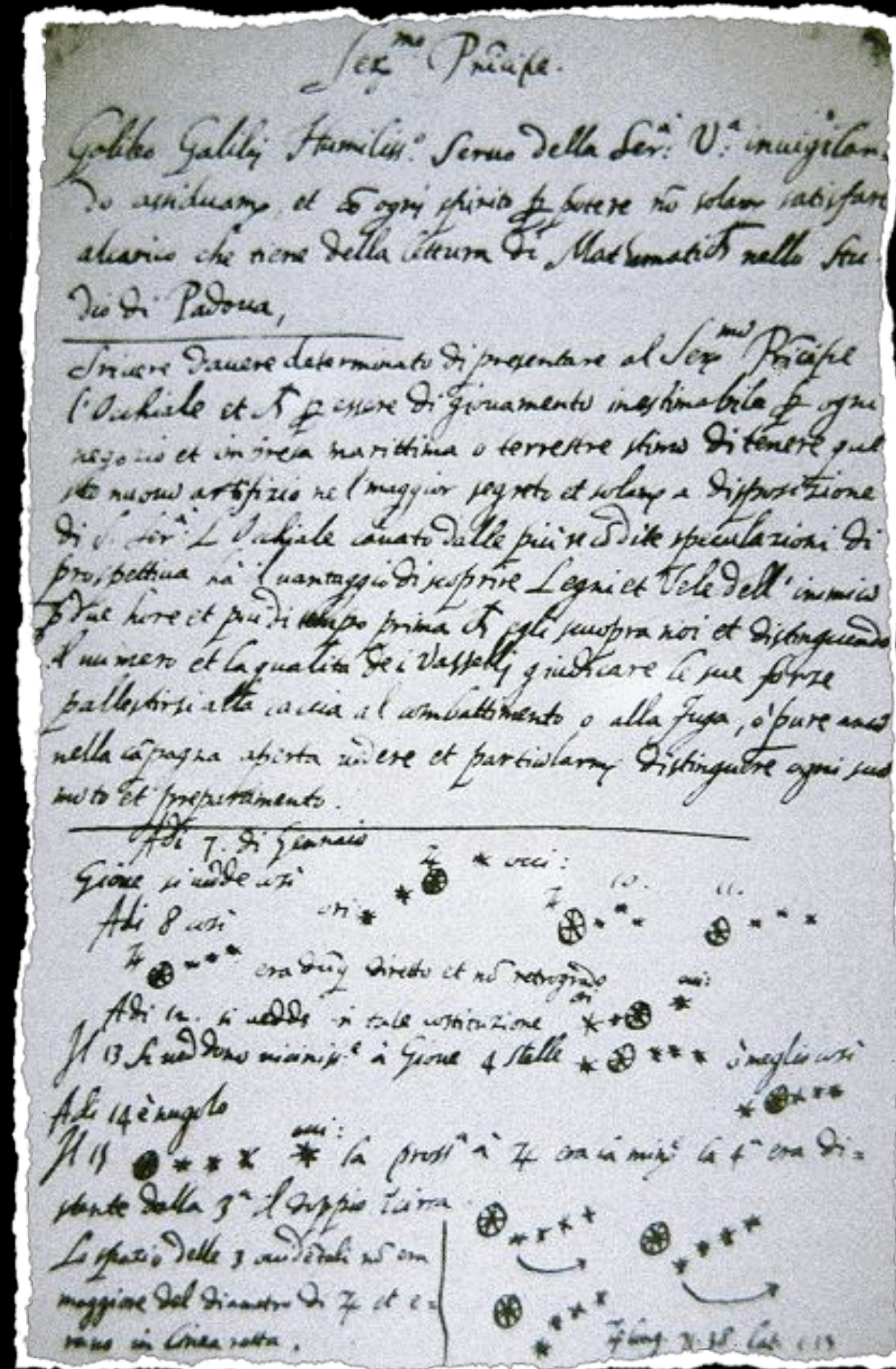
2009



2015



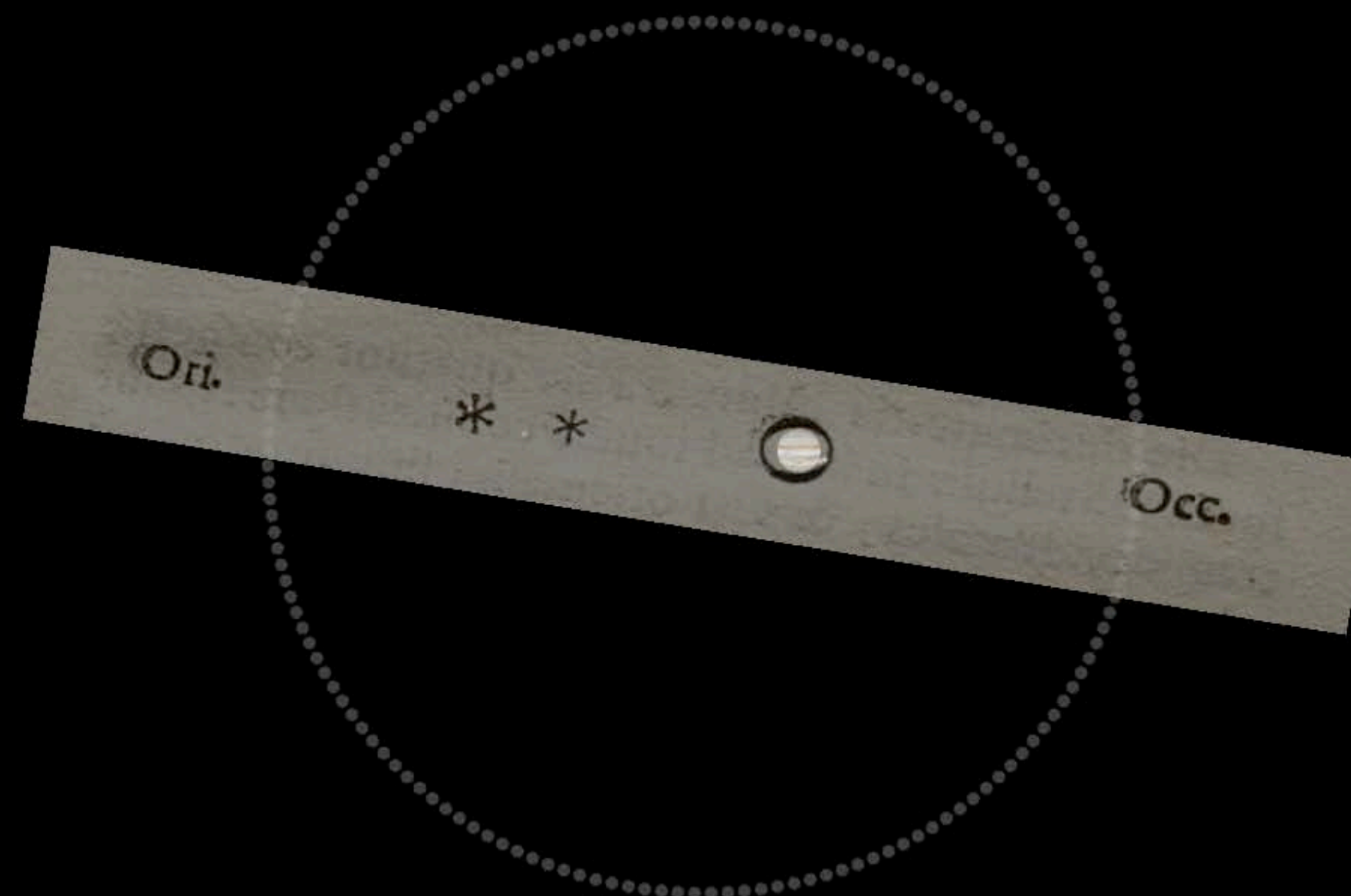
Galileo Discovers Jupiter's Moons, 1610





January 11, 1610

1610



2009

3D PDF interactivity in a “Paper”

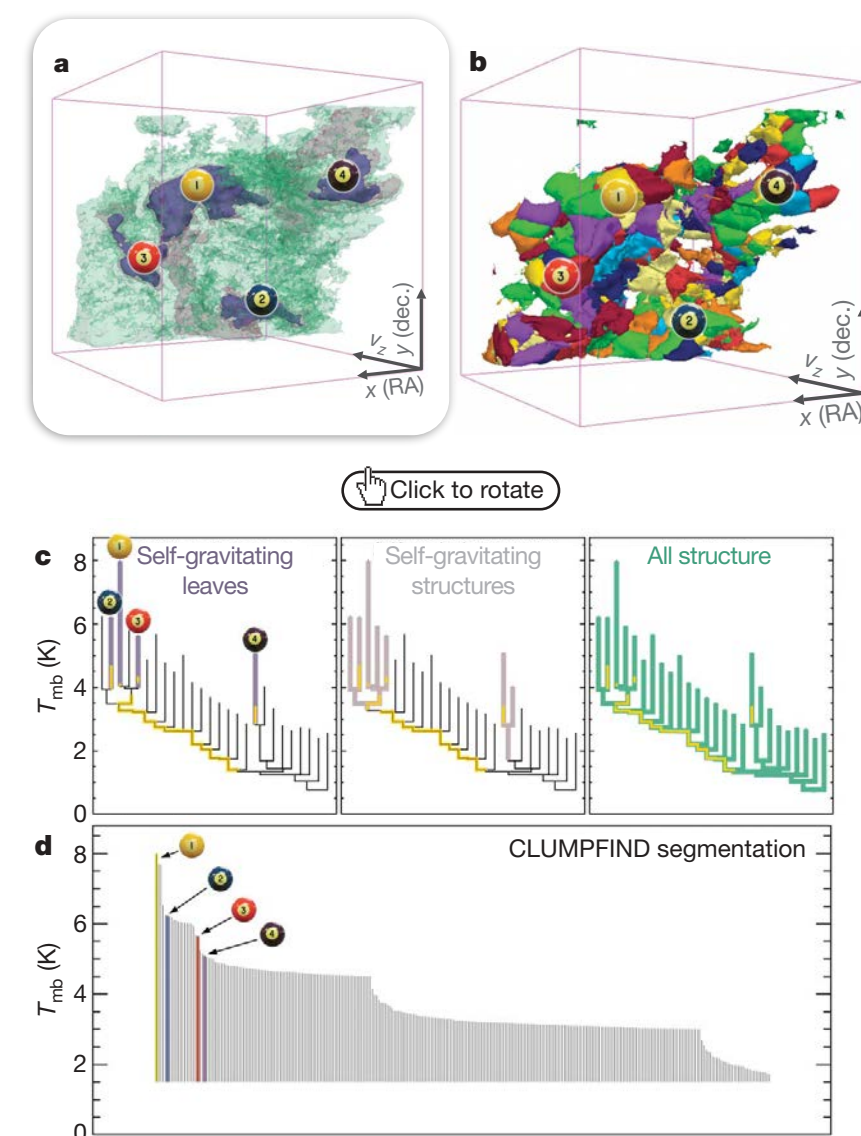


Figure 2 | Comparison of the ‘dendrogram’ and ‘CLUMPFIND’ feature-identification algorithms as applied to ^{13}CO emission from the L1448 region of Perseus. **a**, 3D visualization of the surfaces indicated by colours in the dendrogram shown in **c**. Purple illustrates the smallest scale self-gravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct self-gravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of T_{mb} (main-beam temperature) test-level values for which the virial parameter is less than 2. The x – y locations of the four ‘self-gravitating’ leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position–position–velocity (p – p – v) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (**c**) to track hierarchical structure, **d** shows a pseudo-dendrogram of the CLUMPFIND segmentation (**b**), with the same four labels used in Fig. 1 and in **a**. As ‘clumps’ are not allowed to belong to larger structures, each pseudo-branch in **d** is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the ‘home’ view in the interactive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front (-0.5 km s^{-1}) to back (8 km s^{-1}).

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND’s two free parameters, the same molecular-line data set⁸ can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, ‘structure trees’⁹ were proposed as a way to characterize clouds’ hierarchical structure

using 2D maps of column density. With the early 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a data cube into an easily visualized representation called a dendrogram. The well-developed in other data-intensive applications of tree methodologies so far have been applied almost exclusively within the astronomical community. ‘merger trees’ are being used with increasing frequency.

Figure 3 and its legend explain the dendrogram process schematically. The dendrogram quantifies the hierarchy of emission merge with each other, as explained in Supplementary Methods. The dendrogram is determined almost entirely by the data, and its structure is insensitive to algorithm parameters. The dendrogram is possible on paper and 2D screen data (see Fig. 3 and its legend). The dendrogram is a cross, which eliminates dimensions. The dendrogram is preserving all information. The dendrogram is a cross, which eliminates dimensions. The dendrogram is preserving all information. The dendrogram is a cross, which eliminates dimensions. The dendrogram is preserving all information.

