The First Galaxies

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Outline

1. Introduction

- 1.1 Observing our past
- 1.2 Cosmological context: the expanding universe
 - 1.2.1 FRW metric, cosmological parameters
 - 1.2.2 Thermal history of the Universe: inflation, nucleosynthesis, matter-radiation equality, recombination, reionization

1.2.3 Cosmic composition: radiation, baryons, neutrinos, dark matter, dark energy

- 1.3 Observational overview
- 2. From recombination to the first galaxies
 - 2.1 Initial conditions (power spectrum and its parameters)
 - $2.2\ {\rm Growth}$ of linear perturbations
 - 2.2.1 The Zel'dovich approximation
 - 2.3 Thermal history during the Dark Ages (adiabatic cooling, Compton heating)
- 3. Nonlinear structure
 - 3.1 Properties of virialized halos $(M_{\rm vir}(z, V_c), \text{ NFW density profile}).$
 - 3.2 Abundance and clustering of dark-matter halos
 - 3.2.1 Press-Schechter mass function
 - 3.2.2 Extended Press-Schechter, merger trees
 - 3.2.3 Improvements to the Press-Schechter approach
 - 3.3 Nonlinear clustering: The halo model
 - 3.4 Numerical simulations of structure formation (basic principles)
- 4. The Intergalactic Medium
 - 4.1 The Lyman- α forest: overview
 - 4.2 The ionizing background
 - 4.3 Modeling the forest
 - 4.3.1 Numerical simulations
 - 4.3.2 Semi-analytic models
 - 4.4 Metal-line systems
- 5. The First Stars

5.1 Chemistry and cooling of primordial gas (including H_2 , HD).

- 5.2 Formation of the first metal-free stars
 - 5.2.1 Fragmentation mass (Jeans mass, Bonor-Ebert)
 - 5.2.2 Early accretion phase and mass saturation (simulations and analytic models)
 - 5.2.3 Feedback (UV illumination, metal enrichment, remnants)
- 5.3 Later generations of stars
- 5.4 Global parameters of high-redshift galaxies
 - 5.4.1 Minimum mass of galaxies (filtering mass as a function of redshift, mini-halos with no cooling or star formation)
 - 5.4.2 Angular momentum
 - 5.4.3 Formation timescales
- 5.5 Gamma-ray Bursts: probing the first stars one star at a time
- 6. The First Black Holes
 - 6.1 Quasars: observational overview
 - 6.2 Potential seeds (massive stars, supermassive stars, direct collapse)
 - 6.3 Accretion theory (α -disks, radiatively inefficient accretion)
 - 6.4 Eddington limited growth $(L_E, \text{ phenomenology of quasars})$
 - 6.5 Mergers of black hole binaries (gravitational wave emission, recoil)
- 7. The epoch of reionization
 - 7.1 Growth of a single ionized region
 - 7.2 The global reionization history
 - 7.3 Statistical description of size distribution and topology of ionized regions
 - 7.4 Radiative transfer (numerical methods)
 - 7.5 Recombination of ionized regions
 - $7.6~\mathrm{The}$ sources of reionization
 - 7.6.1 Massive stars
 - 7.6.2 Quasars; limits on hard sources
 - 7.6.3 Exotic reionization scenarios
 - 7.7 Helium reionization
- 8. Feedback in the Early Universe
 - 8.1 Radiative feedback
 - 8.1.1 Heating of the intergalactic medium; minihalos and the clumping factor
 - 8.1.2 Photoheating and the suppression of low-mass galaxies
 - 8.1.3 Recombination radiation
 - 8.2 Large-scale mechanical feedback
 - 8.3 Chemical enrichment
- 9. The Ly α line as a probe of the early universe

- 9.1 Emission from galaxies
- 9.2 Scattering in the intergalactic medium (Ly α damping wing, Ly α halos)
- 9.3 The Gunn-Peterson trough and the ${\rm Ly}\alpha$ for est
- 10. The 21cm line
 - 10.1 Atomic physics (from first principles)
 - 10.2 Interaction with gas and UV/X-ray radiation backgrounds
 - 10.2 Statistical and imaging tools
 - 10.3 Observational prospects (with images)
 - 10.4 The transition to the post-reionization Universe

11. Galaxy surveys

- 11.1 Stellar populations
- 11.2 Galaxy evolution
- 11.3 Luminosity functions
- 11.4 Ly α and Lyman-break galaxies as probes of reionization
- 11.5 Molecules, dust, and the interstellar medium

12. Other observational probes of the first stars and galaxies

- 12.1 The cosmic microwave background
- 12.2 Low-redshift signatures

12.2.1 Clues from the intergalactic medium (thermal history, metal enrichment)

12.2.2 The fossil record of the local group