A dendrogram of a spectral cube can be constructed of key physical properties of the data, such as radius (R), luminosity (L), and virial parameter (α). The volumes can have any shape, and the dendrogram can be used to track hierarchical structure. The significance of the especially elongated features (Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}} L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^2 \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an ‘observed’ virial parameter, $\alpha_{\text{obs}} = 5\sigma_v^2 R / GM_{\text{lum}}$. In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{\text{obs}} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p – p – v space where self-gravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields¹⁶, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

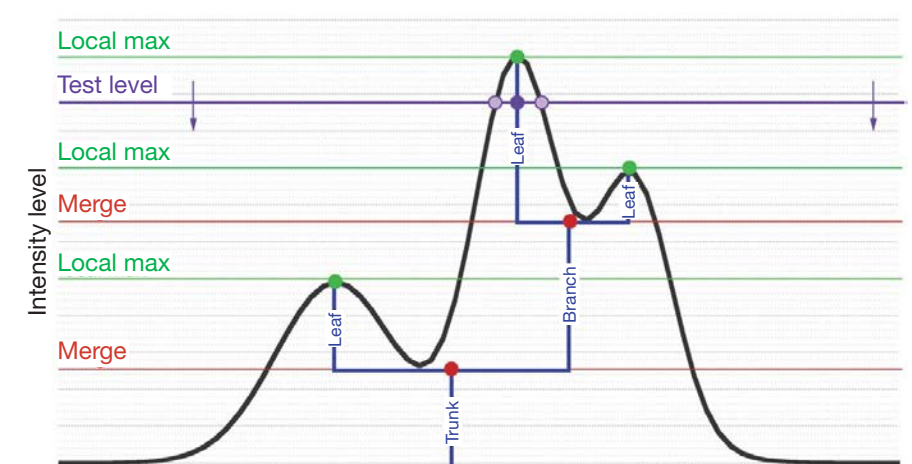
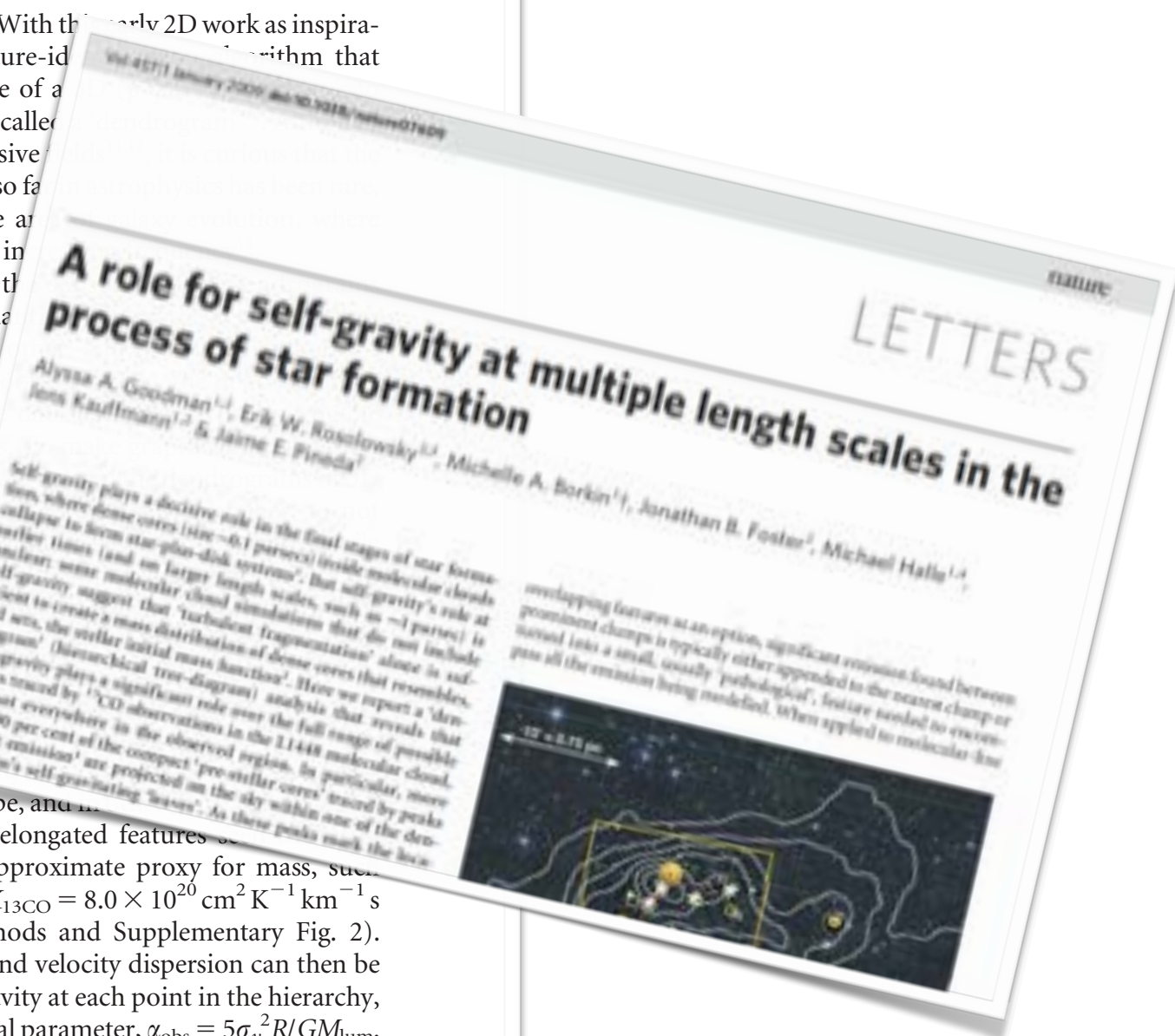


Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional emission profile (black). The dendrogram (blue) can be constructed by ‘dropping’ a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimension, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from ‘isosurface’ rather than ‘point’ intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.



Goodman et al. 2009, Nature,
cf: Fluke et al. 2009

Create

1 / 4

131%

Tools

Comment

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Vol 457 | 1 January 2009 | doi:10.1038/nature07609

nature

LETTERS

A role for self-gravity at multiple length scales in the process of star formation

Alyssa A. Goodman^{1,2}, Erik W. Rosolowsky^{2,3}, Michelle A. Borkin^{1†}, Jonathan B. Foster², Michael Halle^{1,4}, Jens Kauffmann^{1,2} & Jaime E. Pineda²

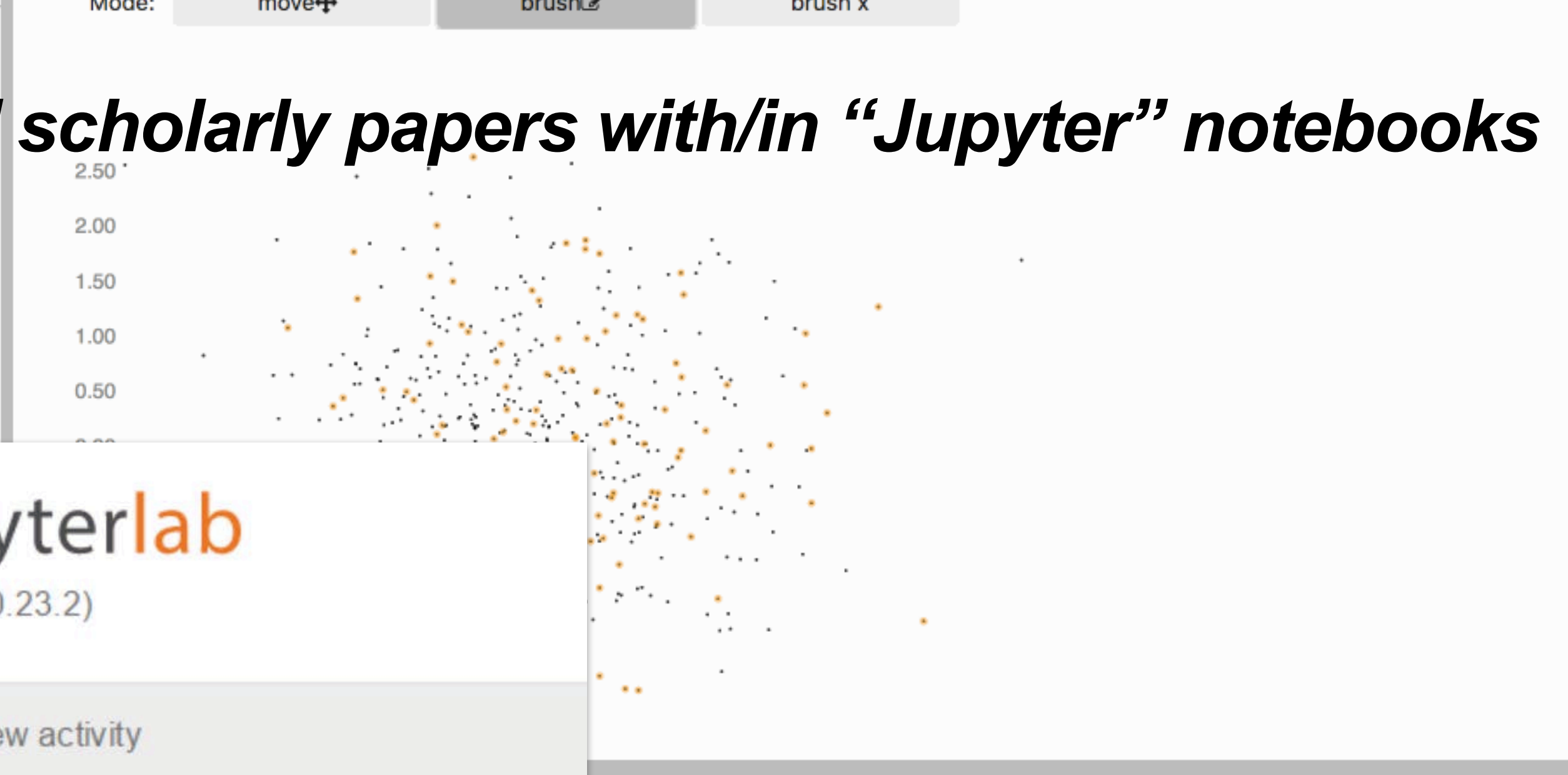
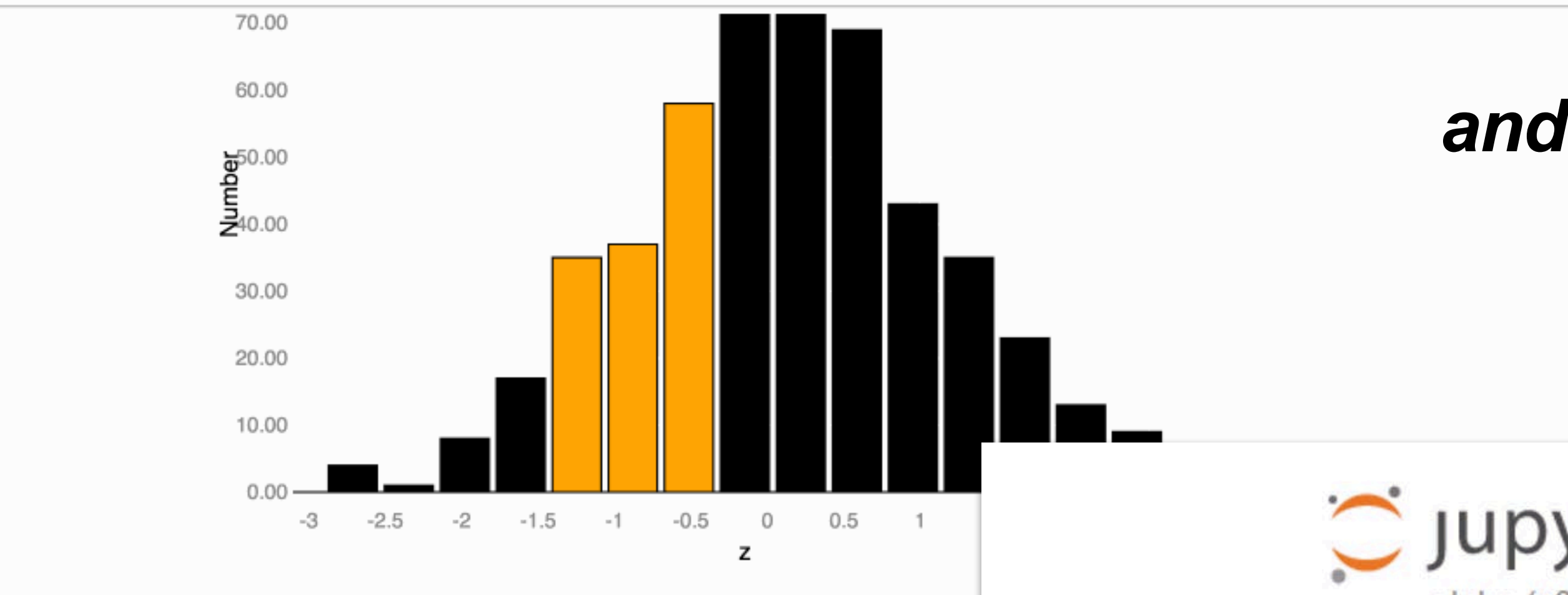
Self-gravity plays a decisive role in the final stages of star formation, where dense cores (size ~ 0.1 parsecs) inside molecular clouds collapse to form star-plus-disk systems¹. But self-gravity's role at earlier times (and on larger length scales, such as ~ 1 parsec) is unclear; some molecular cloud simulations that do not include self-gravity suggest that 'turbulent fragmentation' alone is sufficient to create a mass distribution of dense cores that resembles, and sets, the stellar initial mass function². Here we report a 'dendrogram' (hierarchical tree-diagram) analysis that reveals that self-gravity plays a significant role over the full range of possible scales traced by ^{13}CO observations in the L1448 molecular cloud, but not everywhere in the observed region. In particular, more than 90 per cent of the compact 'pre-stellar cores' traced by peaks of dust emission³ are projected on the sky within one of the dendrogram's self-gravitating 'leaves'. As these peaks mark the locations of already-forming stars, or of those probably about to form, a self-gravitating cocoon seems a critical condition for their exist-

overlapping features as an option, significant emission found between prominent clumps is typically either appended to the nearest clump or turned into a small, usually 'pathological', feature needed to encompass all the emission being modelled. When applied to molecular-line

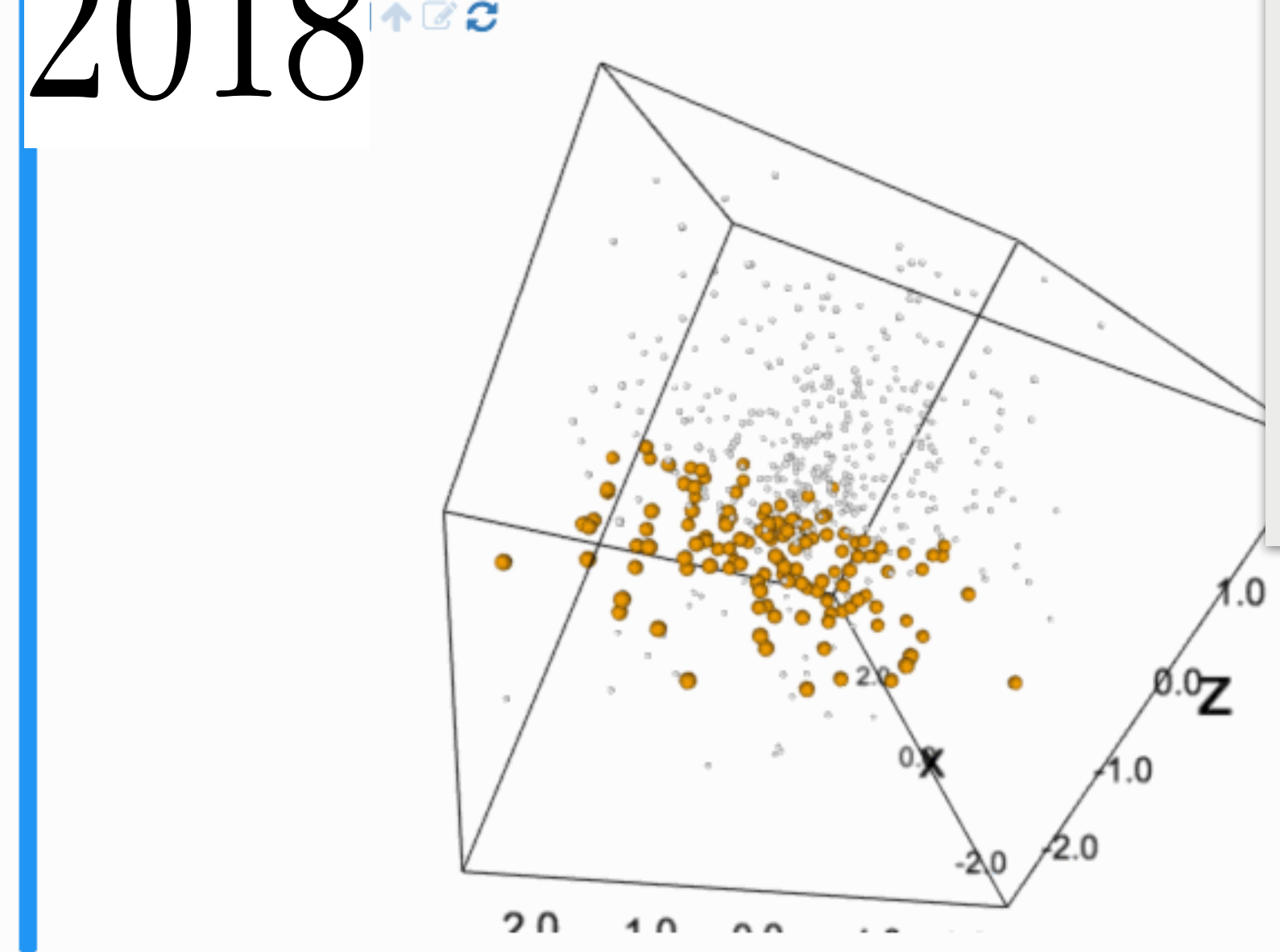
The image shows a dark-field astronomical photograph of the L1448 molecular cloud. Overlaid on the image are white contour lines representing the intensity of ^{13}CO emission. A yellow rectangular box highlights a specific region of interest. Within and around this box, several small, bright spots are marked with yellow circles and numbers (1, 2, 3, 4), representing pre-stellar cores. A scale bar at the top left of the image indicates a length of $10' \approx 0.75 \text{ pc}$.


and scholarly papers with/in “Jupyter” notebooks

2018




```
In [4]: app.scatter3d('x', 'y', 'z');
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alpha (v0.23.2)


Start a new activity



Notebook

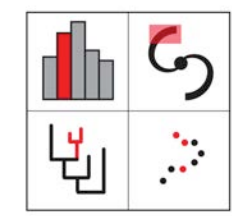


Code Console

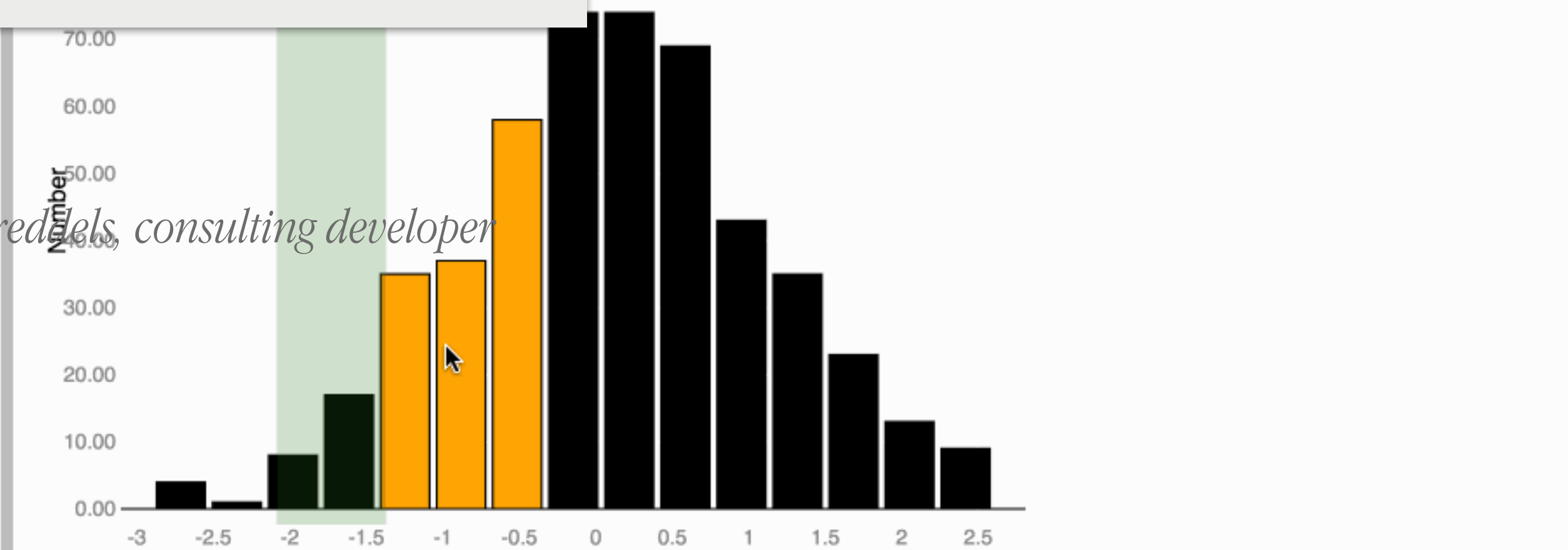


Text Editor

Video courtesy of Maarten Breddele, consulting developer



glue *in the browser*
multidimensional data exploration

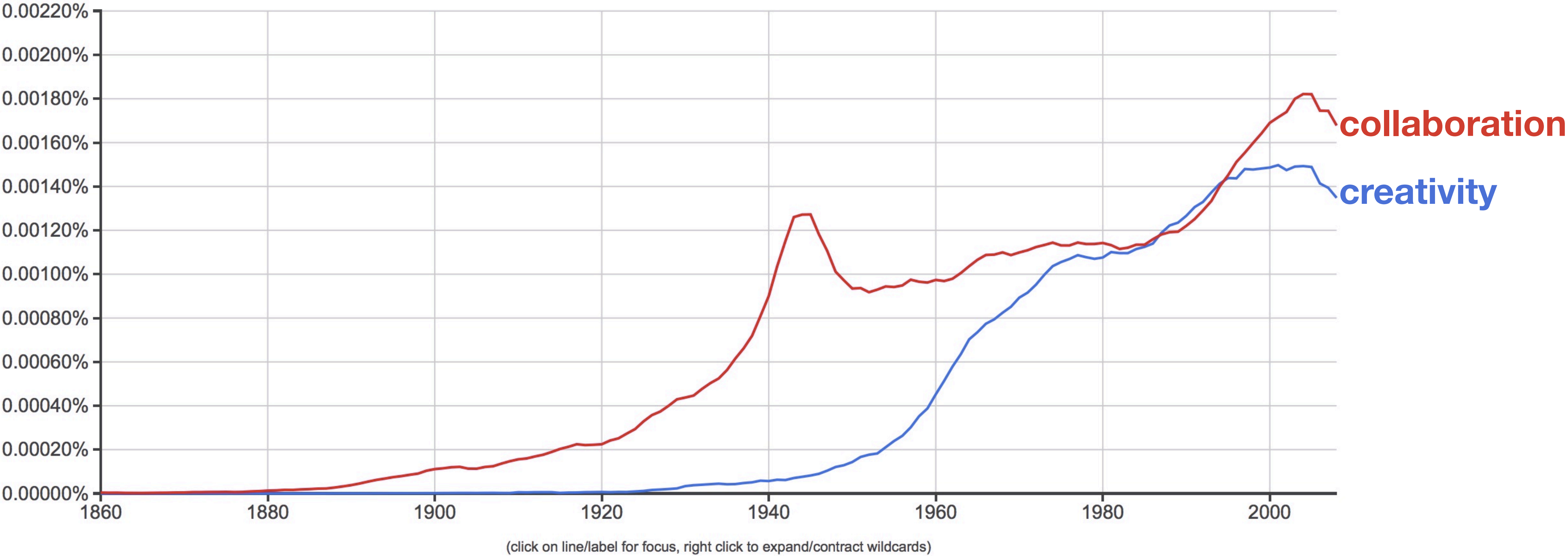


Google Books Ngram Viewer

Graph these comma-separated phrases: ☒ case-insensitive

between and from the corpus with smoothing of .

[Embed Chart](#)



ASTRONOMY REWIND

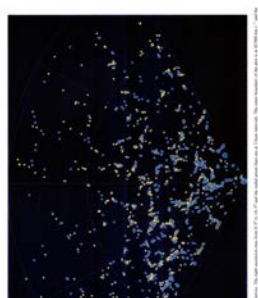
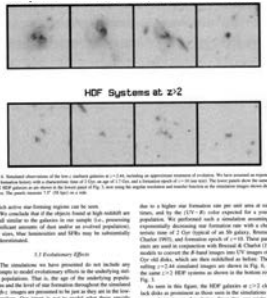
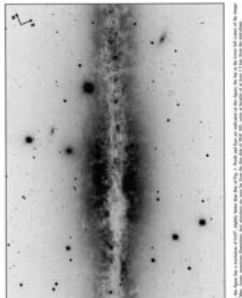
ABOUTCLASSIFYTALKCOLLECTFEEDBACK

agoodman

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"Your contributions unlock the information from old astronomy journals. Thank you and enjoy the images!"

ABOUT ASTRONOMY REWIND

This project is part of an ongoing NASA-funded effort aimed at turning the SAO/NASA Astrophysics Data System (ADS) into a data resource. The result will be a database of astro-referenced images, i.e., images of the sky for which coordinates, orientation, and pixel scale will be publicly available through NASA data archives, [the Astronomy Image Explorer](#), and [World Wide Telescope](#), thanks to your help!






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Plot Layers - 1D Histogram

Plot Options - 1D Histogram

ROOTNAME

Declination

Right Ascension

Number

Log SCI

Custom Parts Organizer Box Included!

11'00"

34'32'30'


ROOTNAME	A_0_2	A_0
0 lc9a04cvq	SCI	2
1 lc9a04cvq	SCI	1
2 lc9a04cvq	SCI	2
3 lc9a04cvq	SCI	1
4 lc9a04cvq	SCI	2
5 lc9a04cvq	SCI	1
6 lc9a04cvq	SCI	2
7 lc9a04cvq	SCI	1




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Data Collection

Data

Subsets

Plot Layers - 3D Volume Rendering

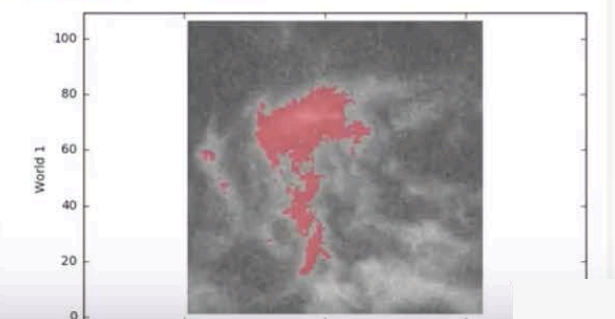
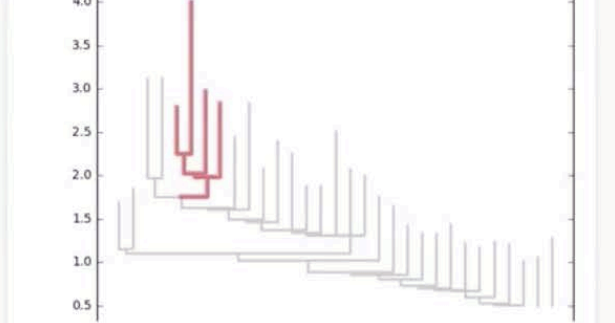
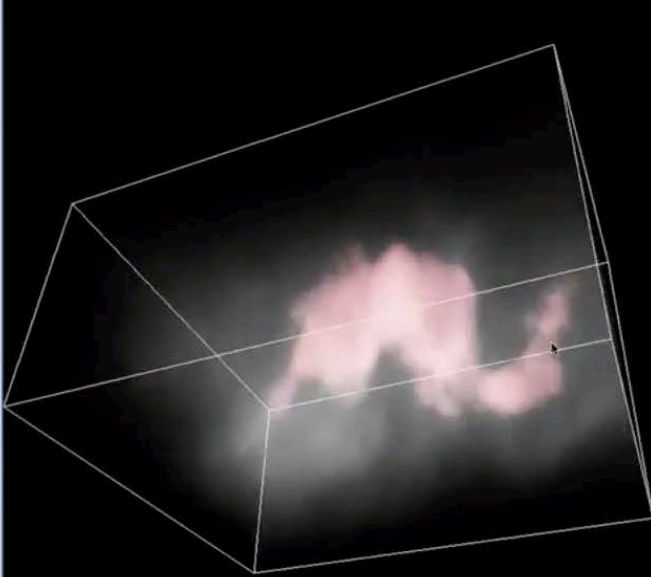
Plot Options - 3D Volume Rendering

Intensity

Color

Alpha

World 1



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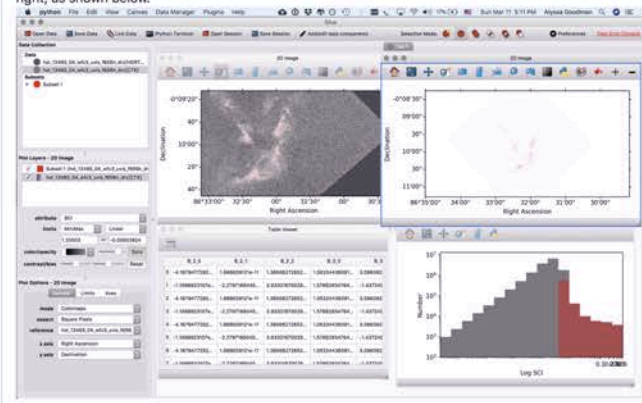
Log axis labels in 1D histograms #1597

aagoodman opened this issue a minute ago · 0 comments

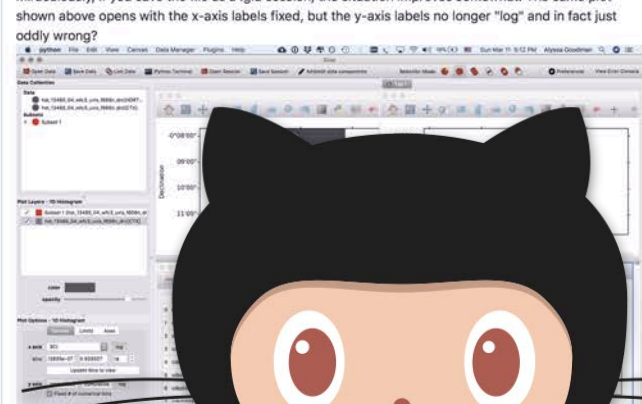
aagoodman commented a minute ago

The .glu file that can reproduce the behavior below is at <https://www.dropbox.com/sh/Op1lw26npax5Gwq/AAC2J4Va98Vp75o6Dtkn6hma7d?dl=0>.

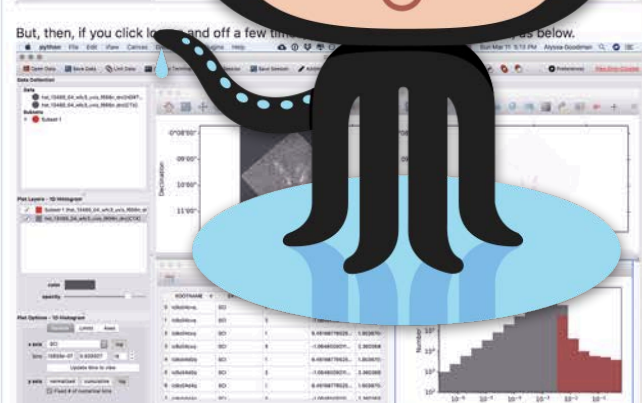
What happens? When you make a histogram and ask for (both, or one) log axes, the tick labeling is not right, as shown below.



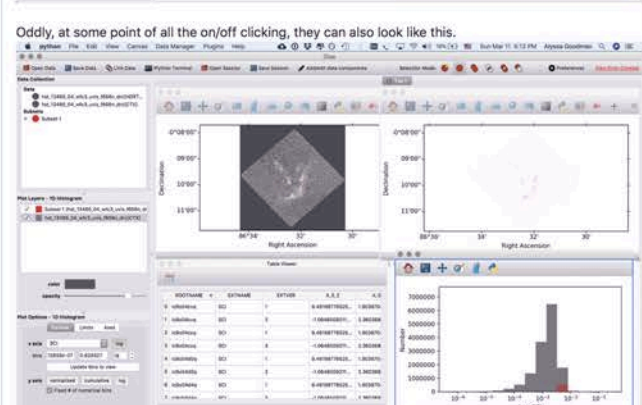
Miraculously, if you save the file as a .glu session, the situation improves somewhat. The same plot shown above opens with the x-axis labels fixed, but the y-axis labels no longer "log" and in fact just oddly wrong?



But, then, if you click to expand off a few bins, the situation improves below.



Oddly, at some point of all the on/off clicking, they can also look like this.



WritePreview

AA- B f

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The Road from **Exploration** to **Explanation**, and Back



HARVARD
UNIVERSITY

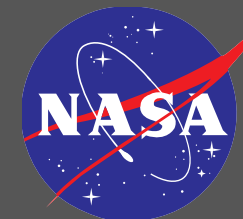
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AMERICAN ASTRONOMICAL SOCIETY
Enhancing and sharing humanity's scientific understanding of the universe since 1899.

Alyssa A. Goodman

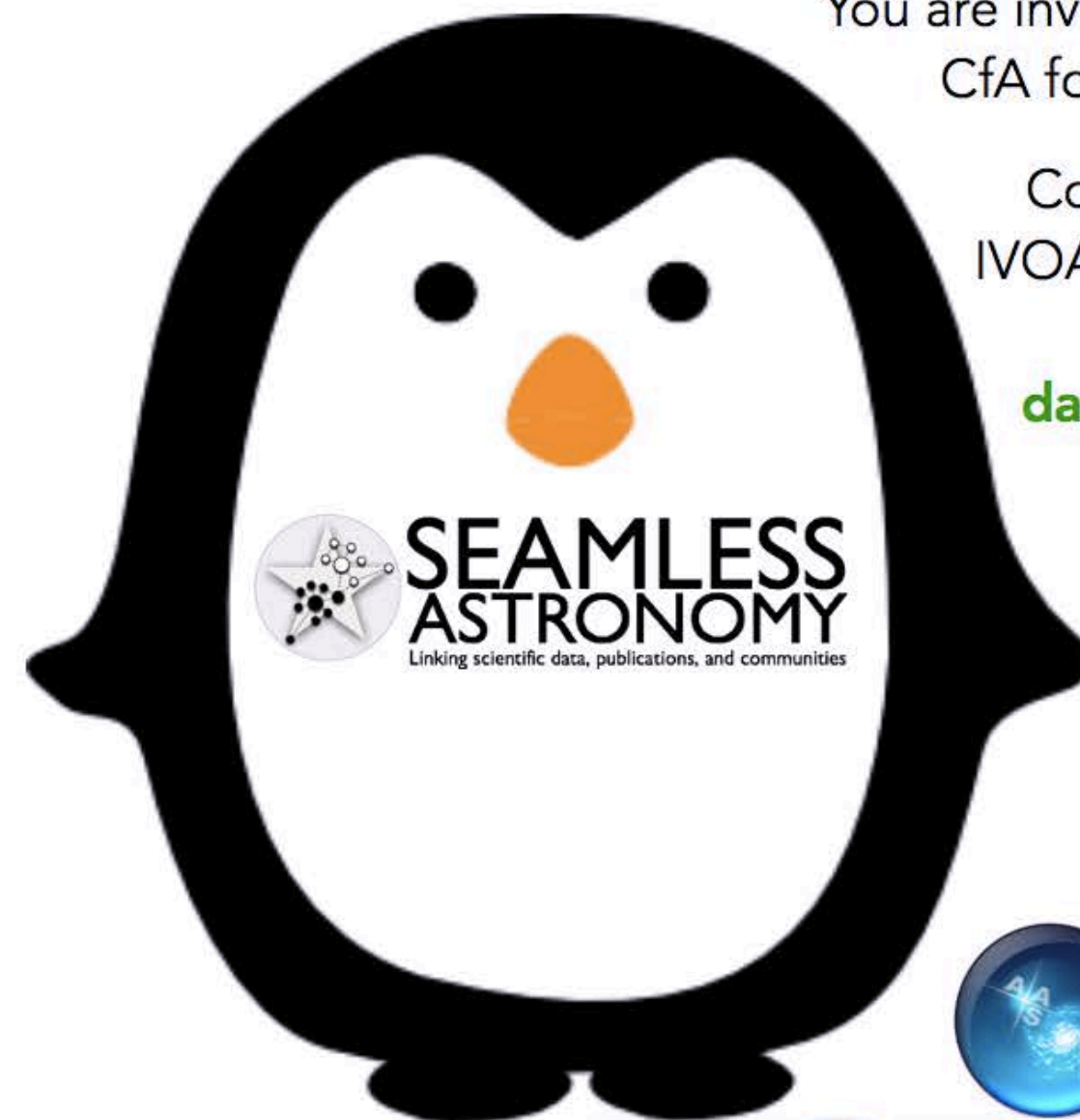
Harvard-Smithsonian Center for Astrophysics + Radcliffe Institute for Advanced Study
@AlyssaAGoodman

THE PARTY PENGUIN IS BACK, LUCKY YOU!



DEMO-FEST 2018

[A PARTY!]



You are invited to join the Seamless Astronomy group of the CfA for 1:1 software demos, snacks & green beverages.

Come learn from ADS, glue, WorldWide Telescope, IVOA, Dataverse, Science Education, Library, Chandra and other experts about how to make **research**, **data-sharing**, **visualization**, **teaching** and **outreach** more "Seamless" & more fun.



MARCH 20, 3:00 TIL 4:30 PM, WOLBACH LIBRARY
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